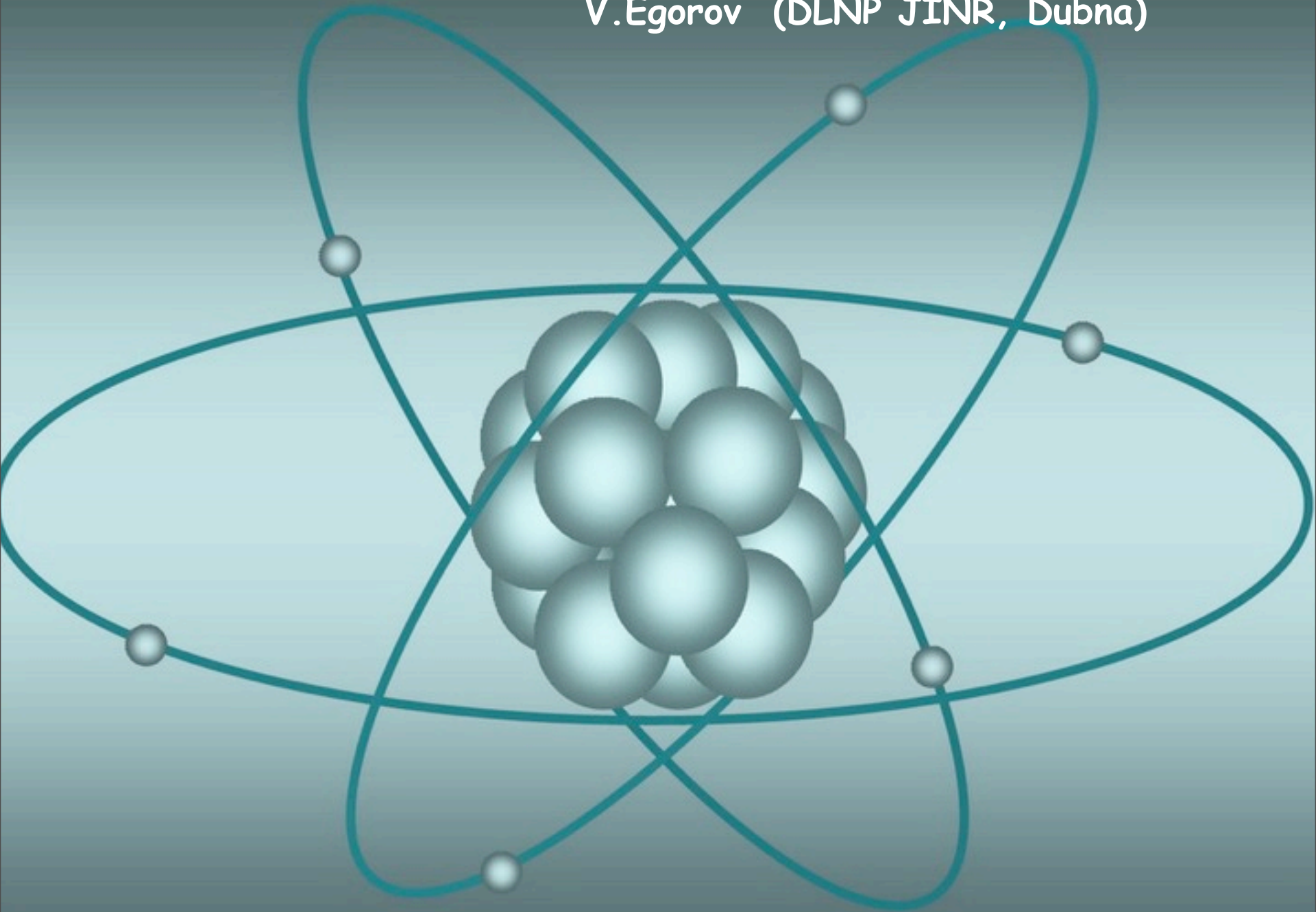
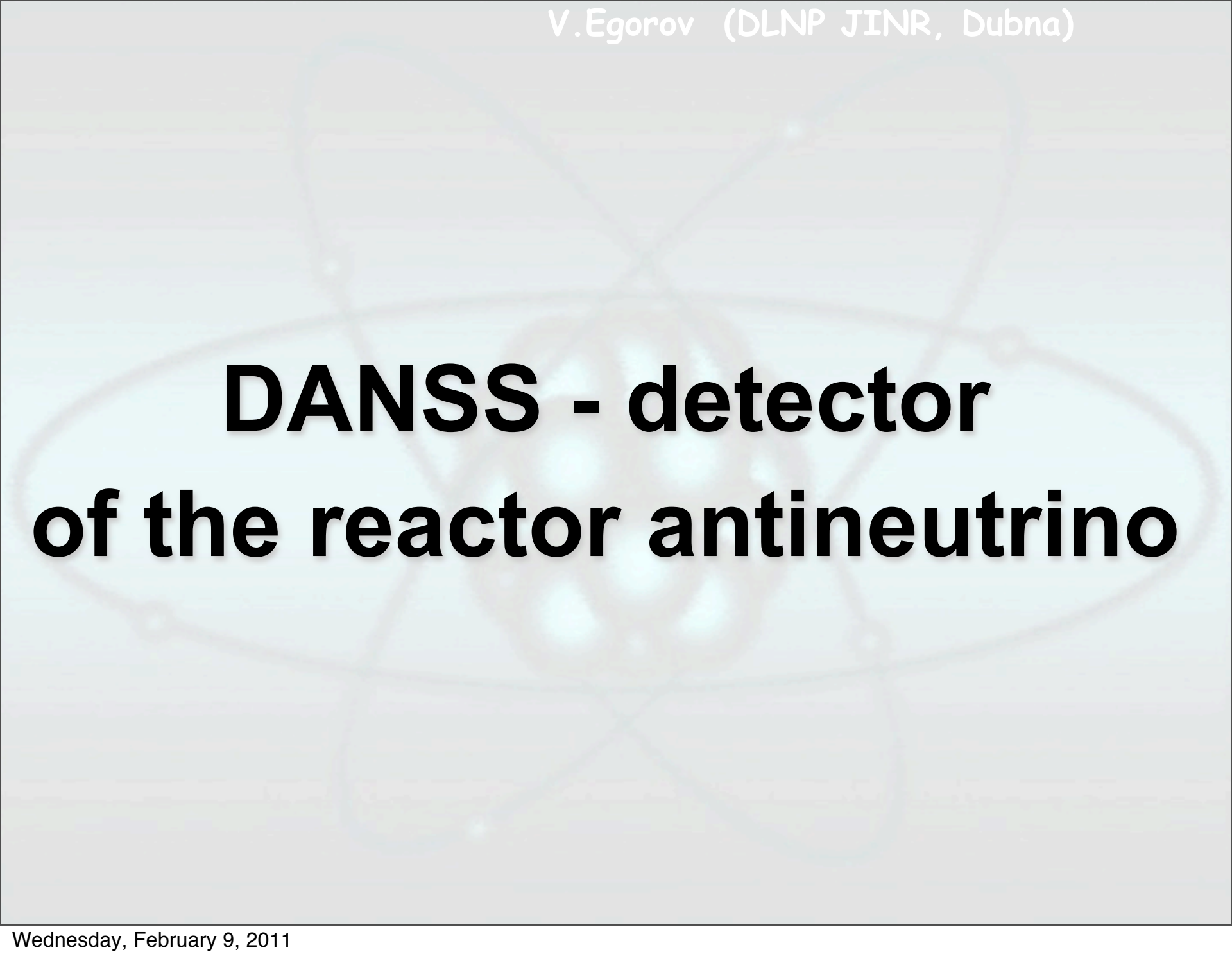


V.Egorov (DLNP JINR, Dubna)



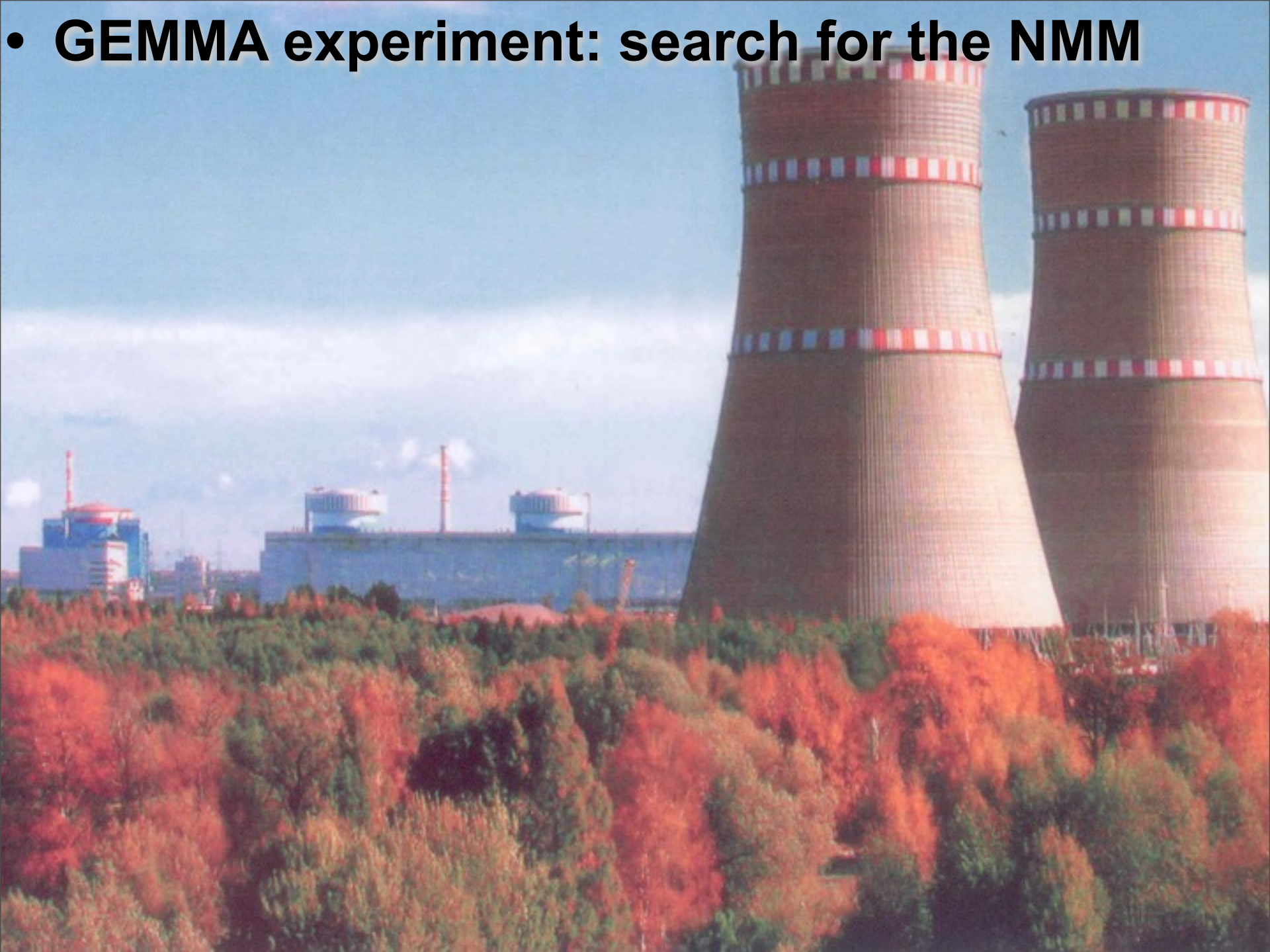


DANSS - detector of the reactor antineutrino

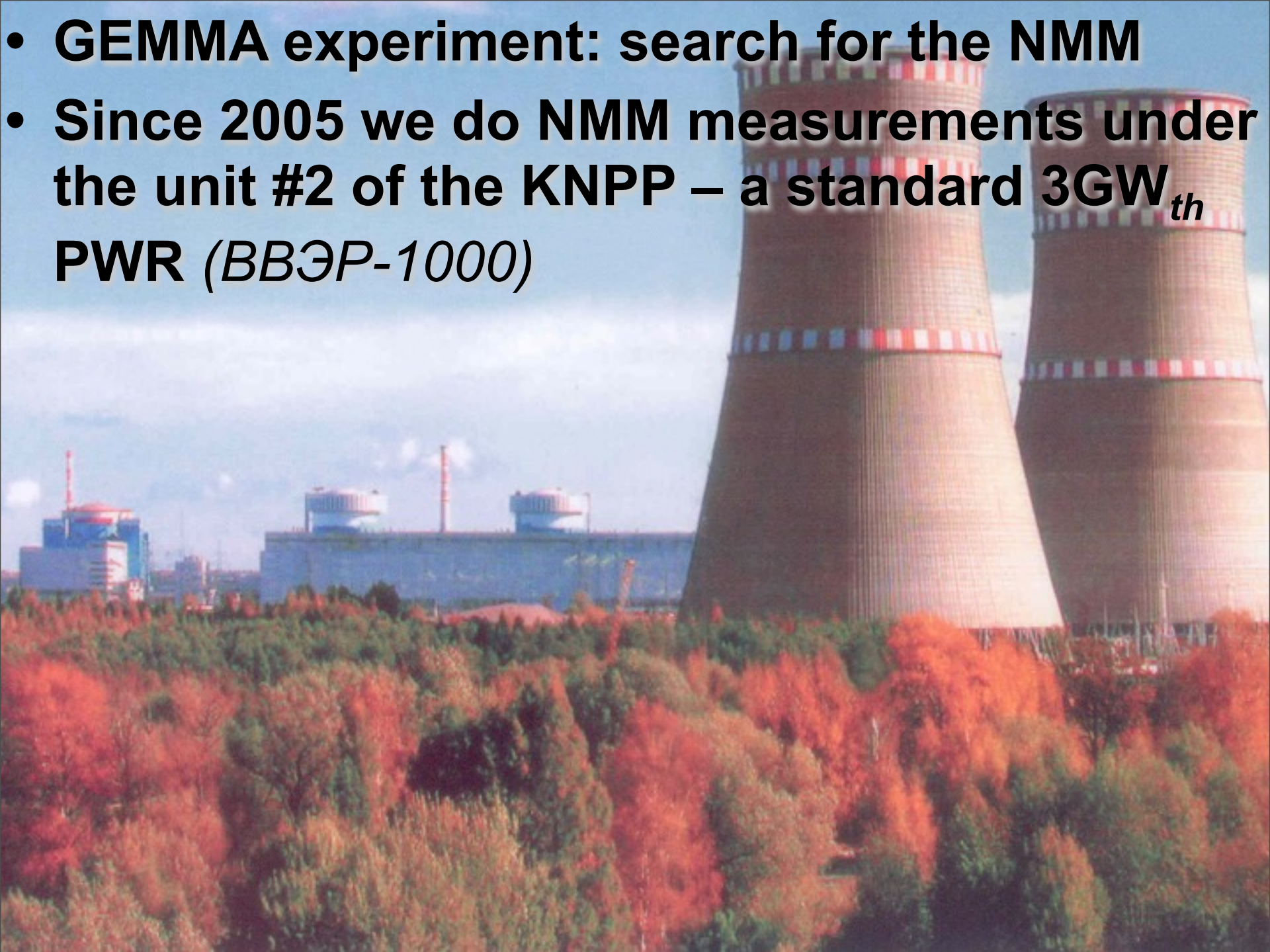


Wednesday, February 9, 2011

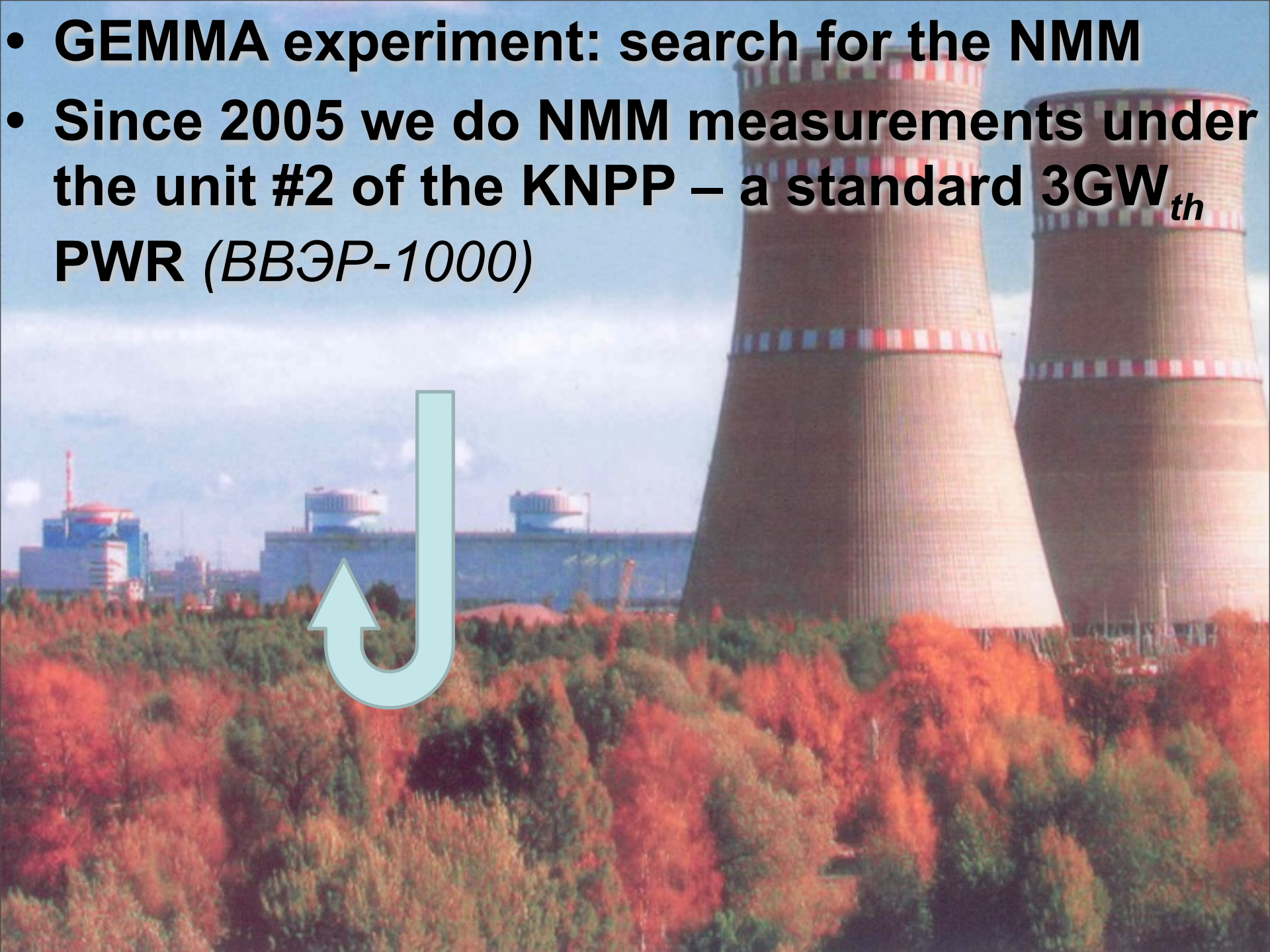
- **GEMMA experiment: search for the NMM**



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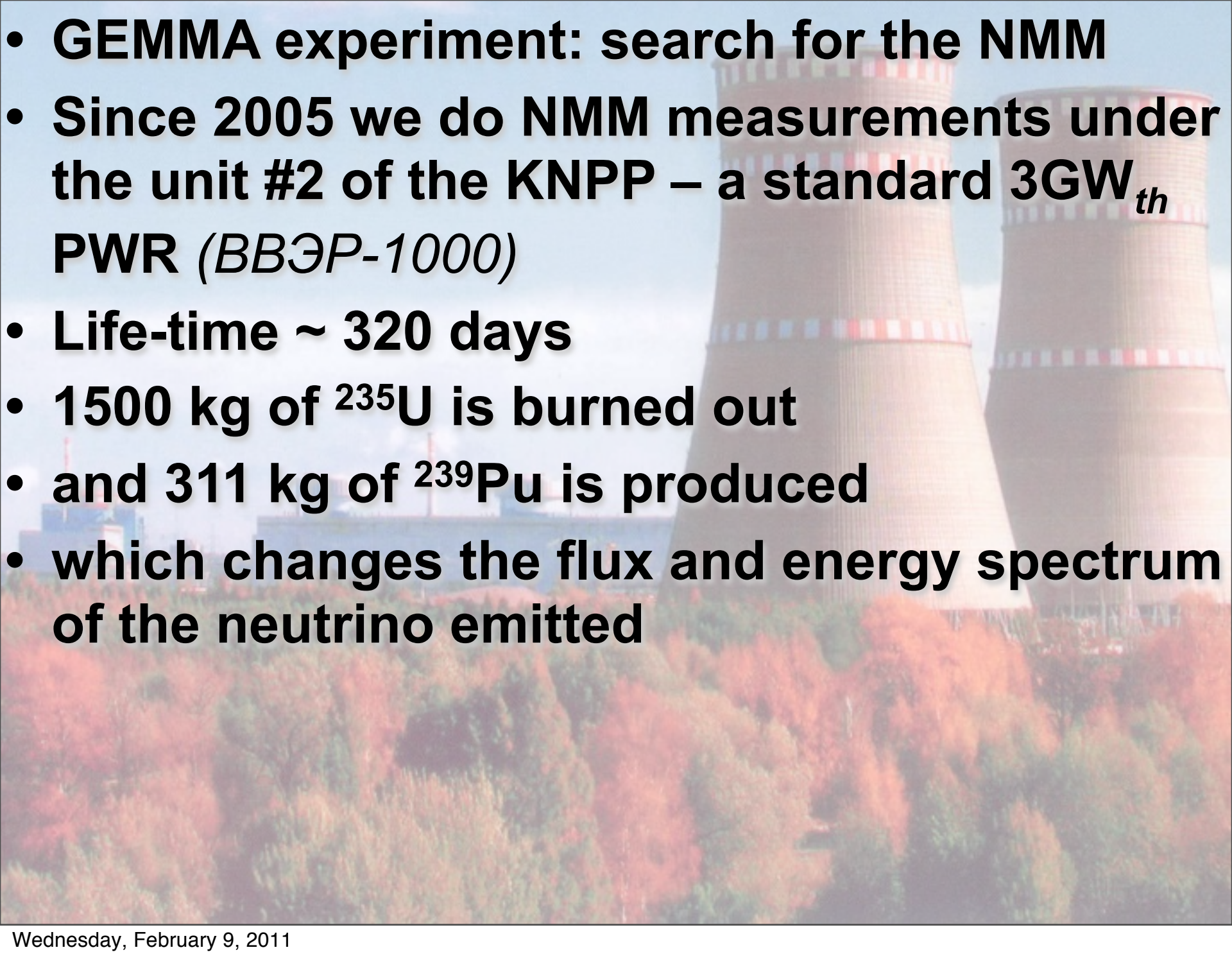
- **GEMMA experiment: search for the NMM**
- **Since 2005 we do NMM measurements under the unit #2 of the KNPP – a standard 3GW_{th} PWR (*BBЭP-1000*)**
- **Life-time ~ 320 days**

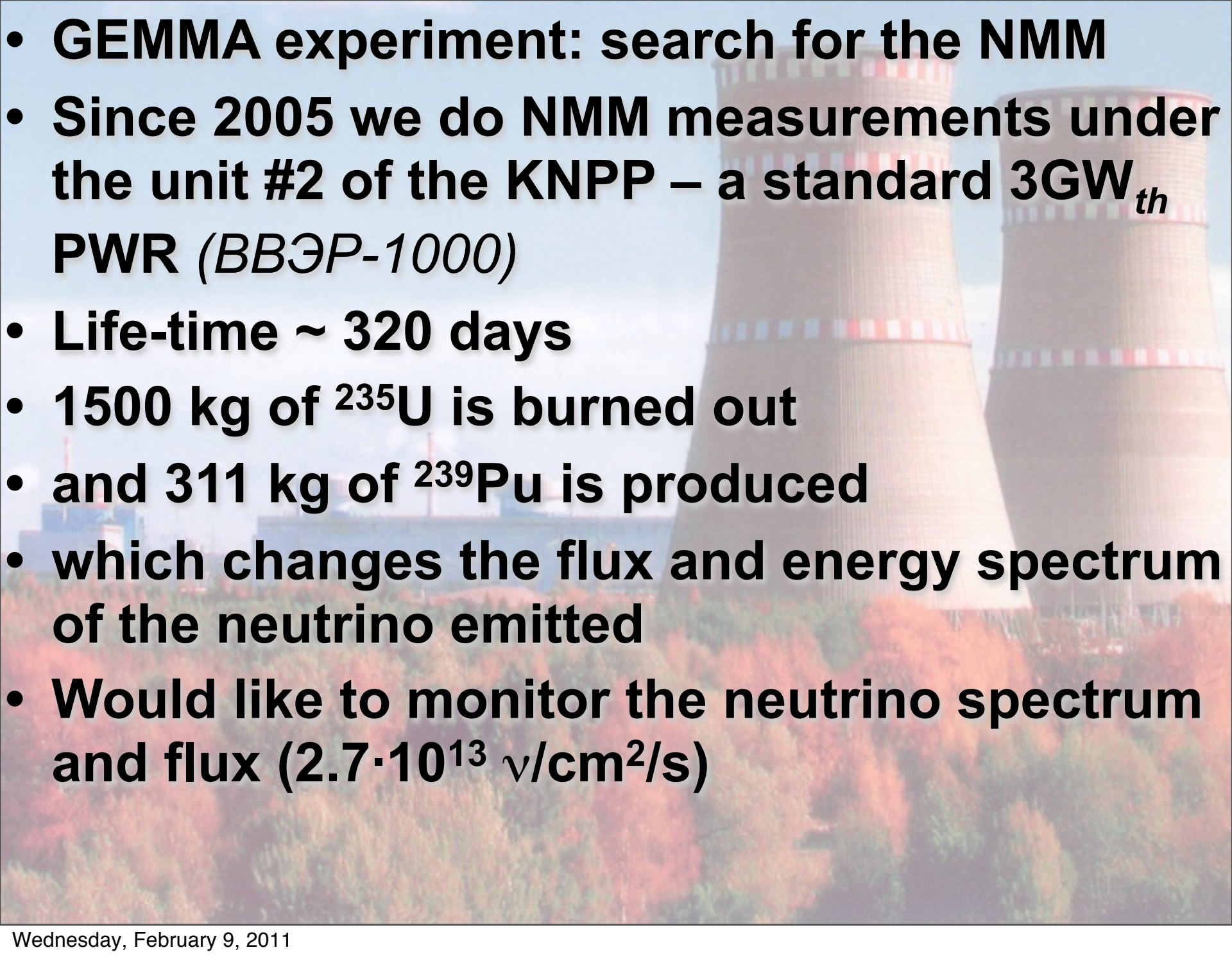


- **GEMMA experiment: search for the NMM**
- **Since 2005 we do NMM measurements under the unit #2 of the KNPP – a standard 3GW_{th} PWR (*BBЭP-1000*)**
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 - **1500 kg of ^{235}U is burned out**
 - **and 311 kg of ^{239}Pu is produced**
 - **which changes the flux and energy spectrum of the neutrino emitted**
 - **Would like to monitor the neutrino spectrum and flux ($2.7 \cdot 10^{13} \nu/\text{cm}^2/\text{s}$)**

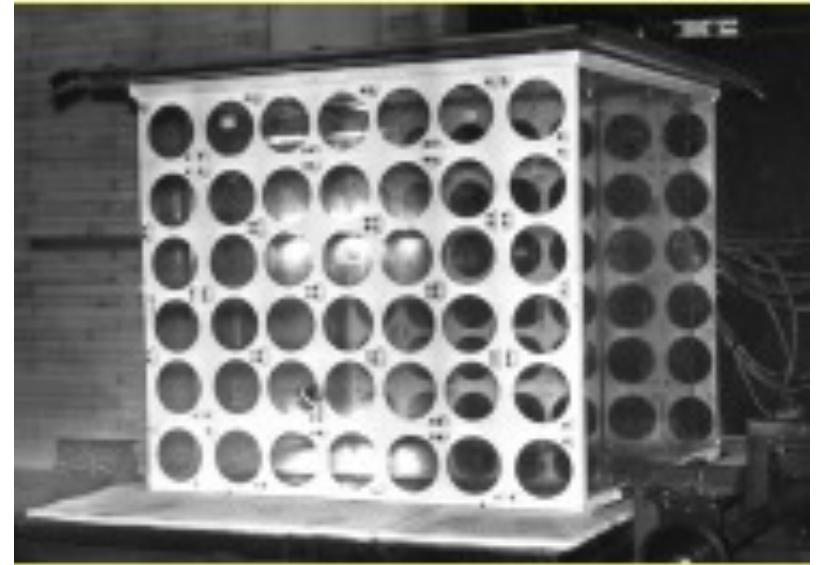
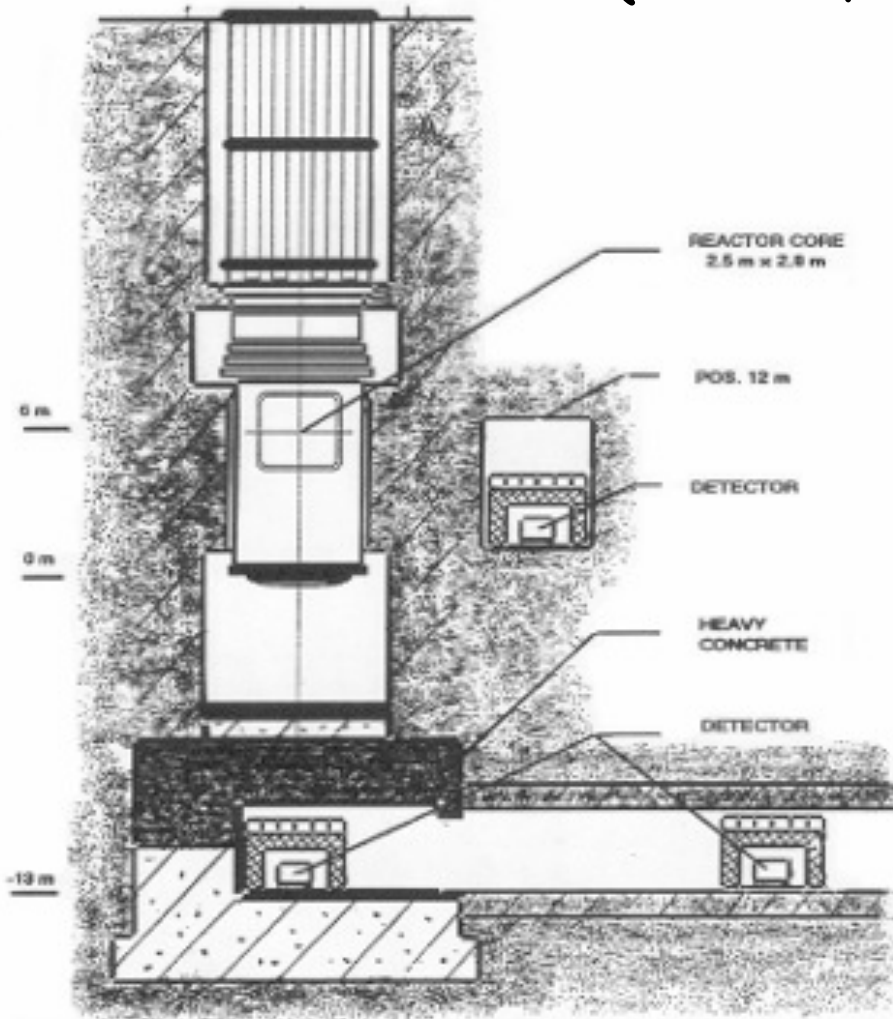


**For the monitoring we needed
anything simple and reliable**

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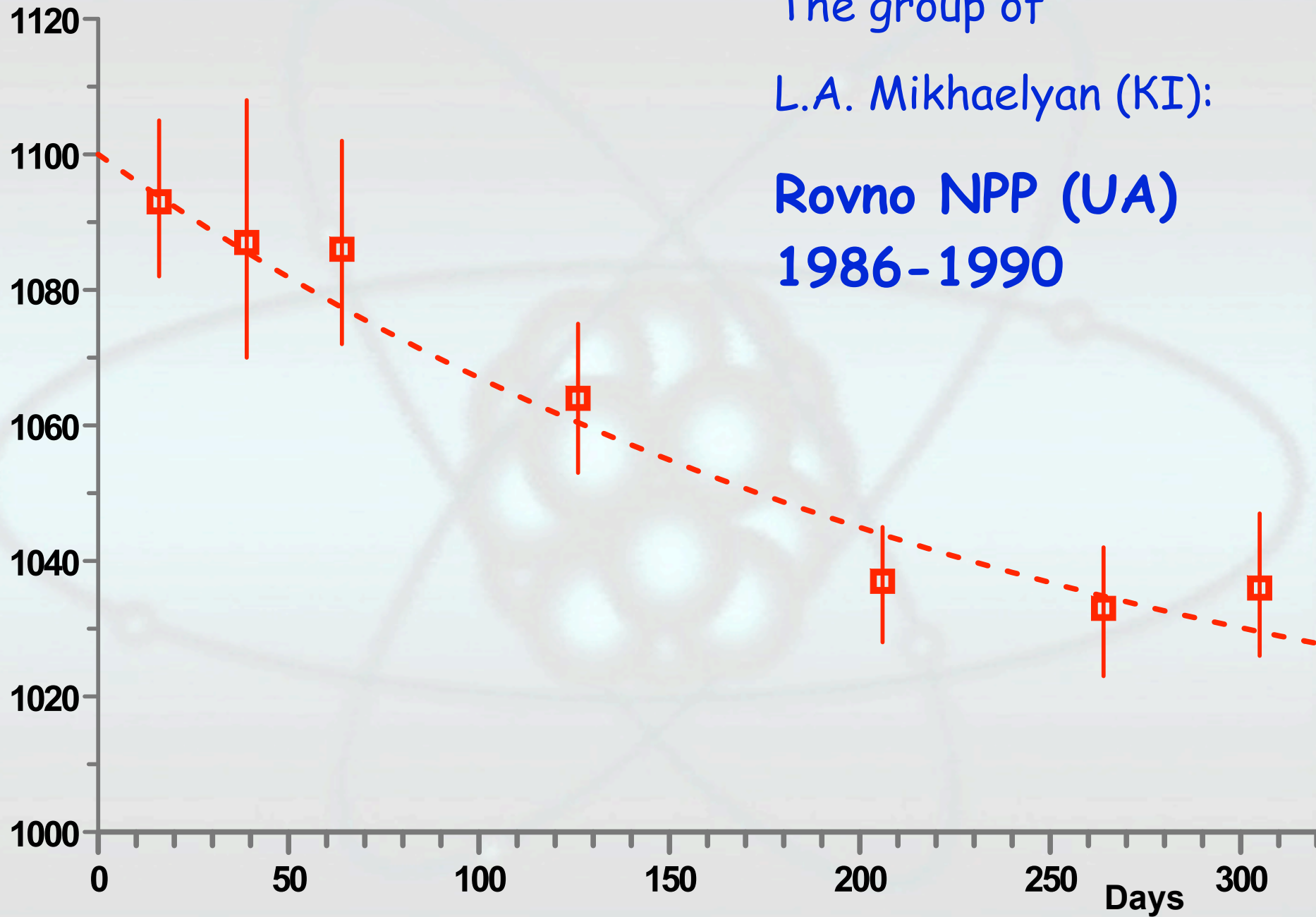


20 years ago our colleagues from KI were the first
who tried to perform the neutrino monitoring
RONS (Rovno, UA) 1986-1990

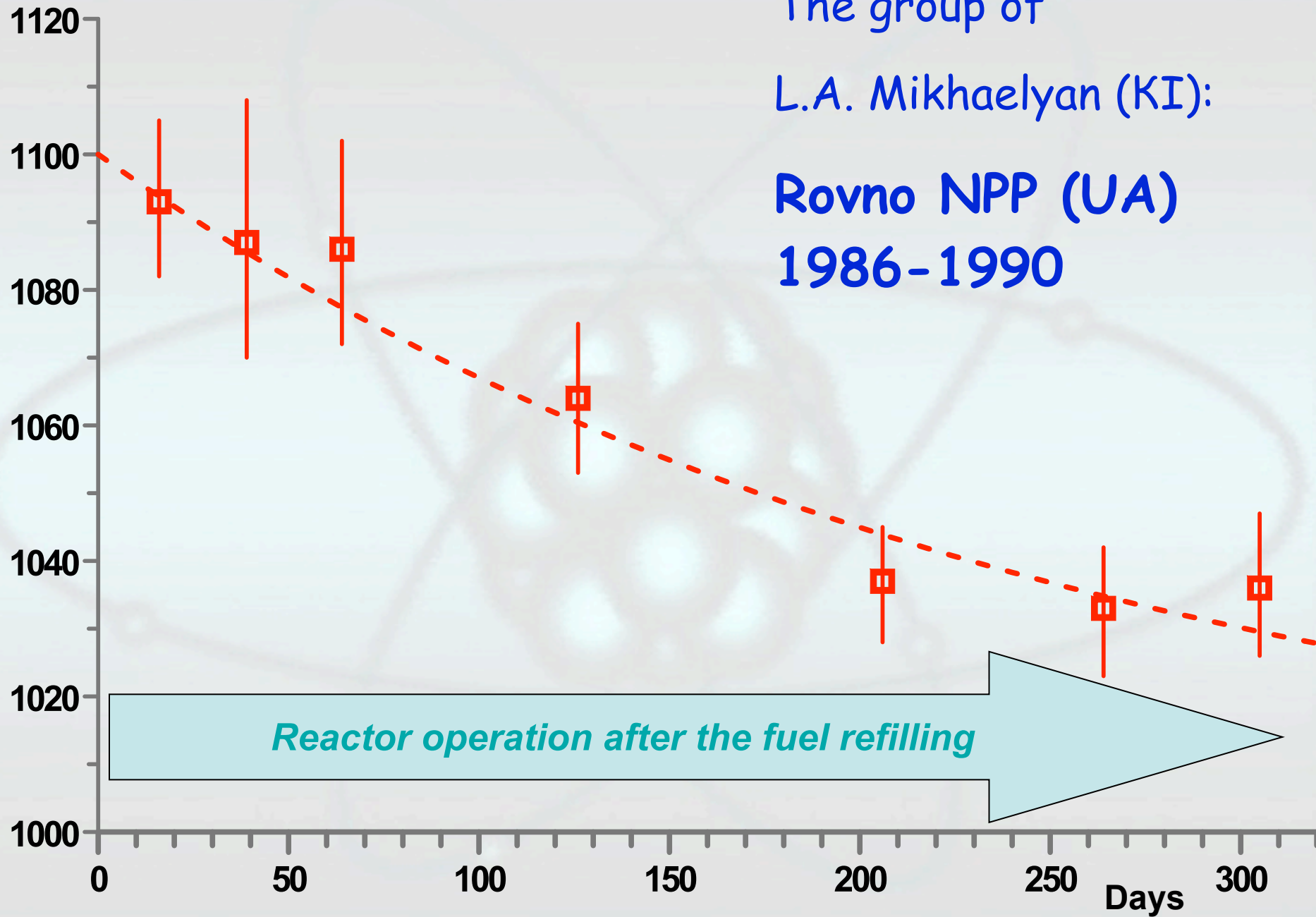


Liquid scintillator ($\sim 1 \text{ m}^3$)
in a **special laboratory**

Counts per 100 000 sec

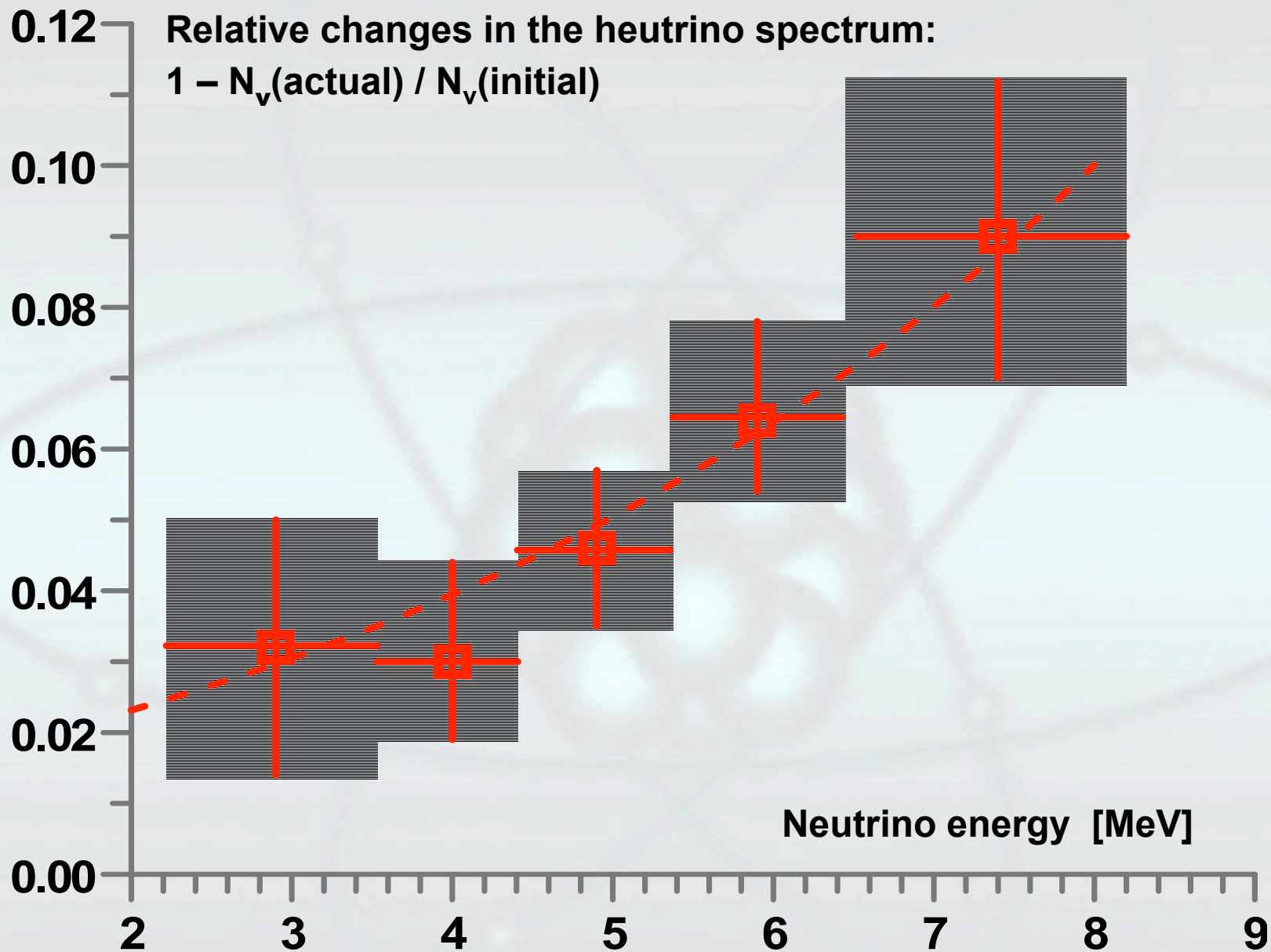


Counts per 100 000 sec



The group of
L.A. Mikhaelyan (KI):
Rovno NPP (UA)
1986-1990

Reactor operation after the fuel refilling





**We want
such**



**We want
such**

«We» - means JINR (Dubna)



**We want
such**

«We» - means JINR (Dubna)

+



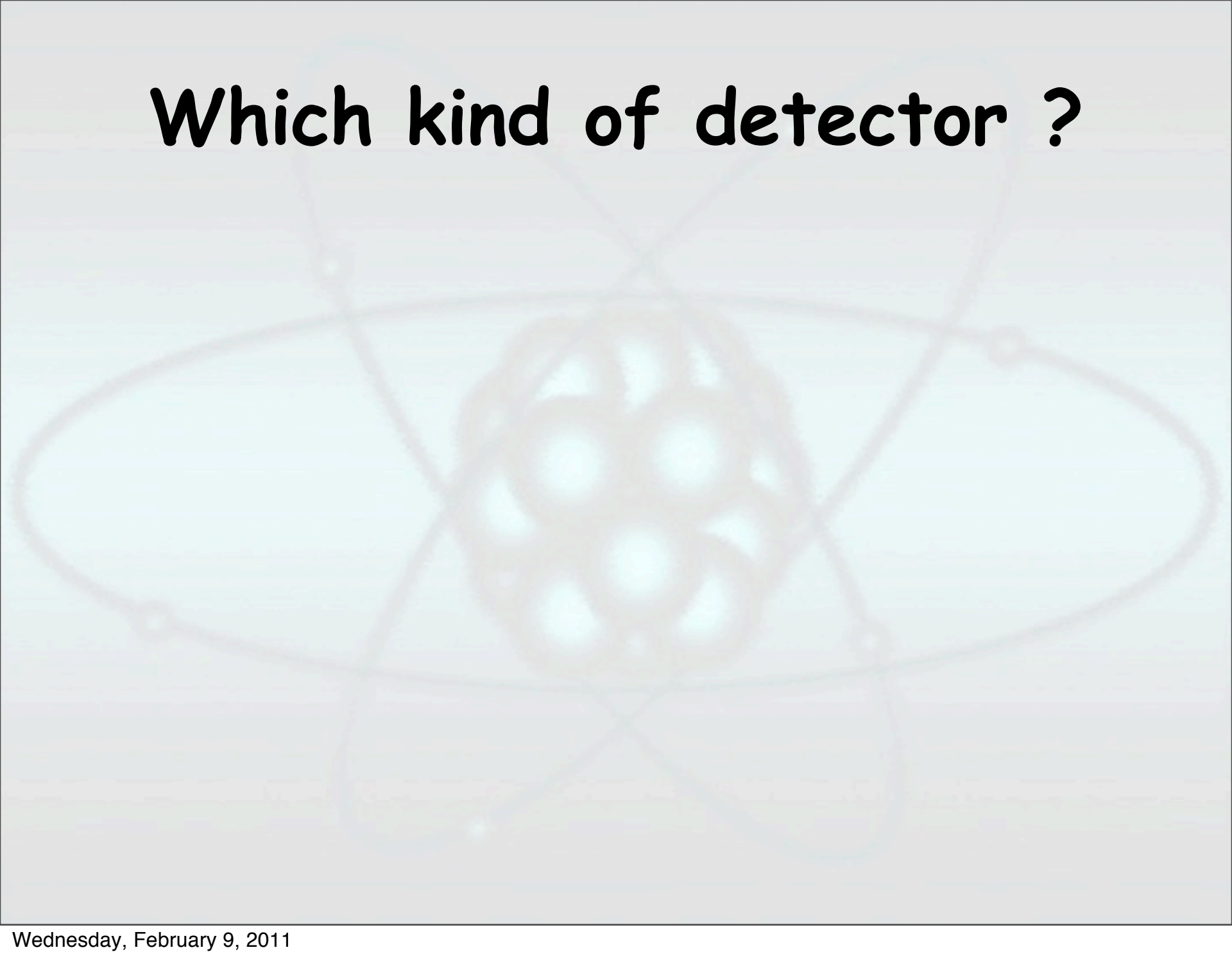
**We want
such**

«We» - means JINR (Dubna)

+

ITEP (Moscow)

Which kind of detector ?



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- Generally $\sim 1 \text{ m}^3$ of liquid scintillator is used

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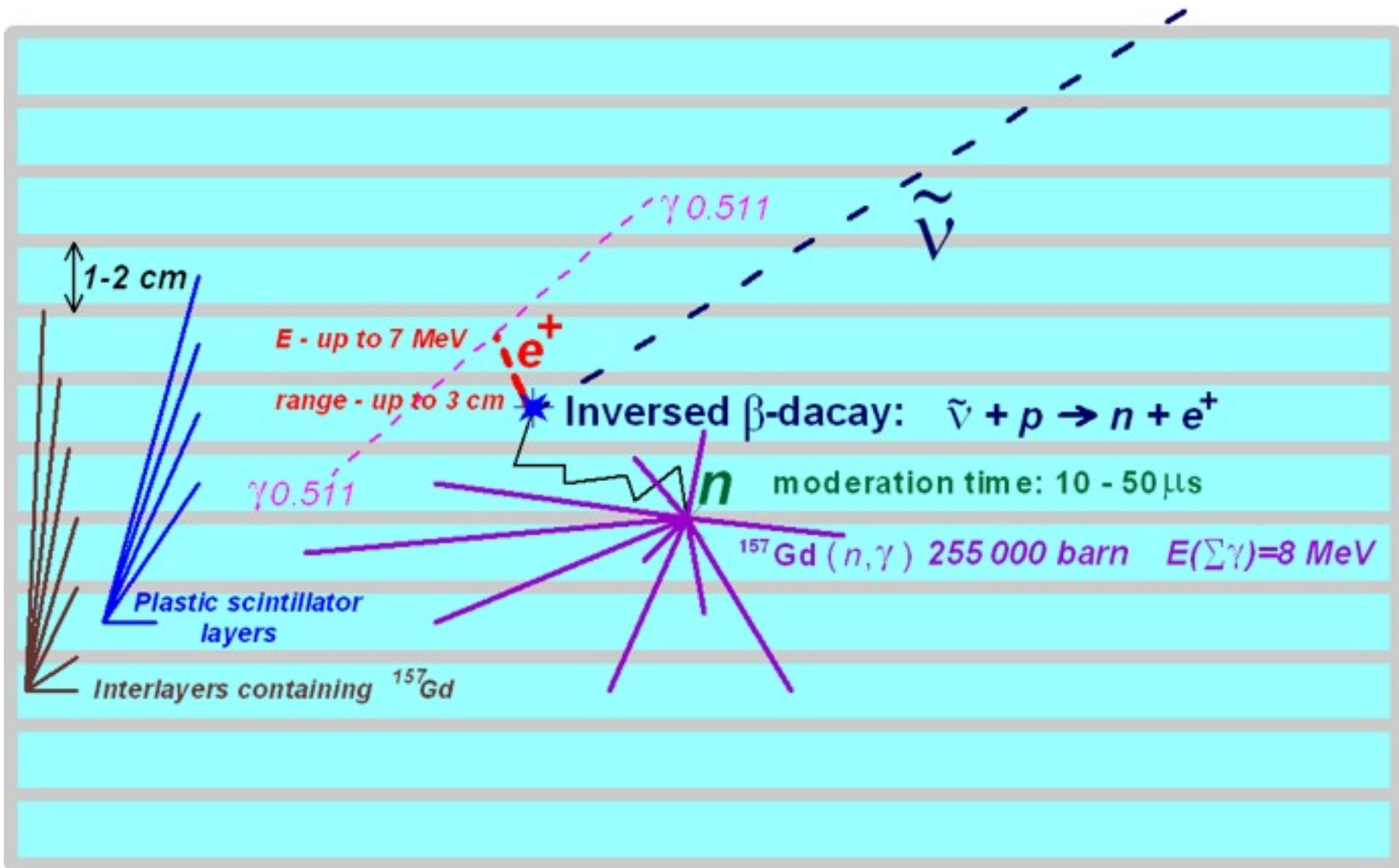
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- ☹️ **But:** is there any **SANE DIRECTOR** of
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close to his reactor ?!

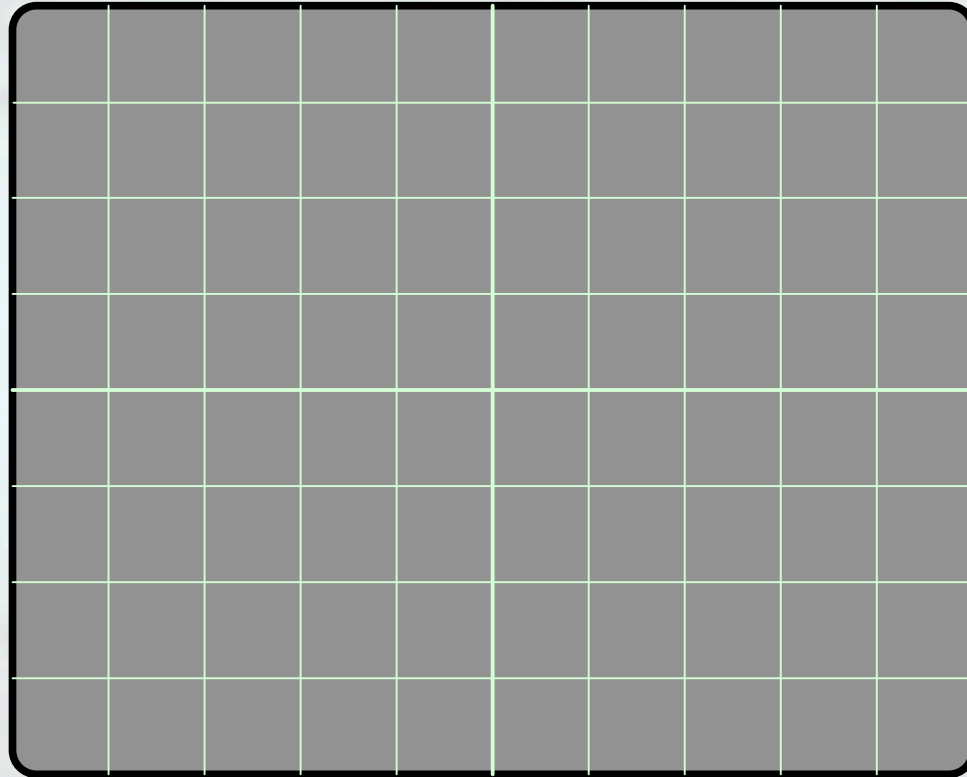
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RONS (SU), NUCIFER (FR), ANGRA (BR),
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- ☹️ **But:** they say it is problematic (Gd...)
- ☹️ **But:** is there any **SANE DIRECTOR** of NPP who allows to bring 1 ton of gasoline close to his reactor ?!
- 😊 **How much better is plastic**

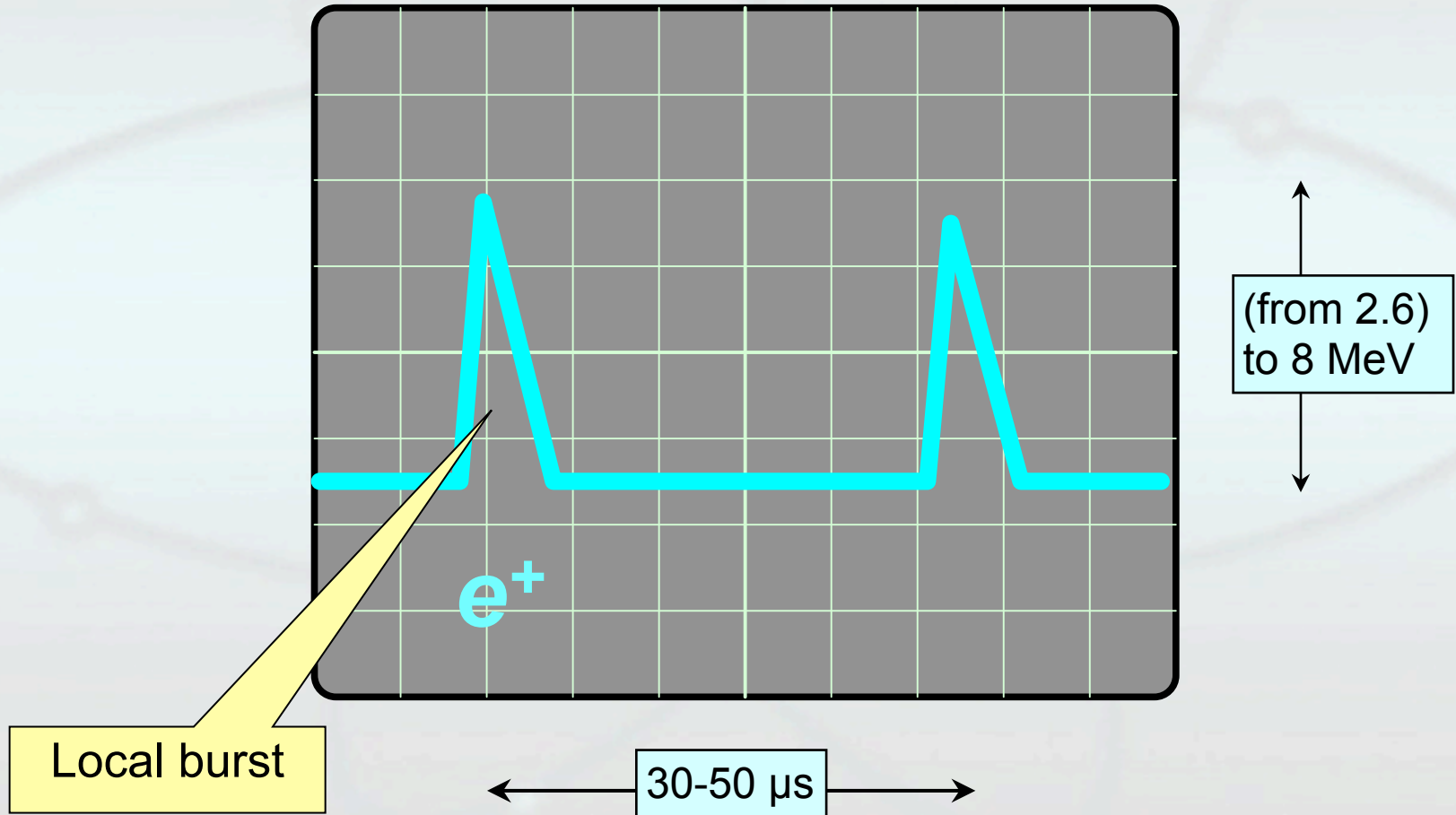
Detection idea: Inversed Beta-Decay



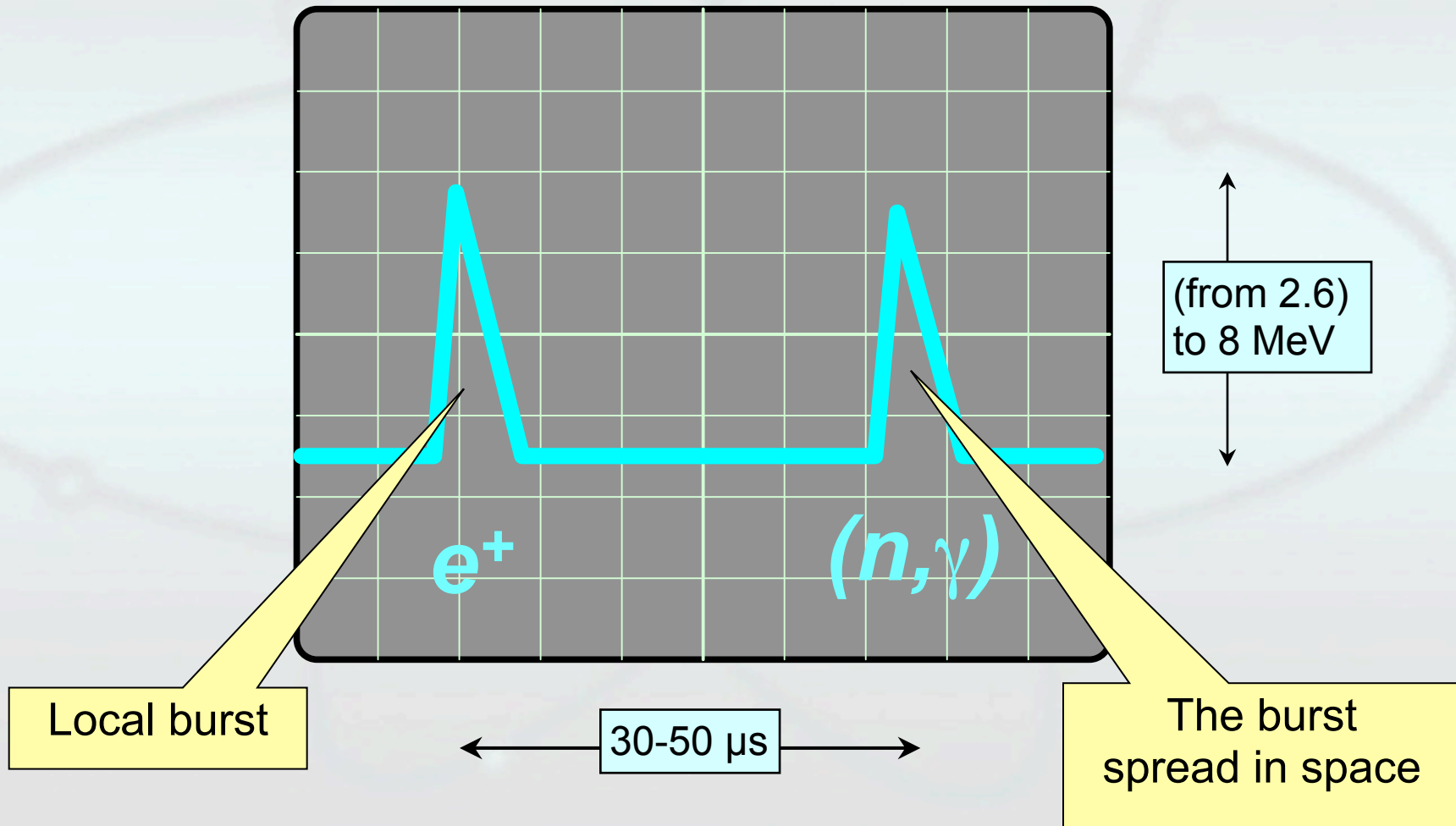
Signature of the IBD registration



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Detection of the reactor neutrino would allow:



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- Measure the actual reactor power (N_ν)
- Deduce the actual fuel composition (E_ν)
- Weak (ν - e) cross-section
- On-line reactor monitoring (tomography?) - especially important in view of the future FBR (with

geometry of neutrino detector

SCINTILLATOR BLOCKS

in X axis

bottom-scinblock.number-of-columns=100

in Y axis

bottom-scinblock.number-of-rows=25

in X axis

bottom-scinblock.width=1 cm

in Y axis

bottom-scinblock.length=4 cm

in Z axis

bottom-scinblock.thickness=1 m

material of capturer layer on the scintillator surface

available materials: Gadolinium Cadmium

bottom-scinblock.n-capt.material=Cadmium

thickness of capturer layer on the X surface

bottom-scinblock.n-capt.x-thickness = 150 micrometer

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NEUDET CHAMBER

global X size

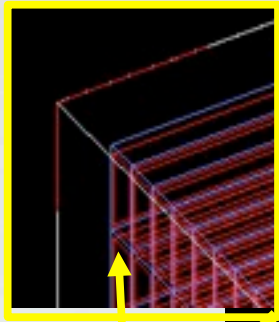
chamber.width = 1.2 m

global Y size

chamber.length = 1.2 m

global Z size

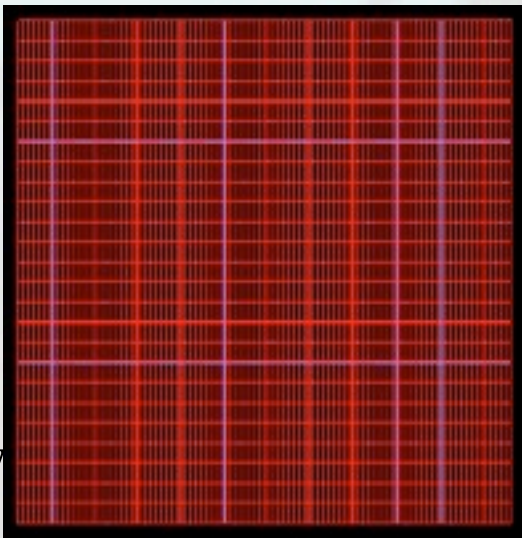
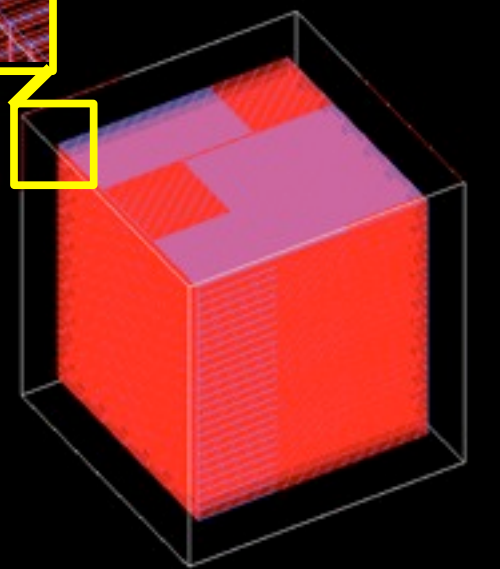
chamber.thickness = 1.2 m



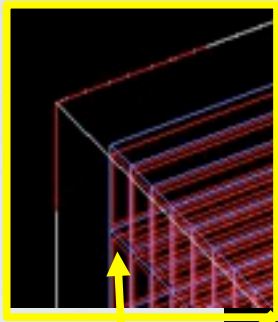
XY-view



Layer of
N-capturer
(Gd,Cd,...)



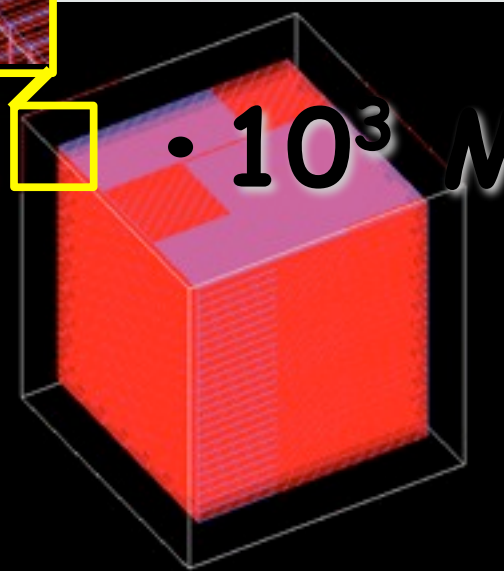
XY-view



XY-view

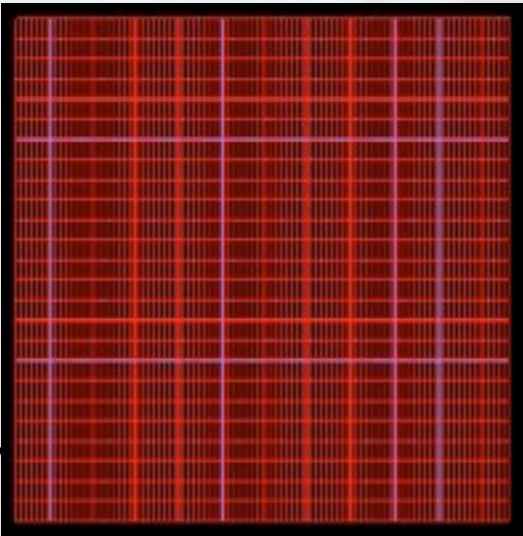


Layer of
N-capturer
(Gd,Cd,...)



$\cdot 10^3$

MC-simulations



XY-view

geometry of neutrino detector

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NEUDET CHAMBER

global X size

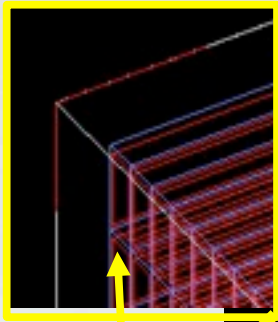
chamber.width = 1.2 m

global Y size

chamber.length = 1.2 m

global Z size

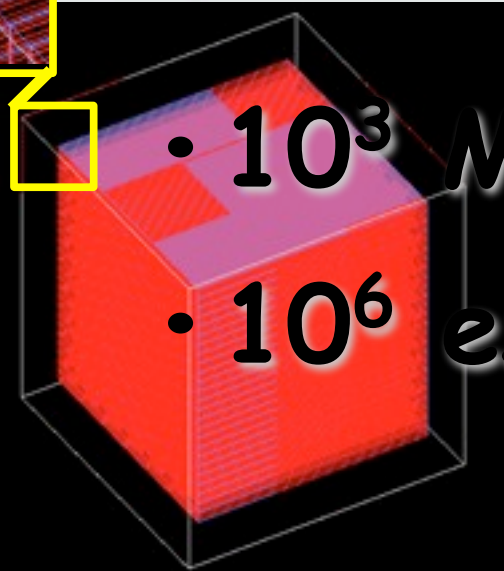
chamber.thickness = 1.2 m



XY-view



Layer of
N-capturer
(Gd,Cd,...)



• 10^3

• 10^6

MC-simulations

experimental tests

geometry of neutrino detector

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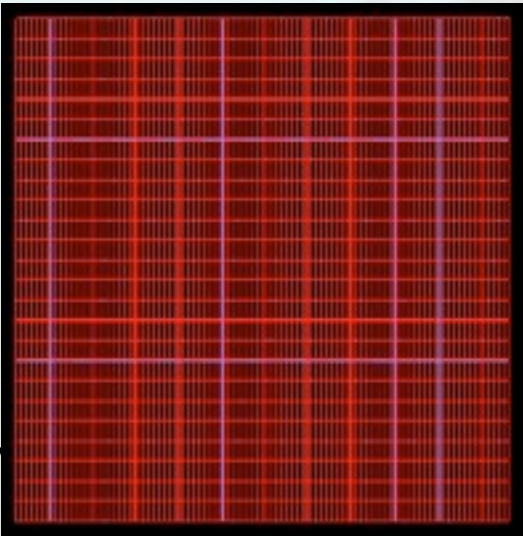
global Y size

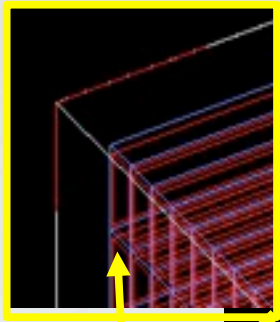
chamber.length = 1.2 m

global Z size

chamber.thickness = 1.2 m

XY-view

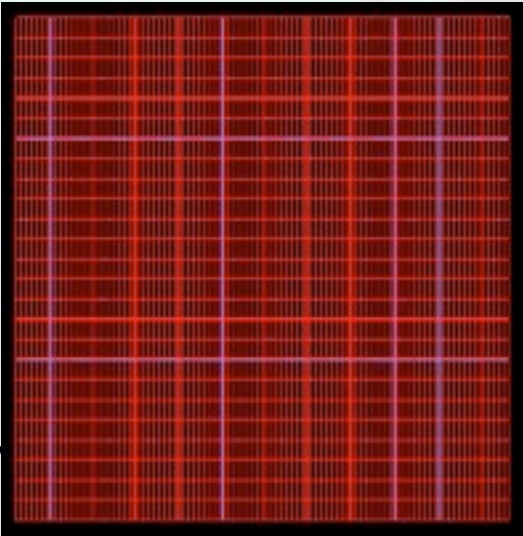
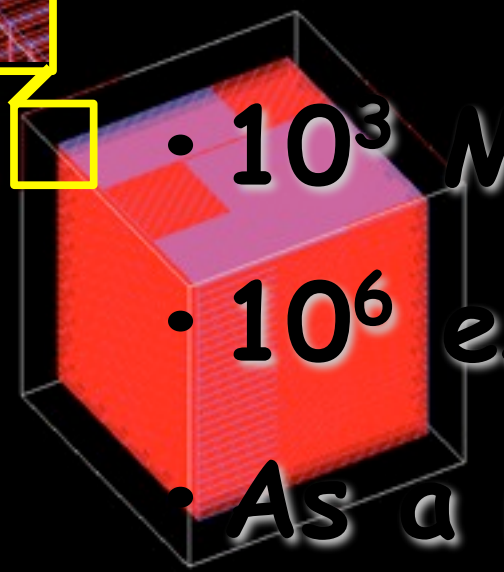




XY-view



Layer of
N-capturer
(Gd,Cd,...)



XY-view

geometry of neutrino detector

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NEUDET CHAMBER

global X size

chamber.width = 1.2 m

global Y size

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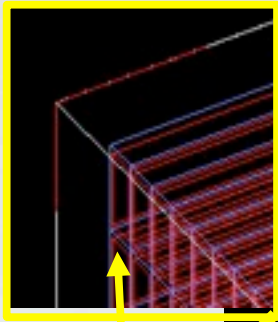
global Z size

chamber.thickness = 1.2 m

MC-simulations

experimental tests

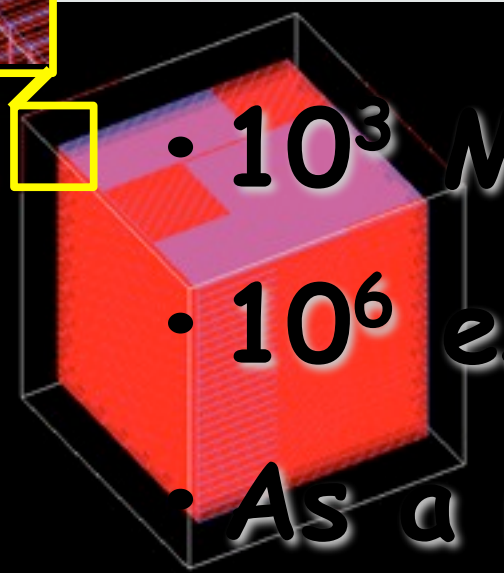
As a result:



XY-view



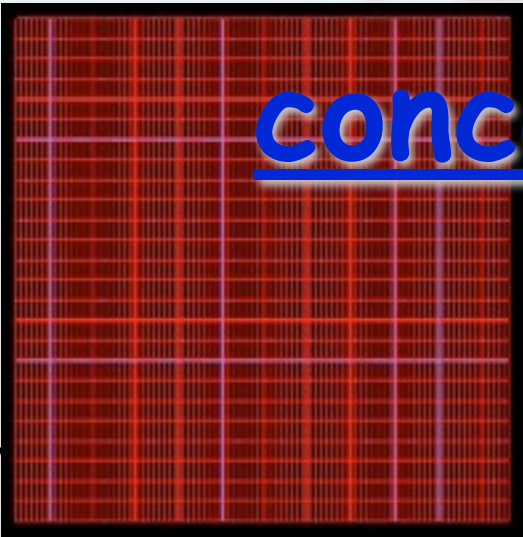
Layer of
N-capturer
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• 10^3

• 10^6

• As a result:



XY-view

geometry of neutrino detector

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NEUDET CHAMBER

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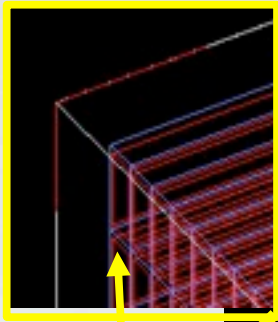
global Y size

chamber.length = 1.2 m

global Z size

chamber.thickness = 1.2 m

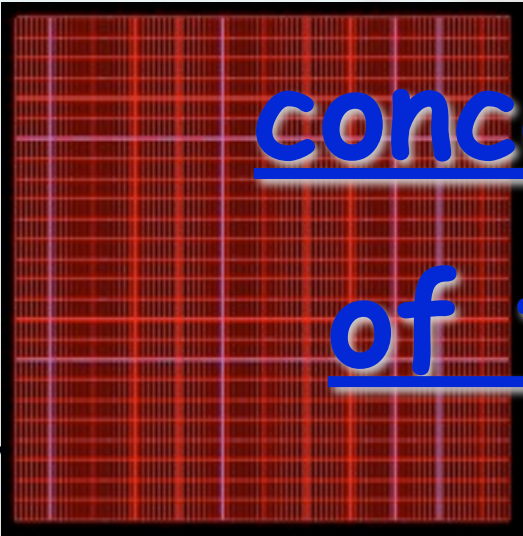
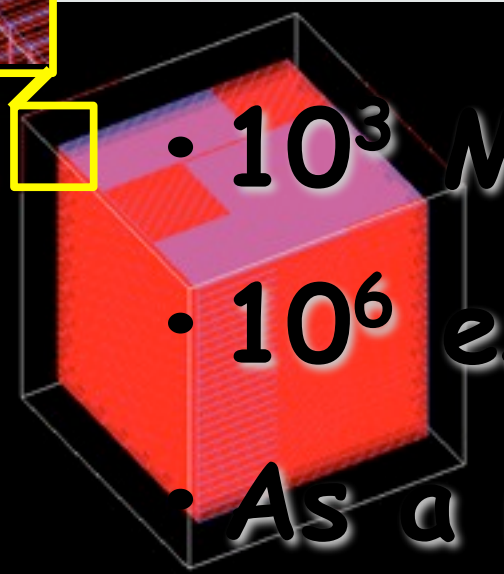
conception



XY-view



Layer of N-capturer (Gd,Cd,...)



XY-view

geometry of neutrino detector

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NEUDET CHAMBER

global X size

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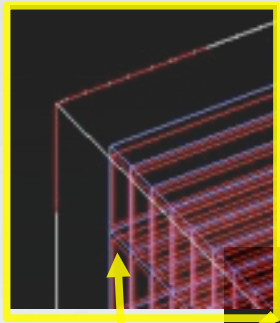
chamber.length = 1.2 m

global Z size

chamber.thickness = 1.2 m

conception

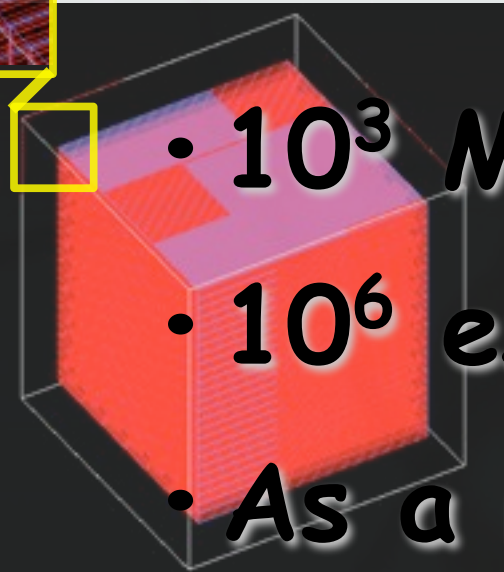
of the spectrometer



XY-view



Layer of N-capturer (Gd,Cd,...)



- 10^3
- 10^6
- As a result:

MC-simulations

experimental tests

As a result:

conception

of the spectrometer

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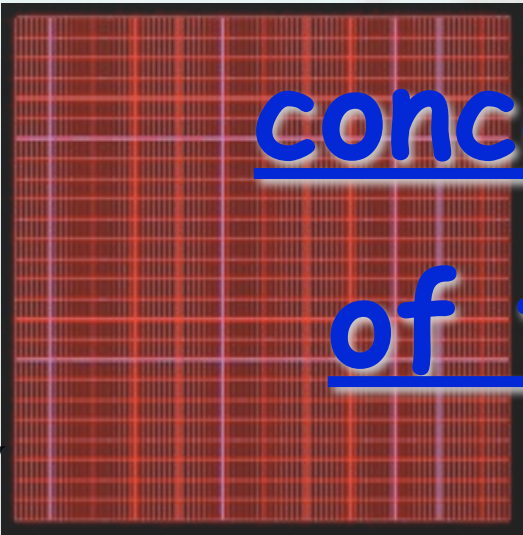
global Y size

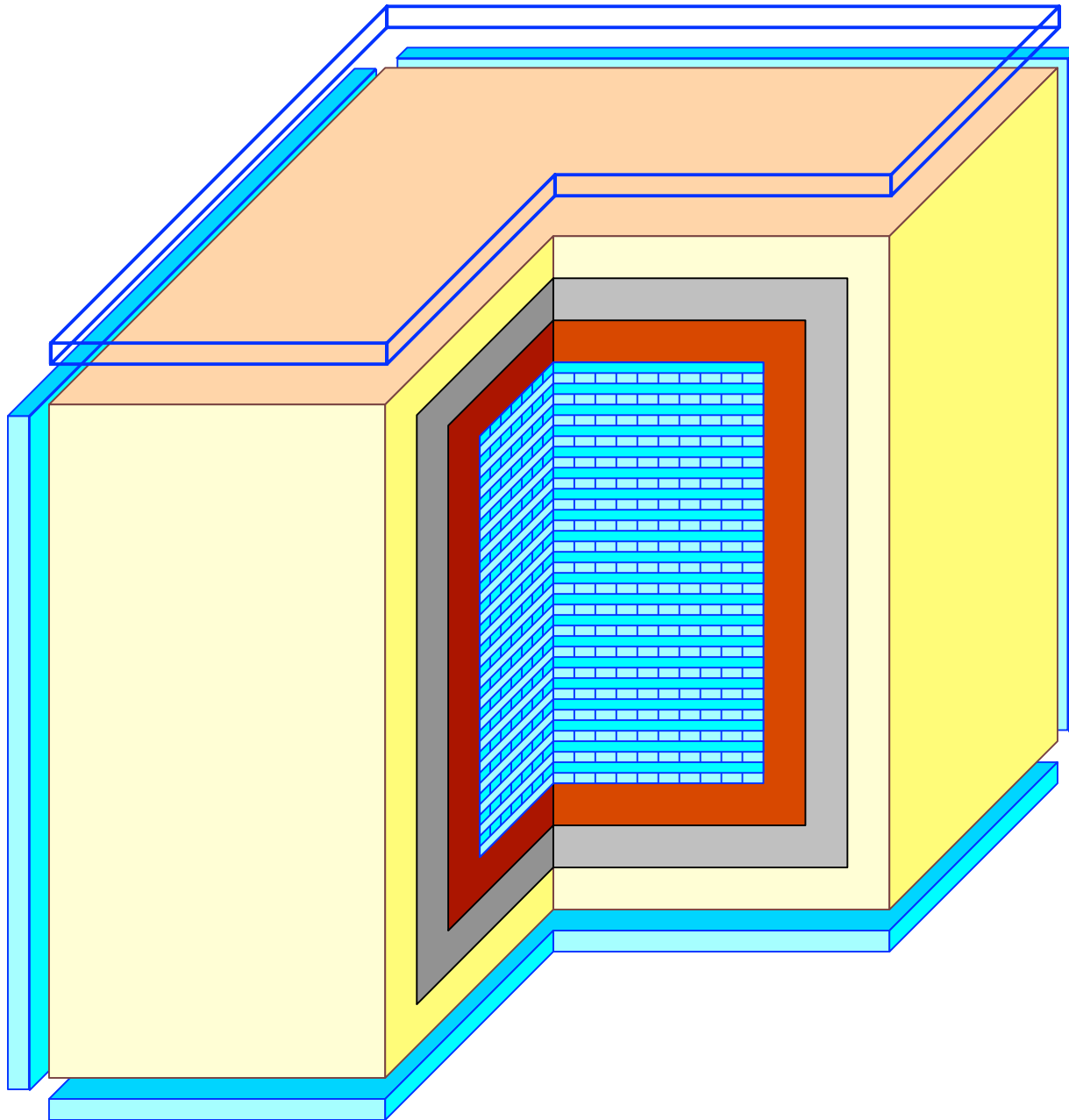
chamber.length = 1.2 m

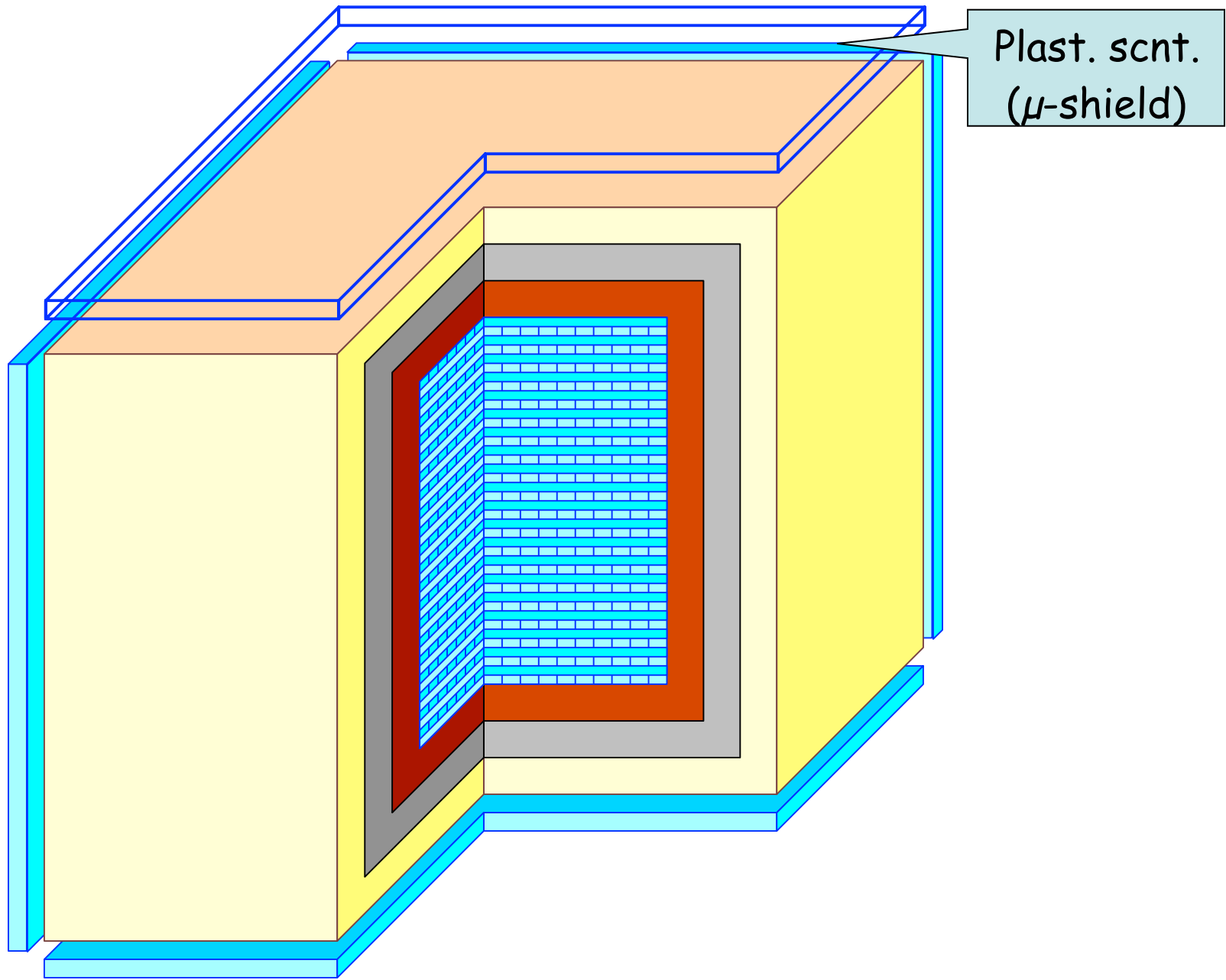
global Z size

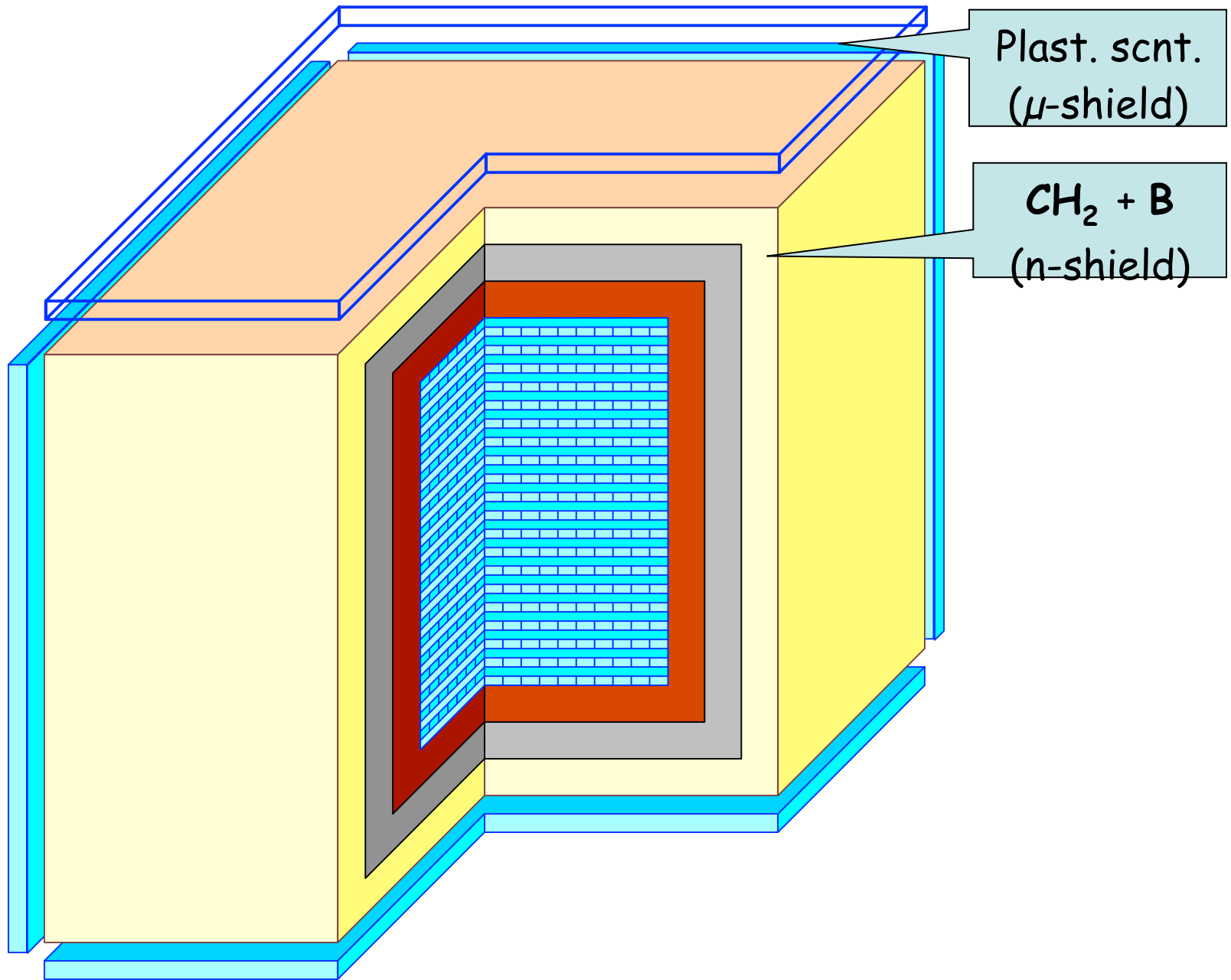
chamber.thickness = 1.2 m

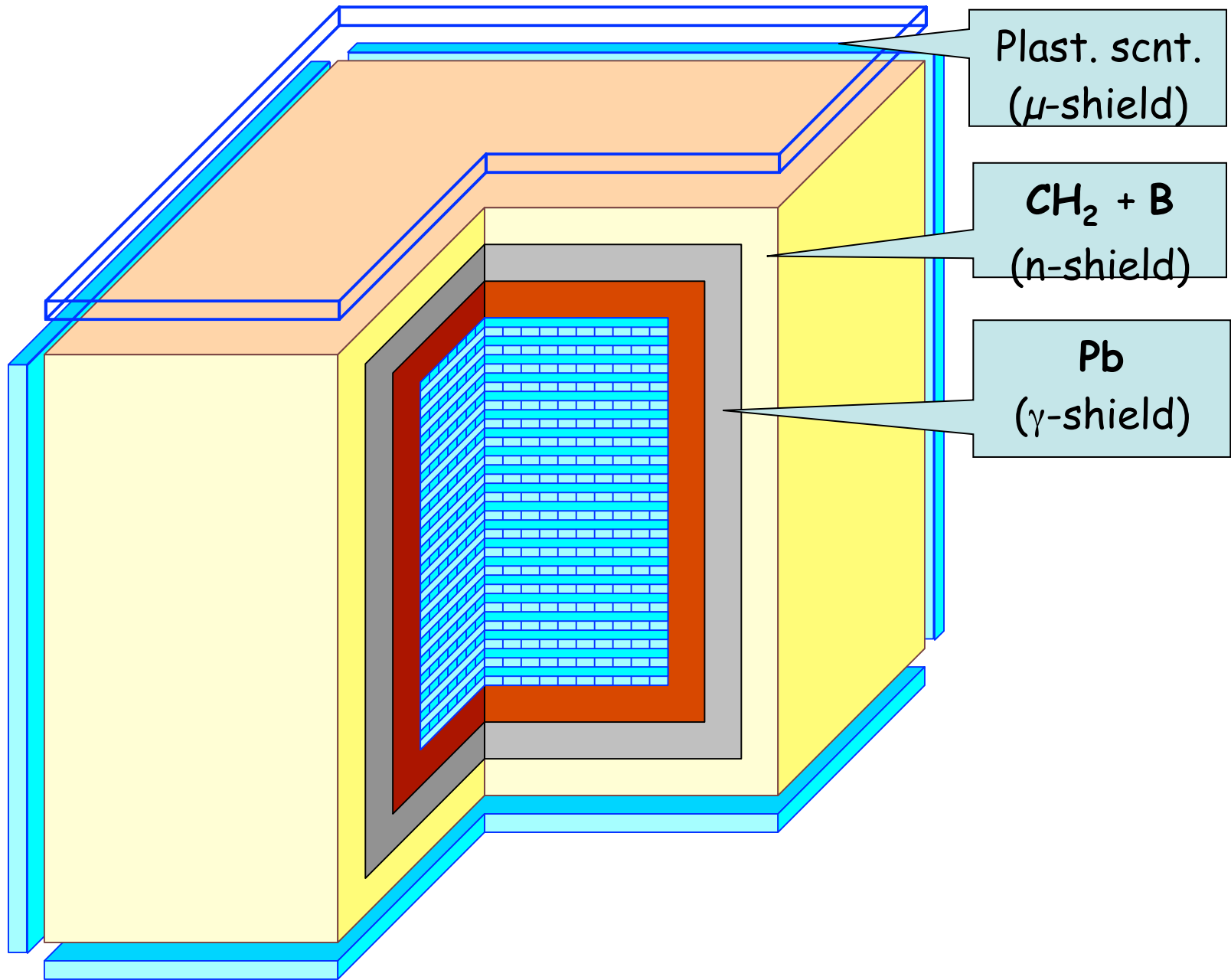
XY-view

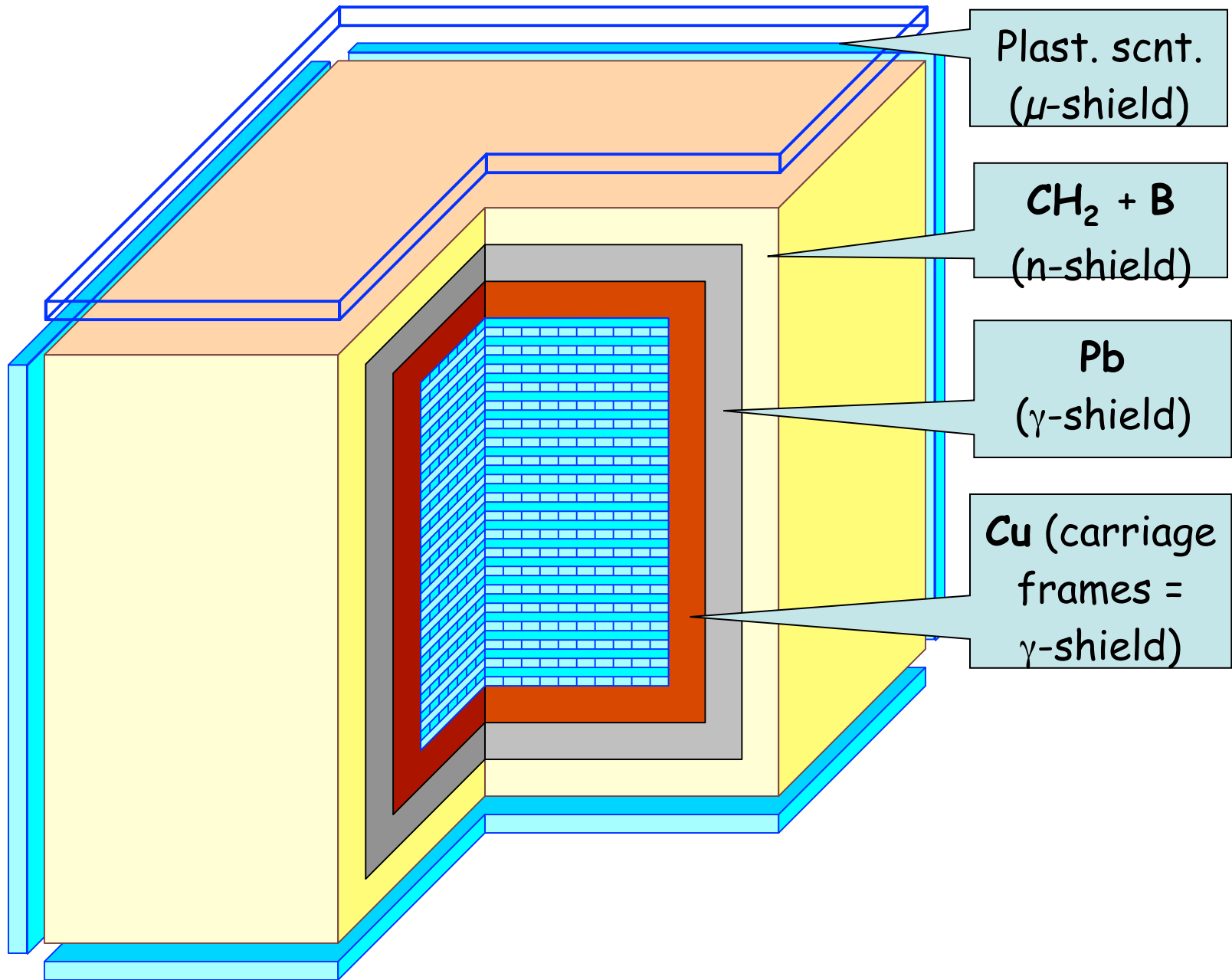


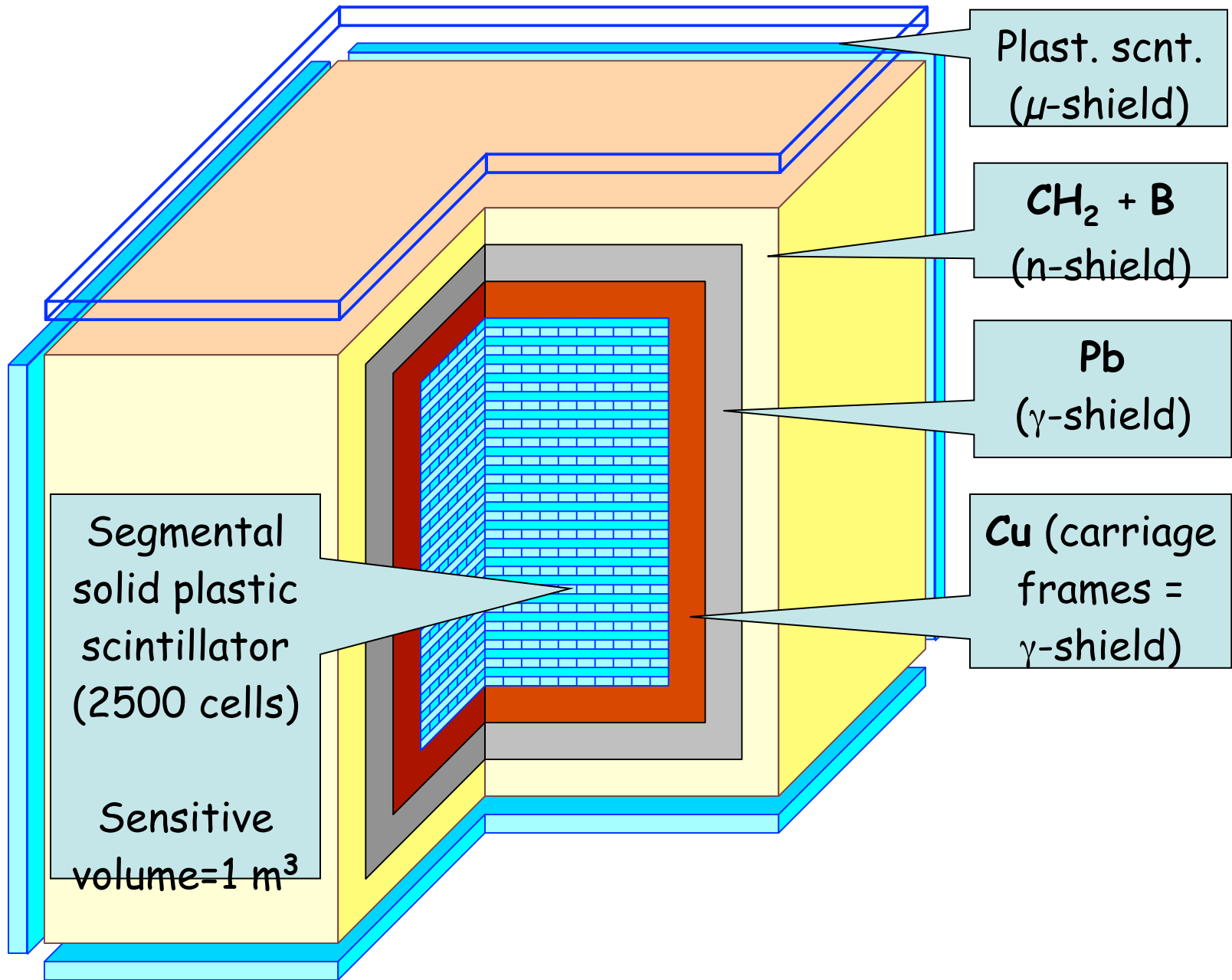


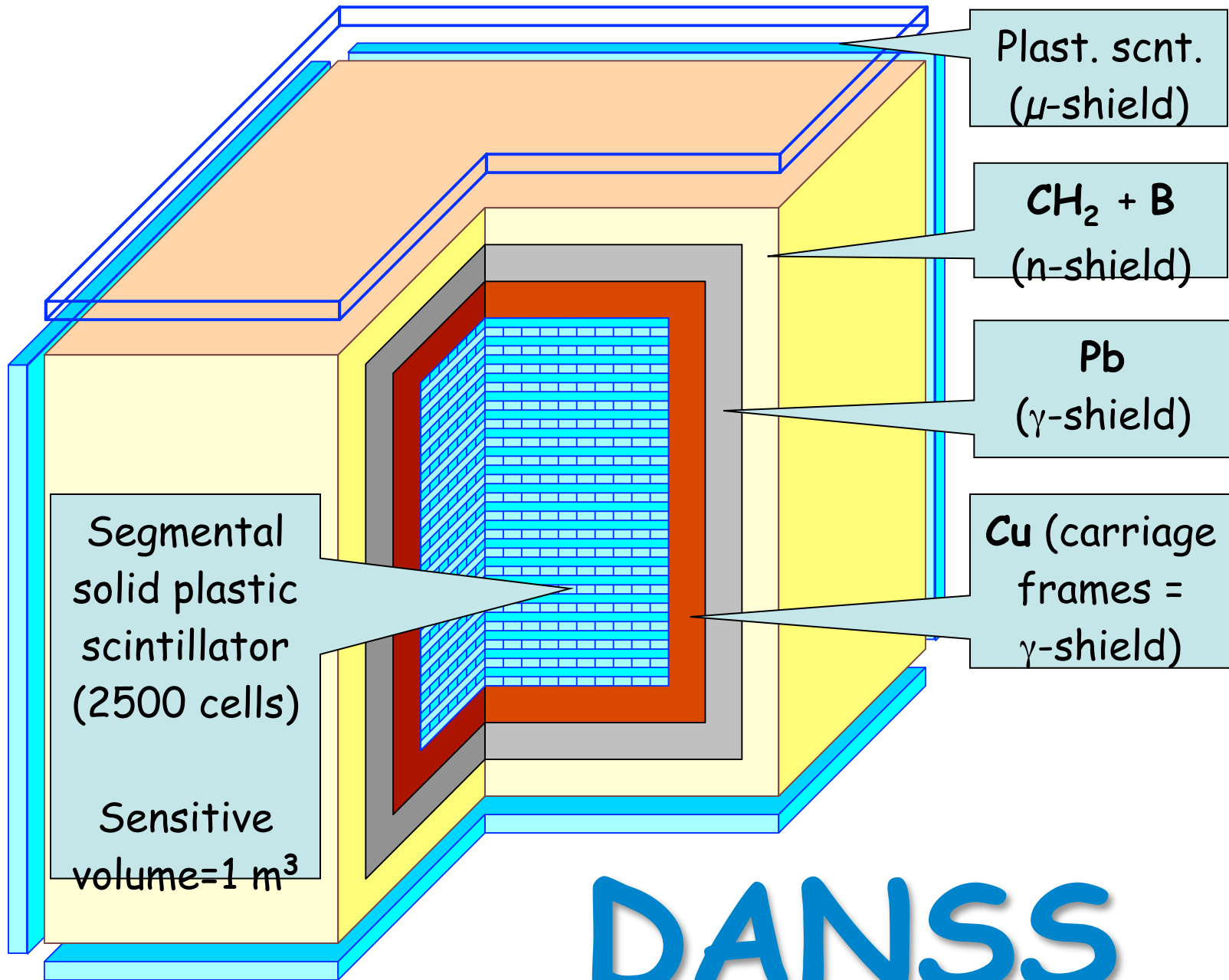




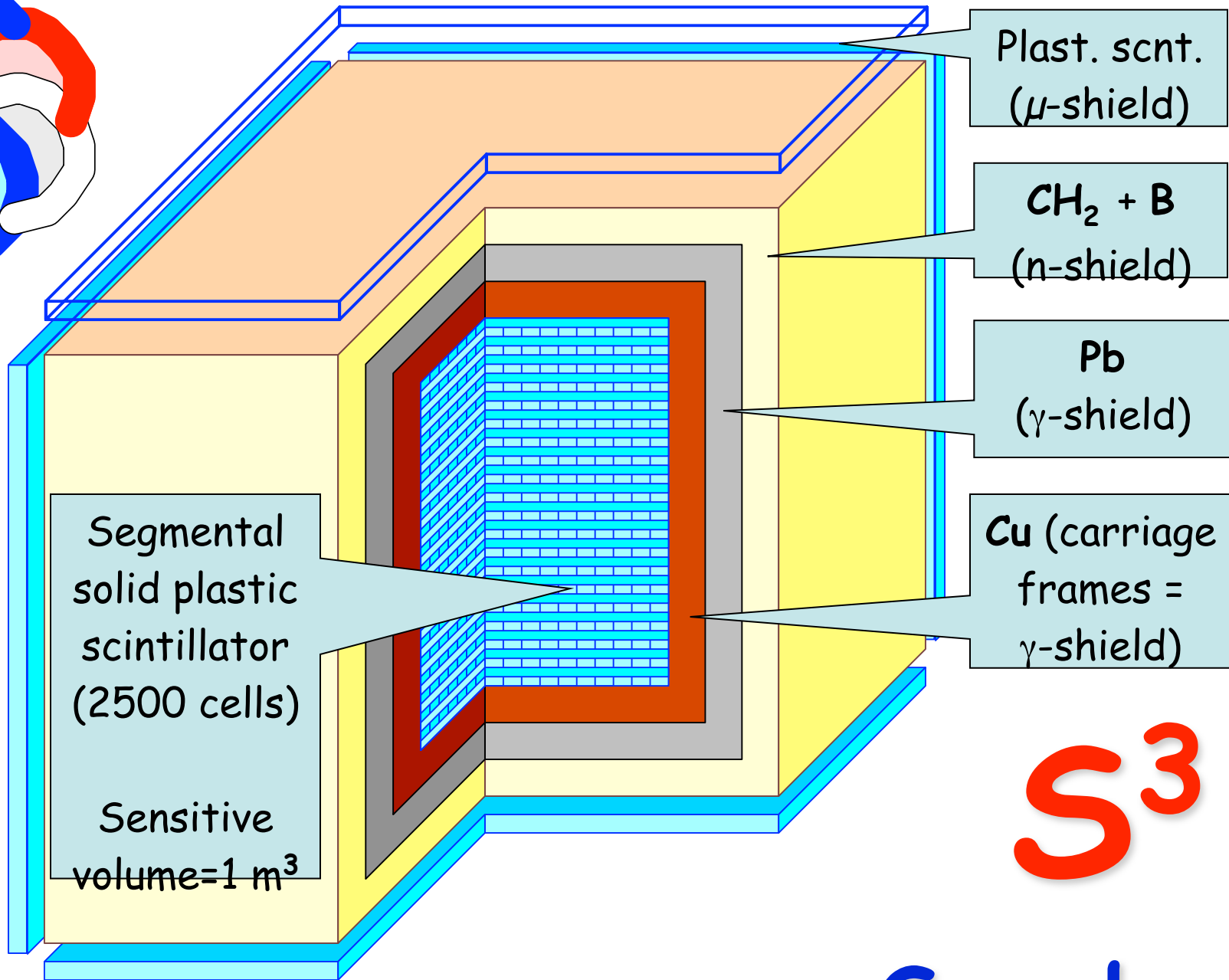
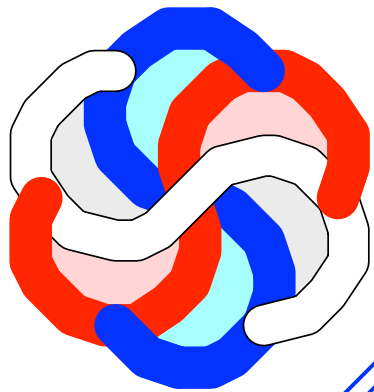








DANSS

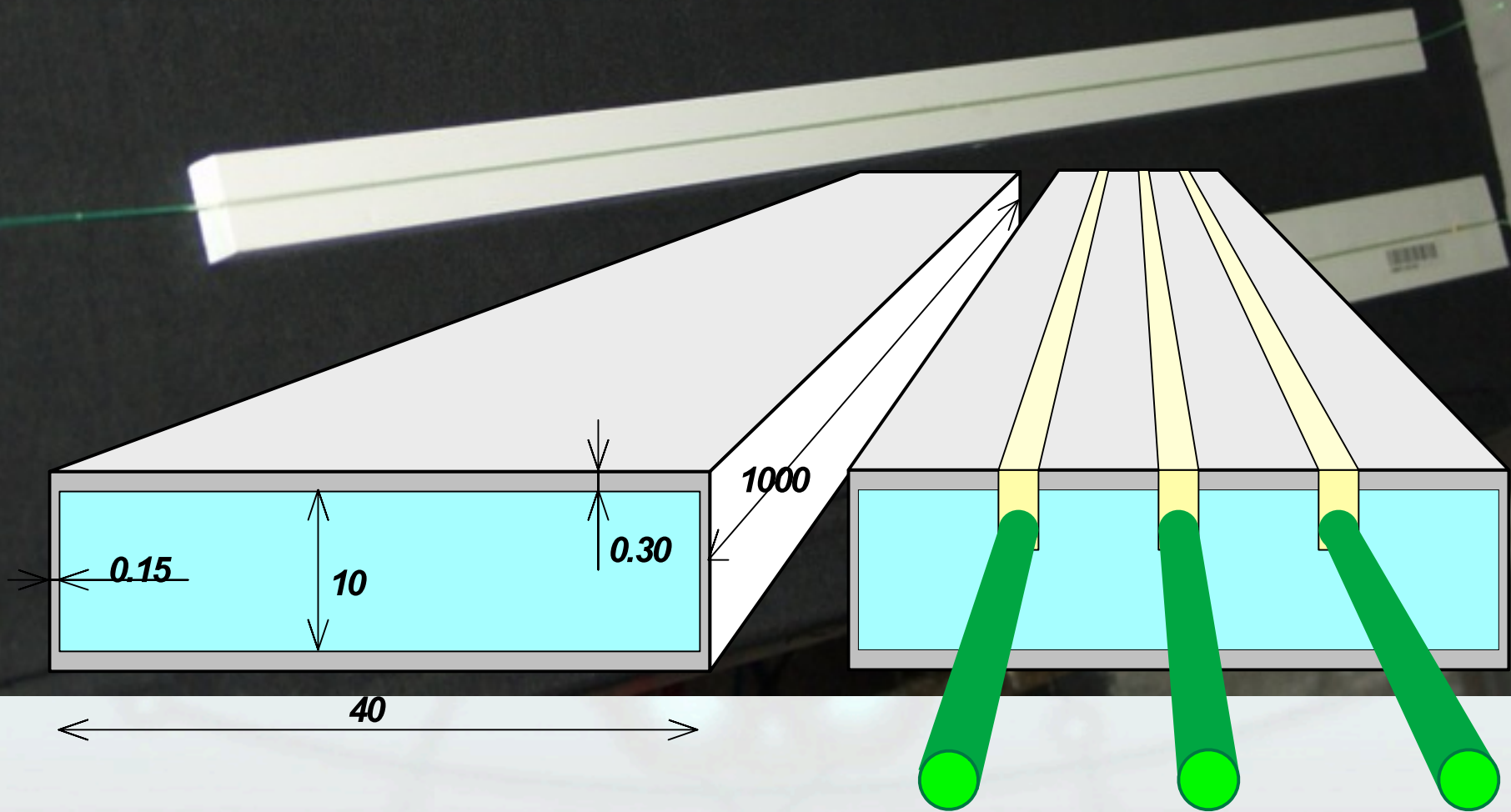


S³

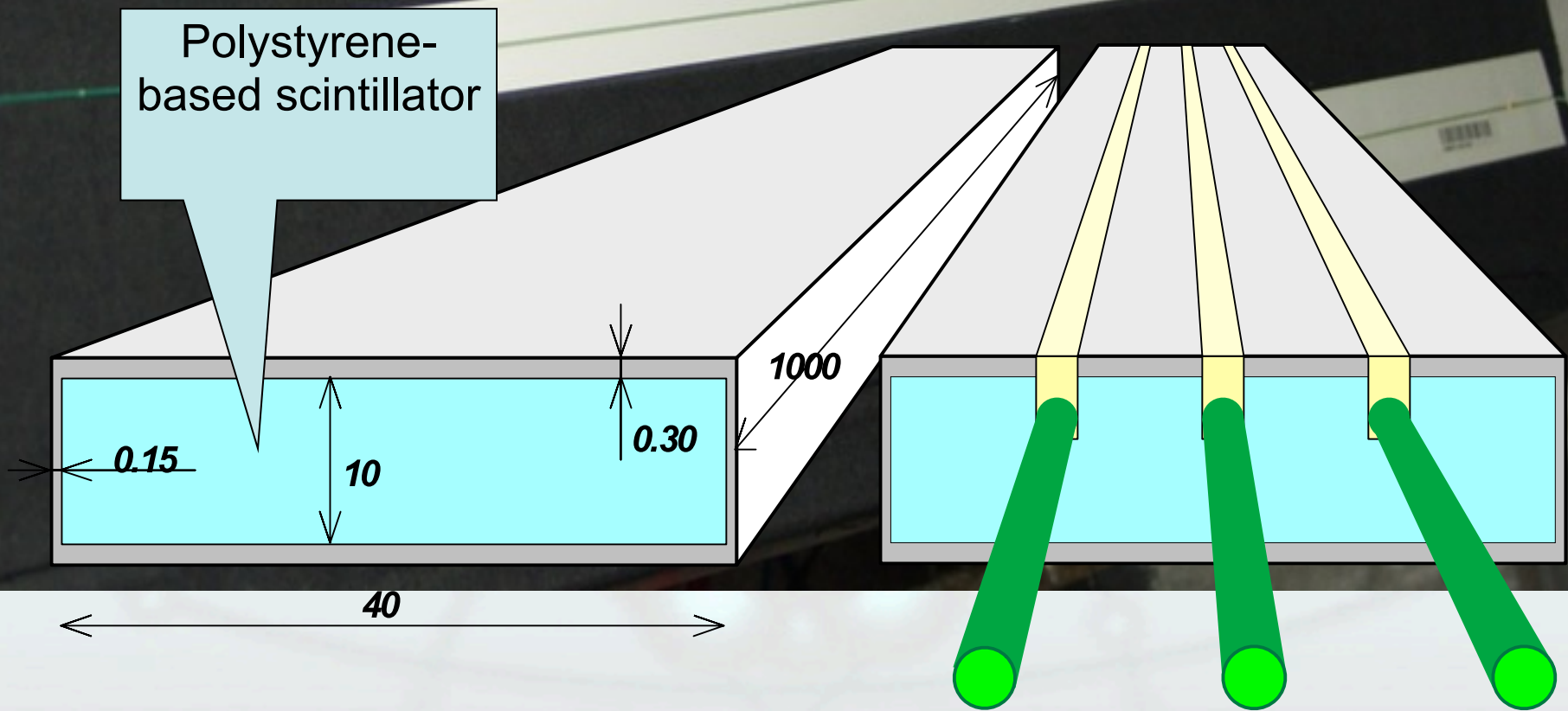
S-cube



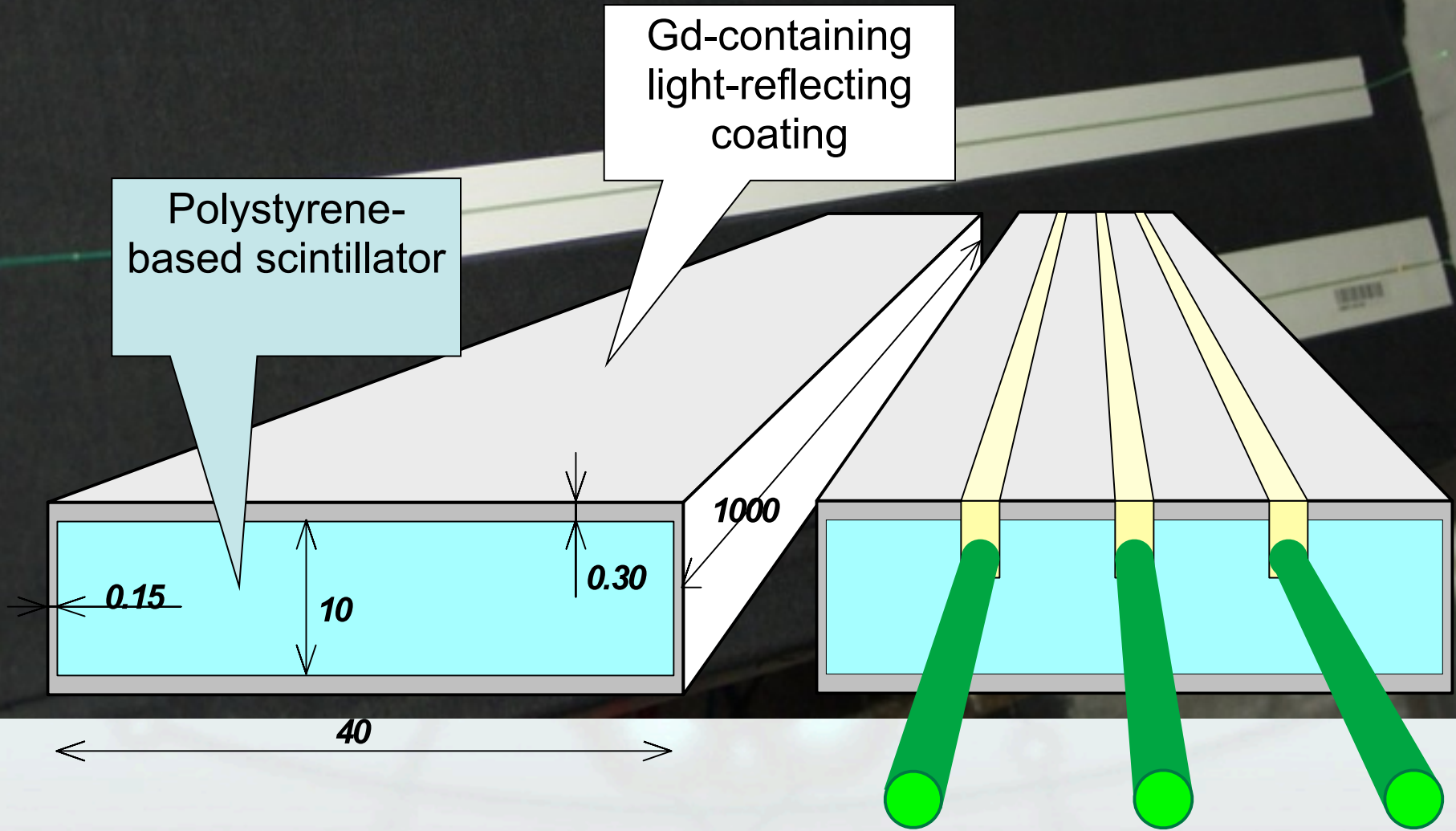
Basic element -



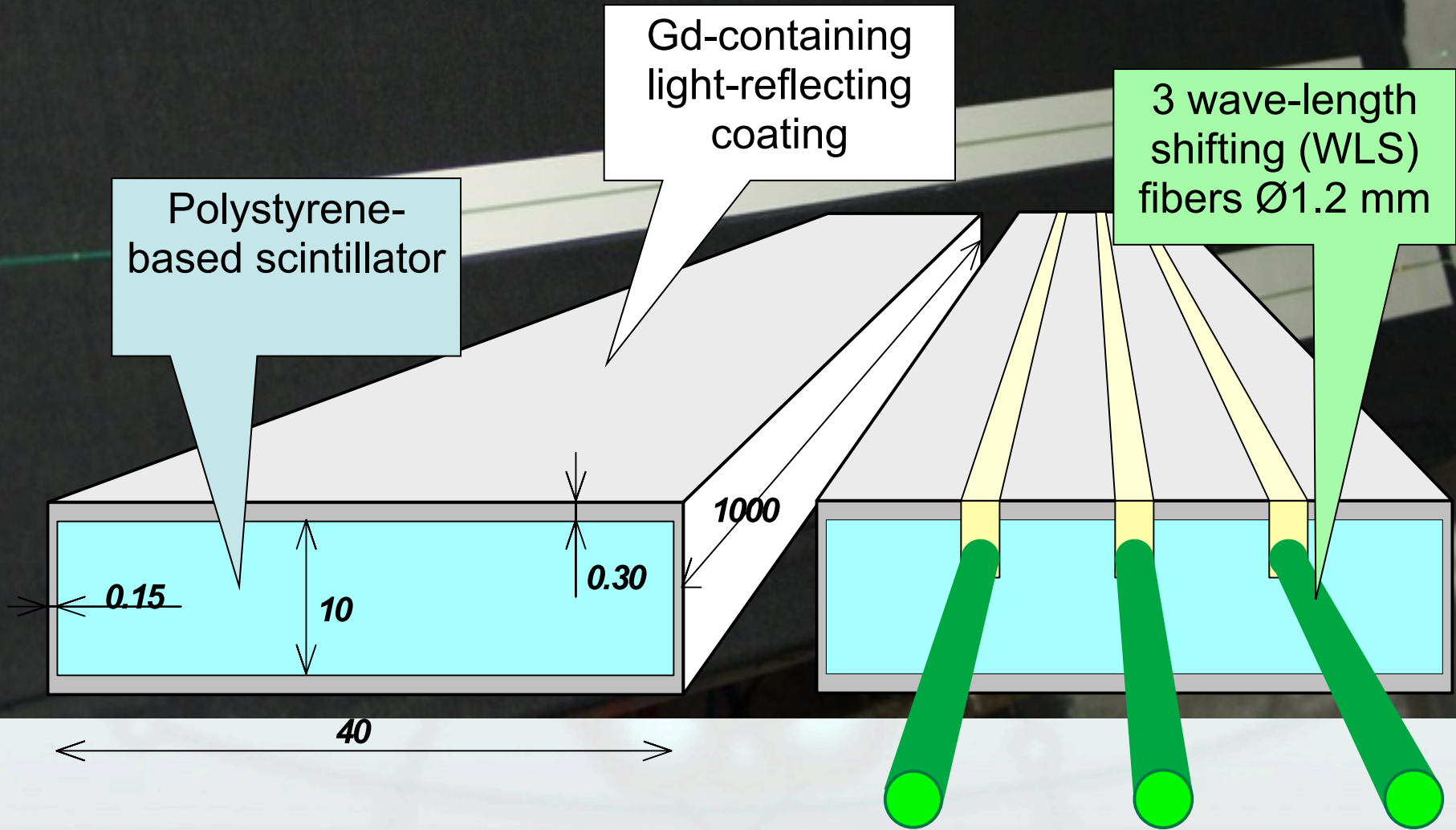
Basic element -



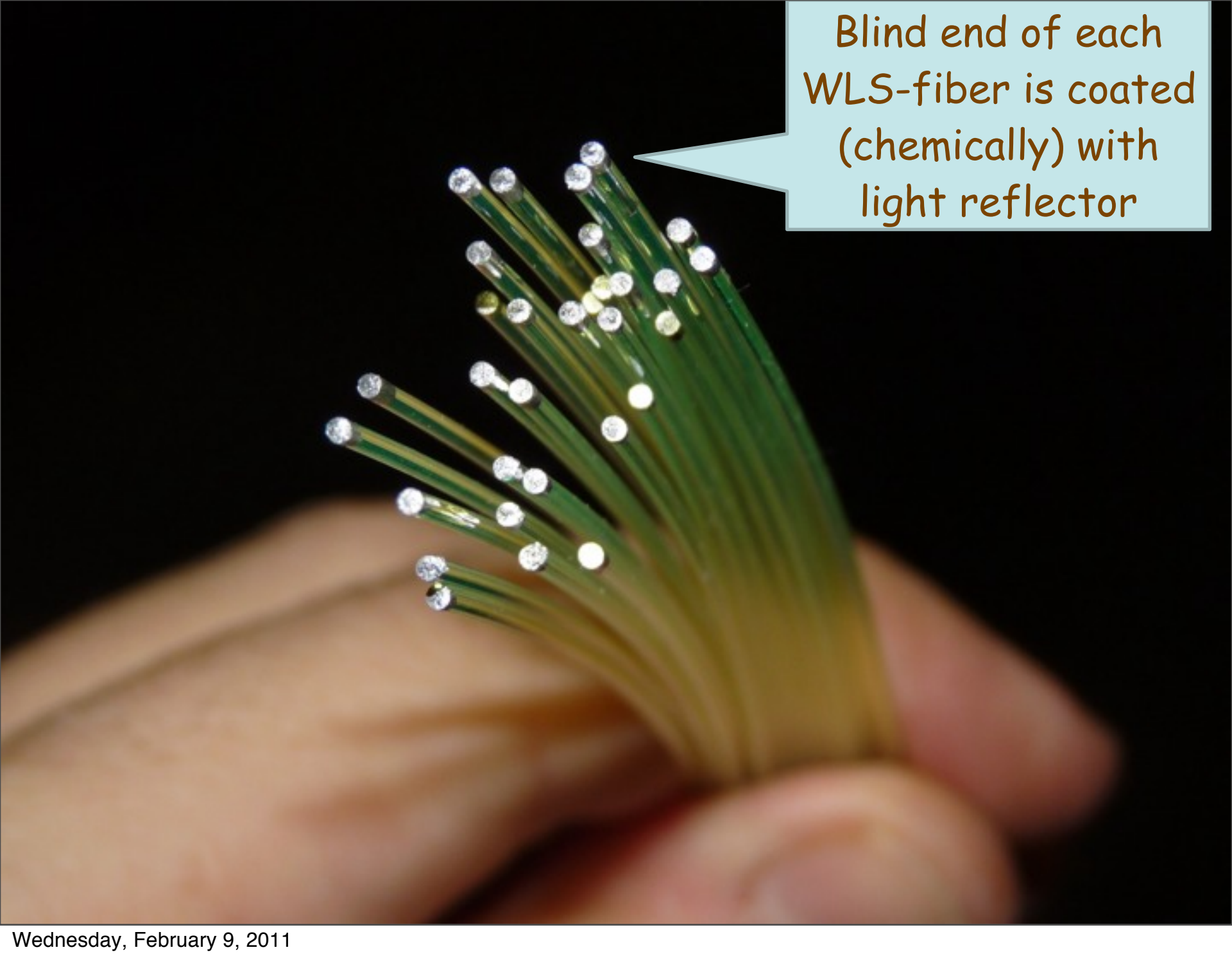
Basic element -



Basic element -



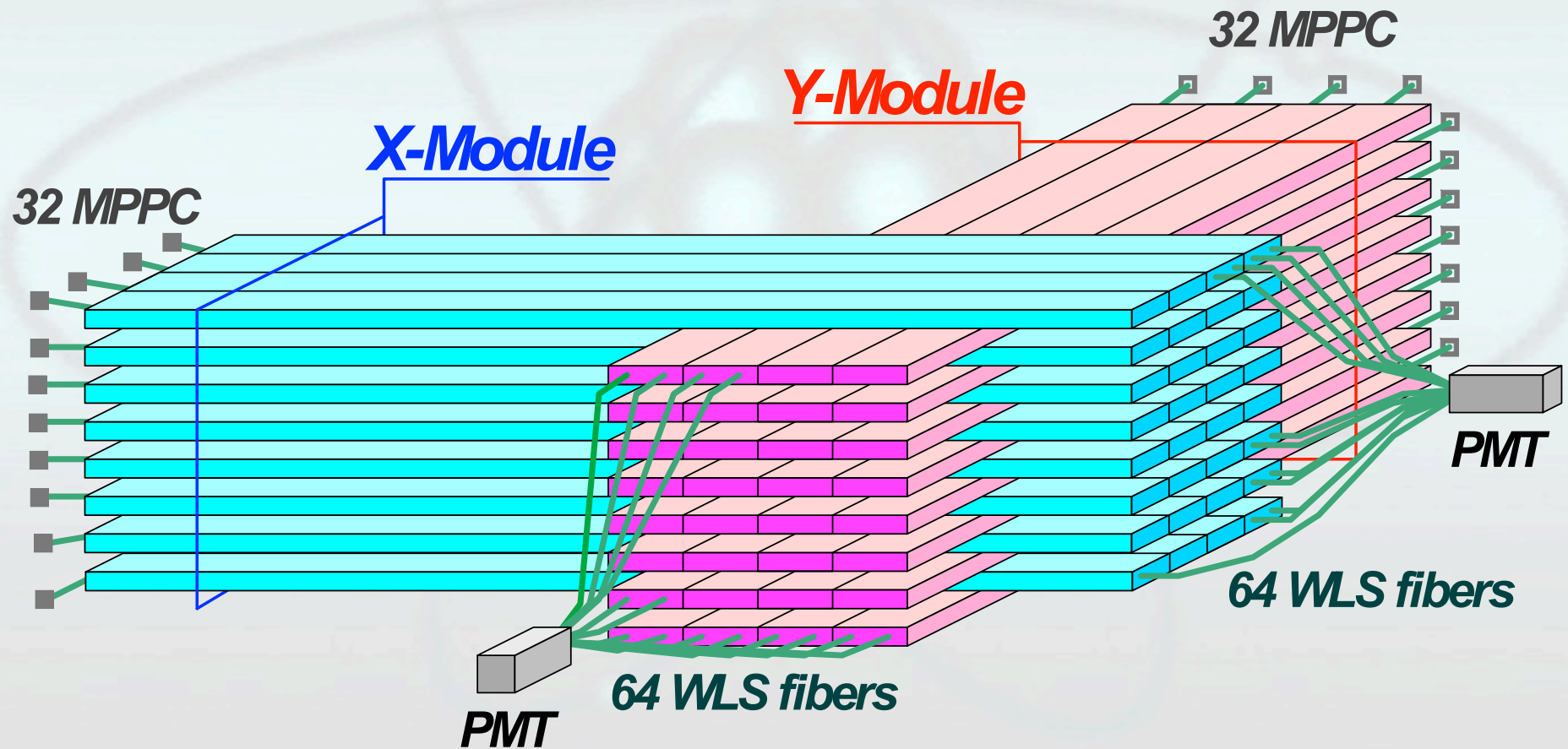
Basic element -

A close-up photograph of a hand holding a bundle of approximately 20-30 thin, green, cylindrical fibers. The fibers are held together at one end and fan out towards the other. The ends of the fibers are coated with a reflective, metallic material, likely silver or aluminum, which appears as small, bright, circular spots. The background is dark, making the green fibers and their reflective ends stand out. A light blue callout box with a pointer is positioned in the upper right corner, containing text that explains the reflective coating.

Blind end of each
WLS-fiber is coated
(chemically) with
light reflector

Modular structure of the detector

A number of strips are combined into intercrossing X- and Y-modules.



Prototype: 2 parallel modules (no Gd-coating)



Light collection

Light collection

- **One PMT per Module (25-50 strips):**

Light collection

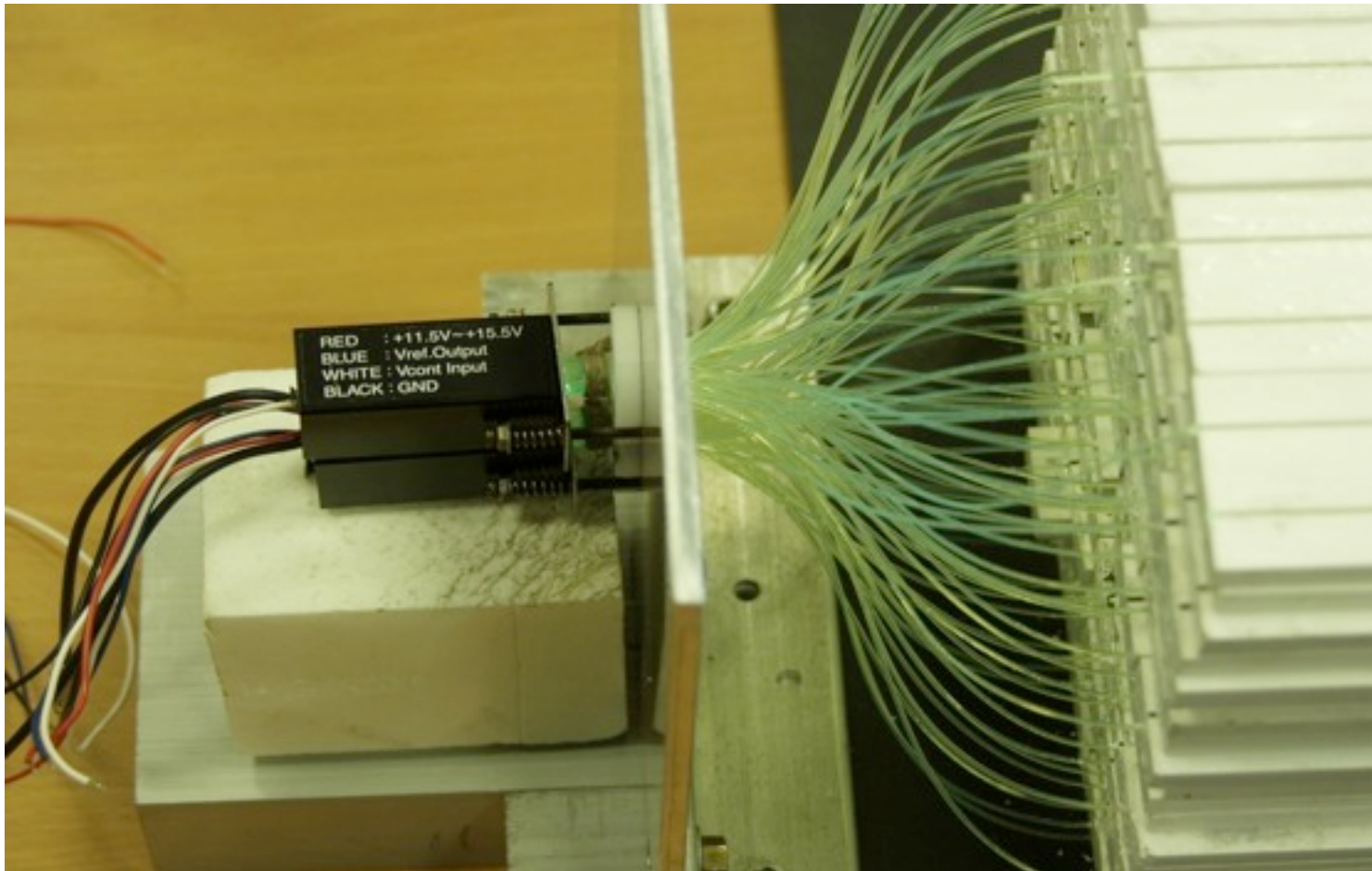
- **One PMT per Module (25-50 strips):**

measure total energy of the photon burst

Light collection

- **One PMT per Module (25-50 strips):**

*measure total energy of the photon burst
produce “hard trigger”*



Light collection

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produce "hard trigger"

Light collection

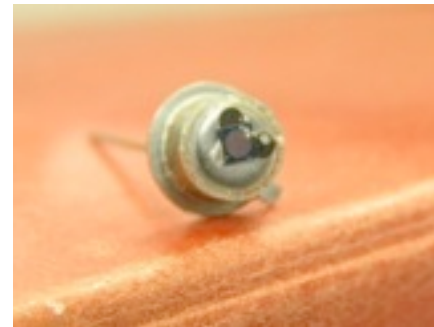
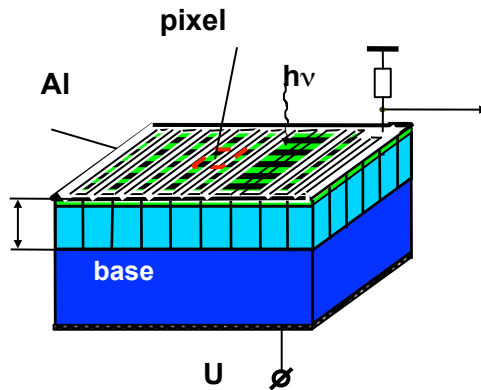
- **One PMT per Module (25-50 strips):**
measure total energy of the photon burst
produce "hard trigger"
- **Individual photo-diod (SiPM or MPPC) per each strip:**

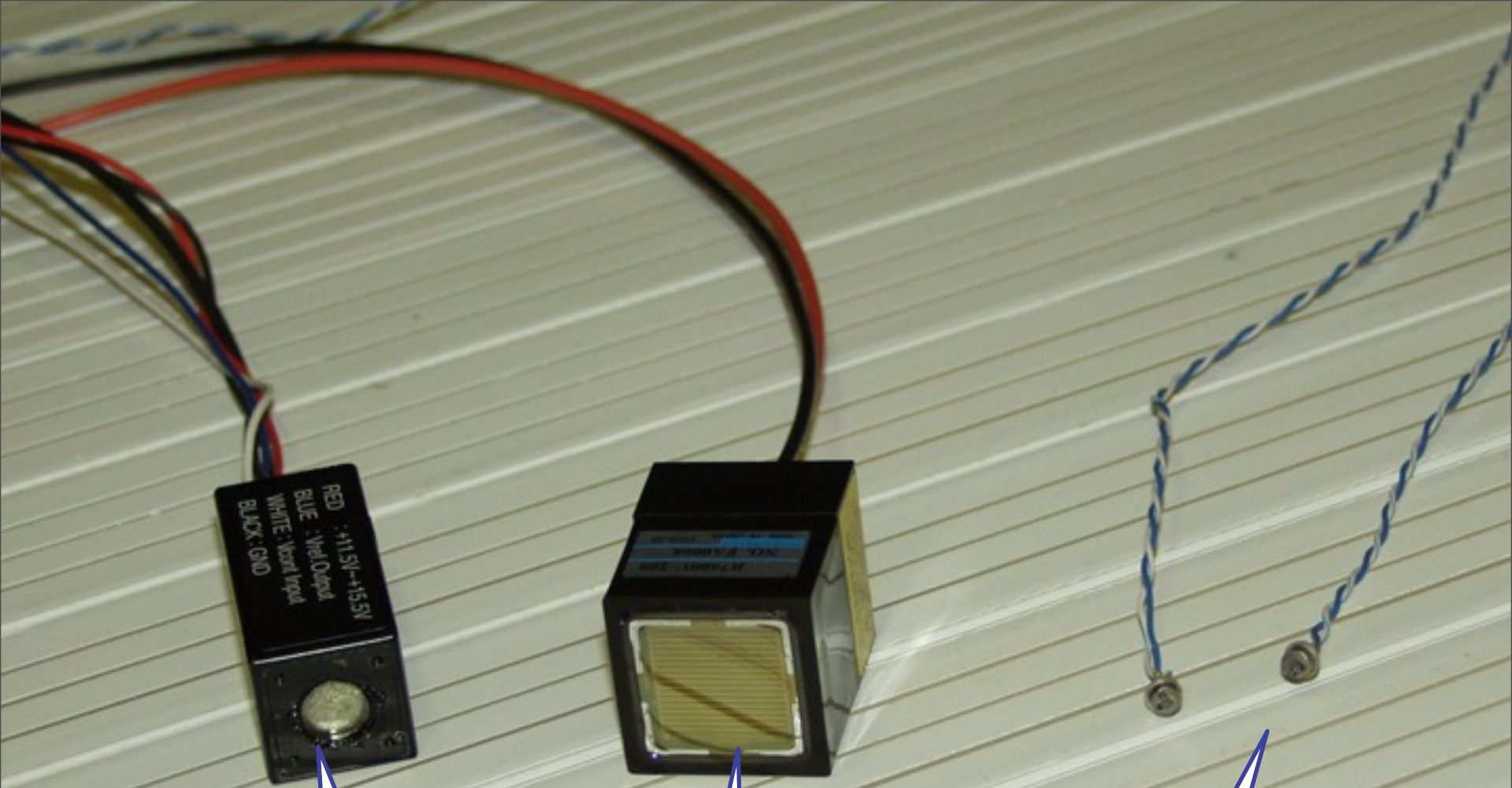
Light collection

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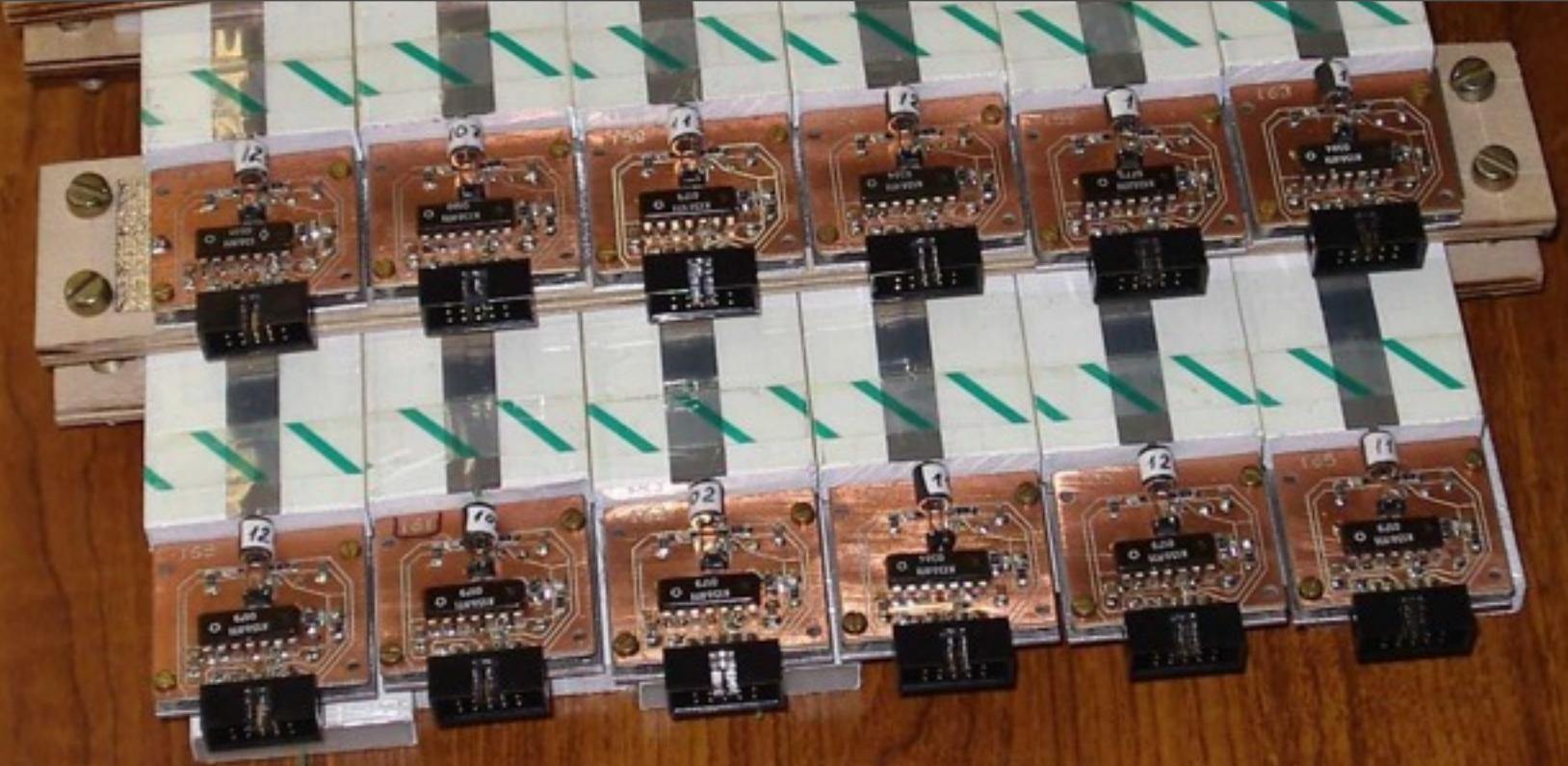




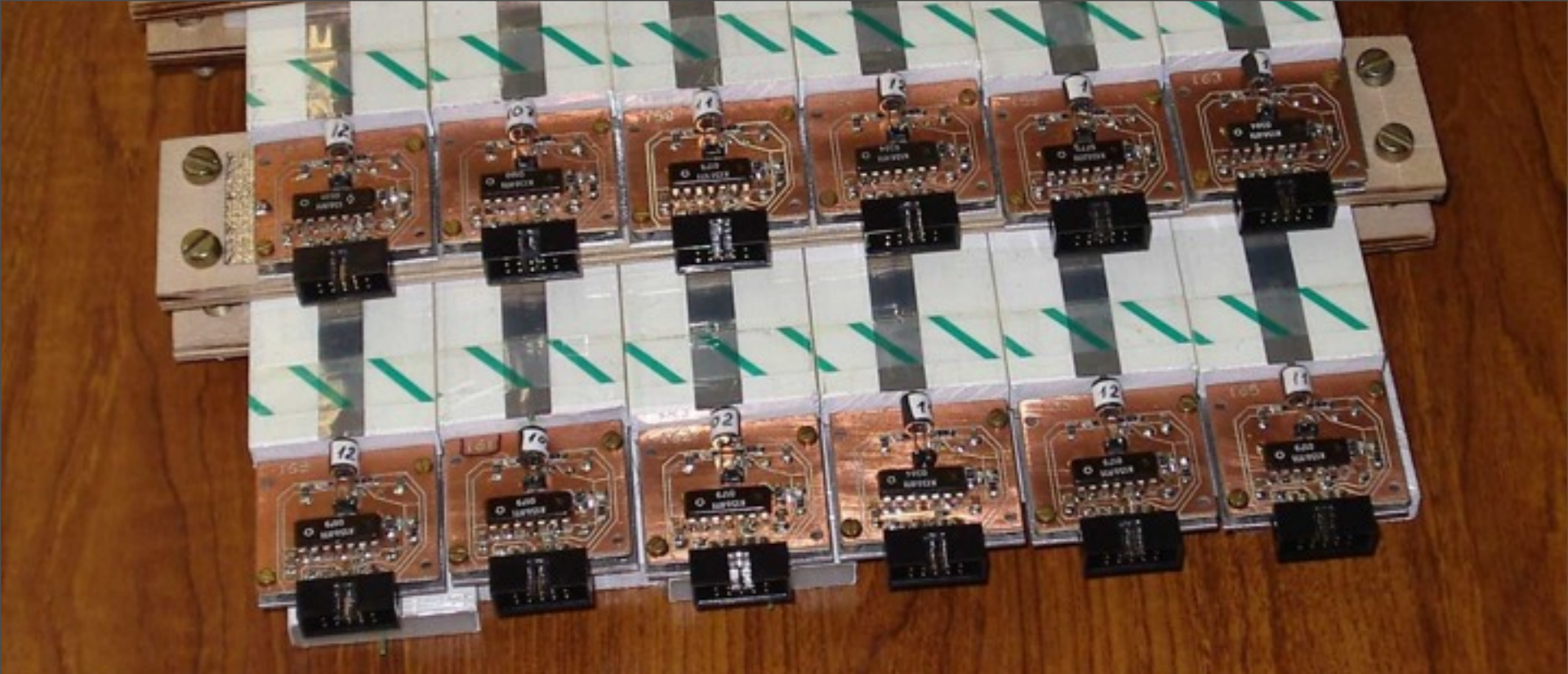
PMT+HV
(Japan)

PMT
(Japan)

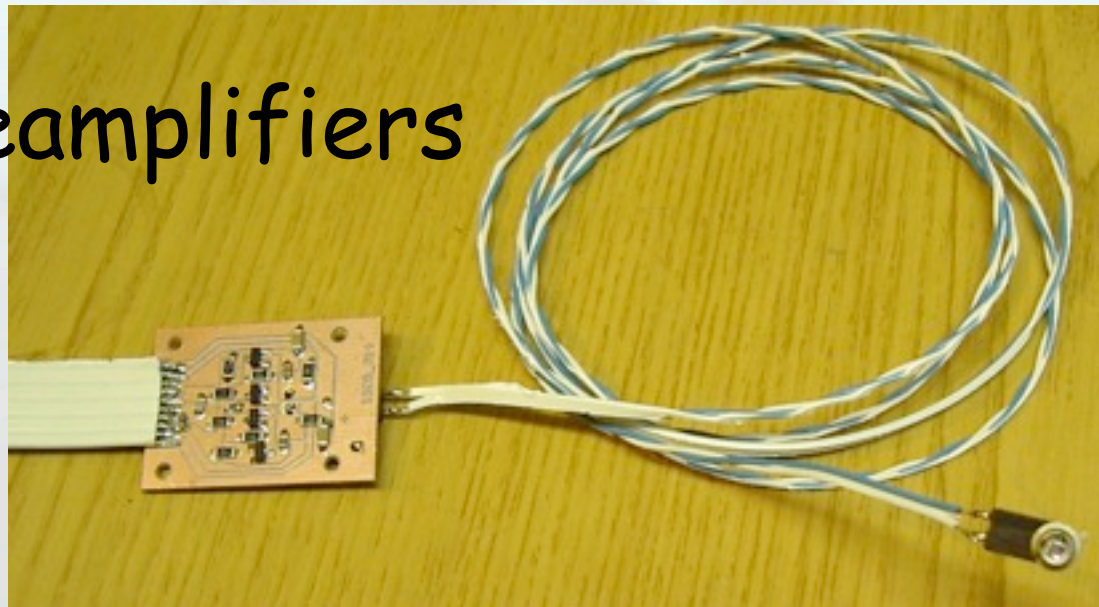
MPPC
(Russia)

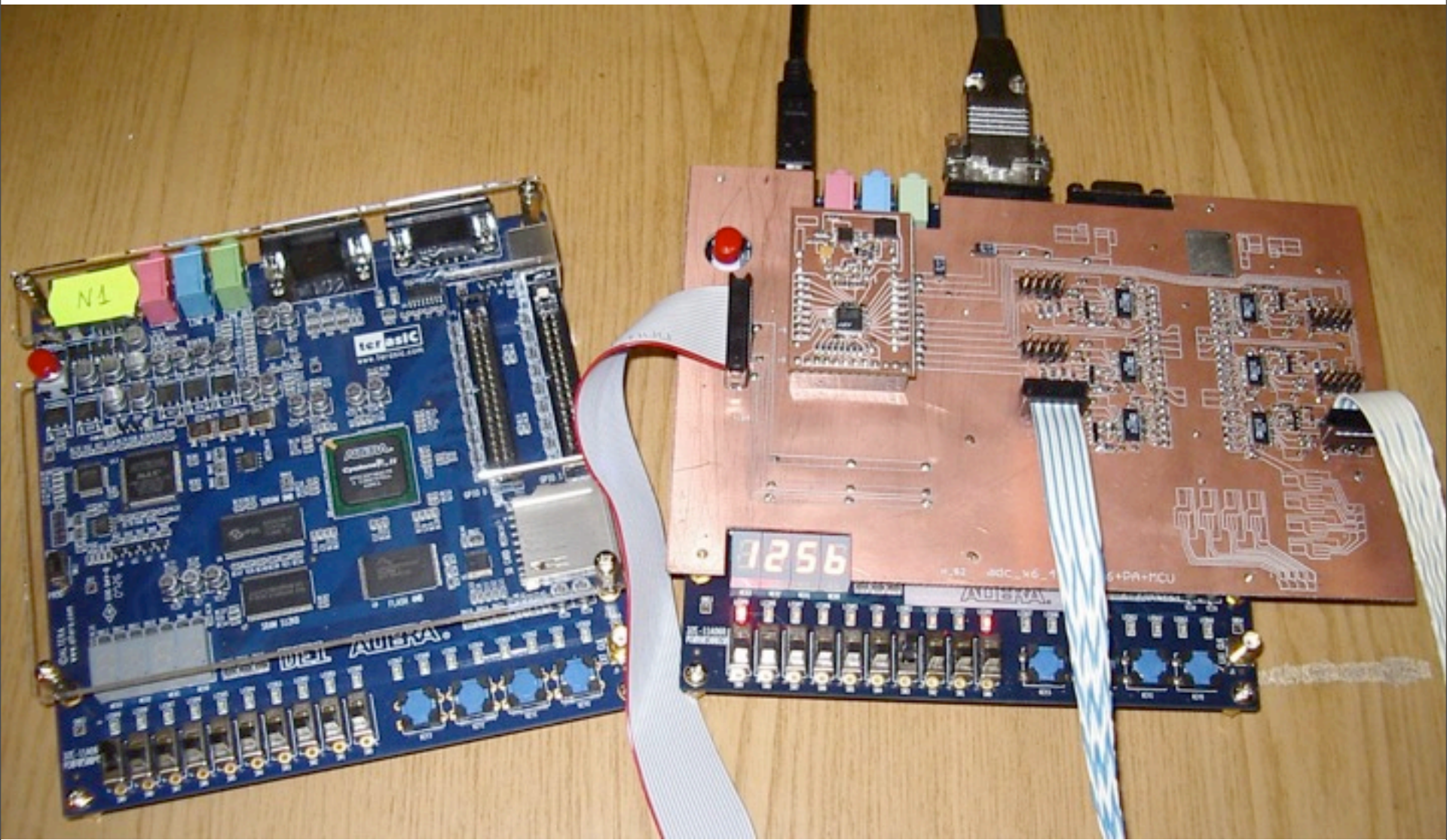


MPPC preamplifiers

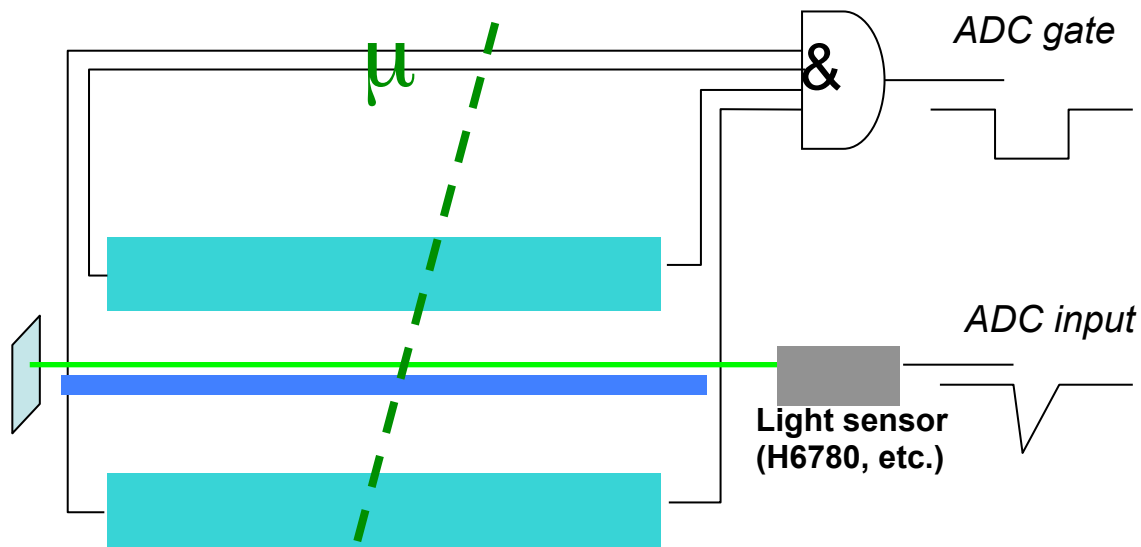


MPPC preamplifiers



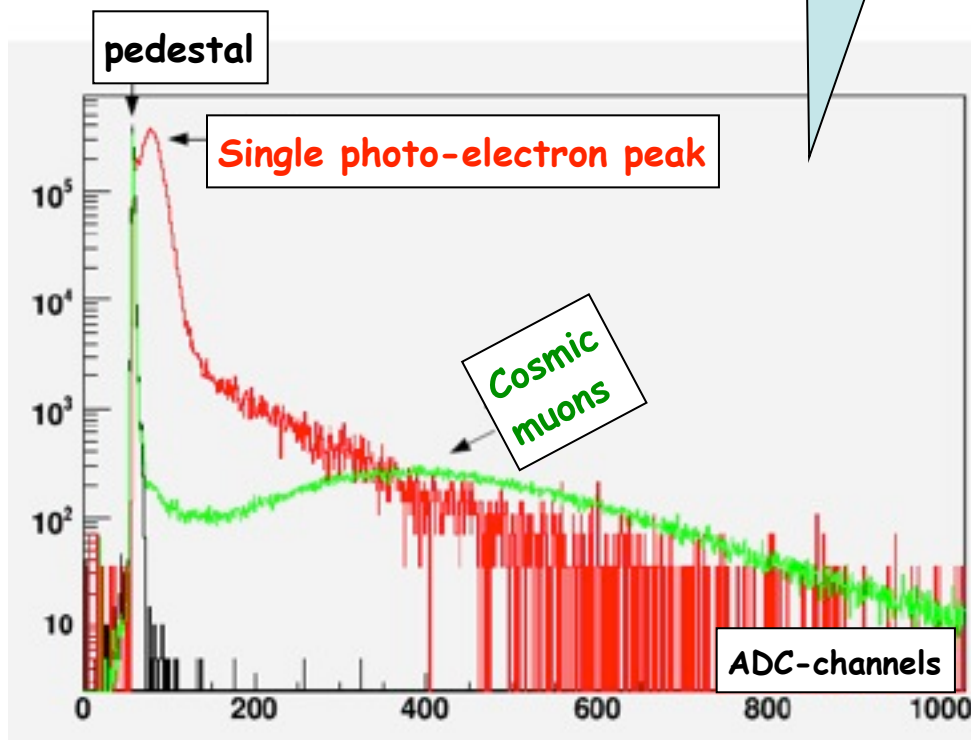


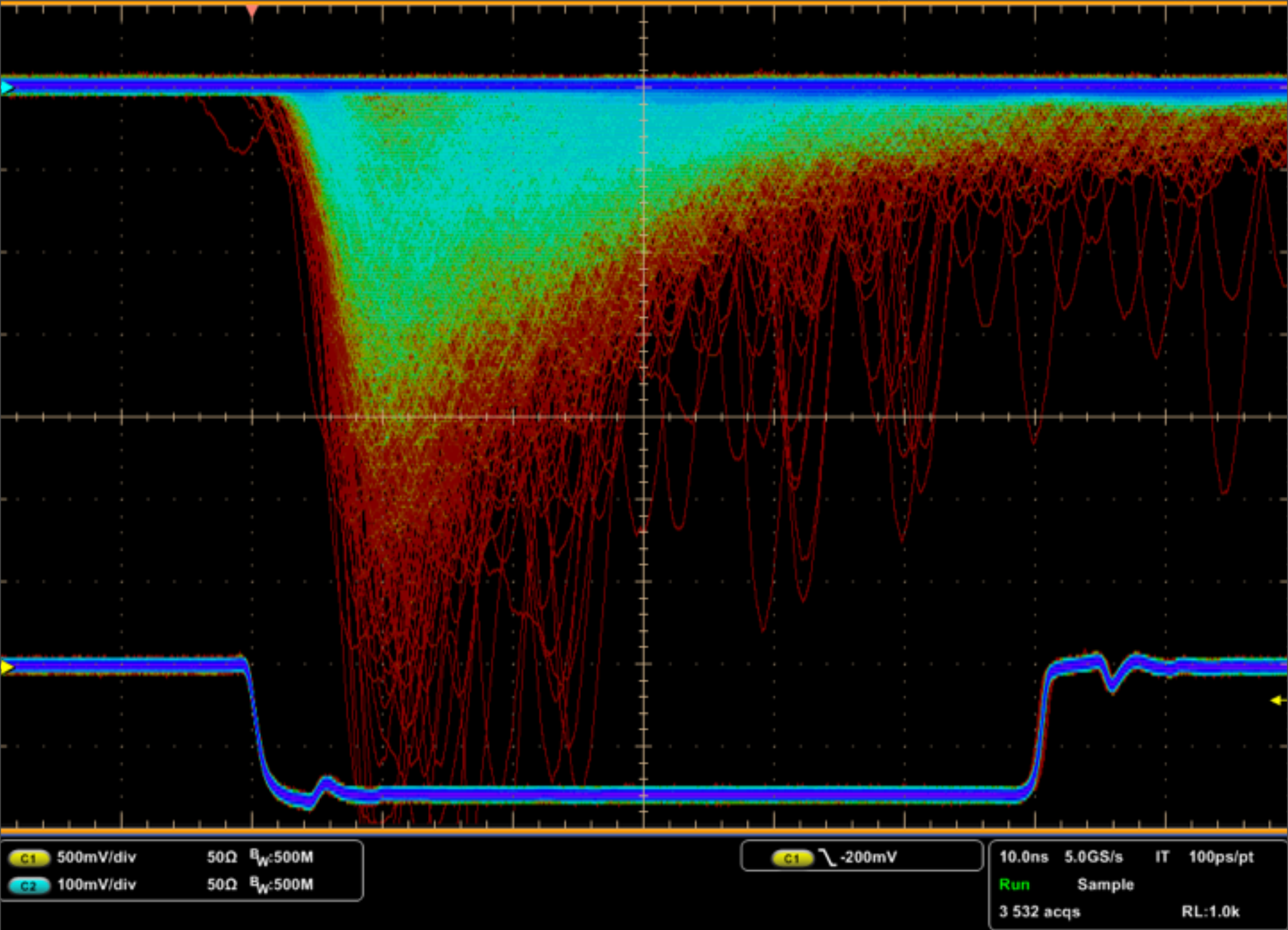
Flash ADC etc.



Typical energy spectrum

Test bench for strips, WLS-fibers and photo sensors

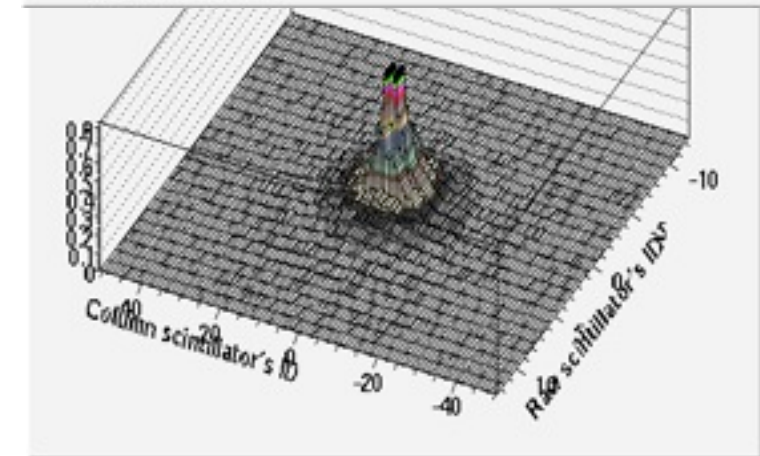
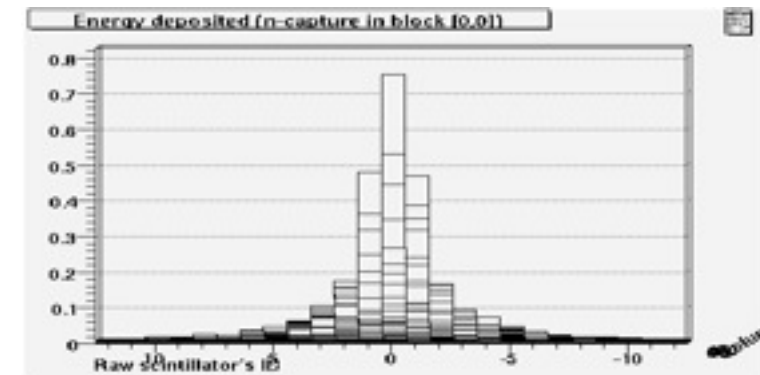
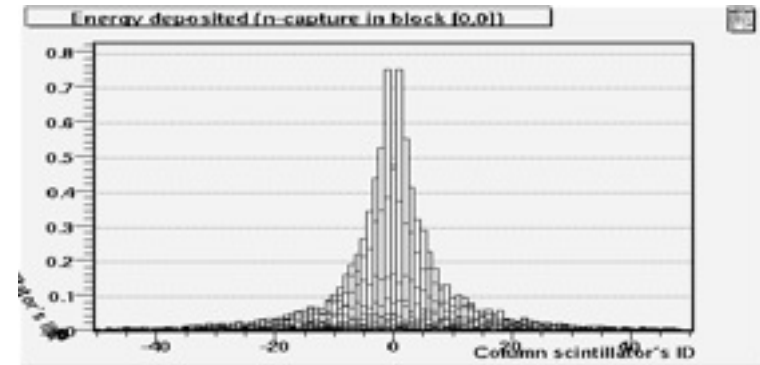
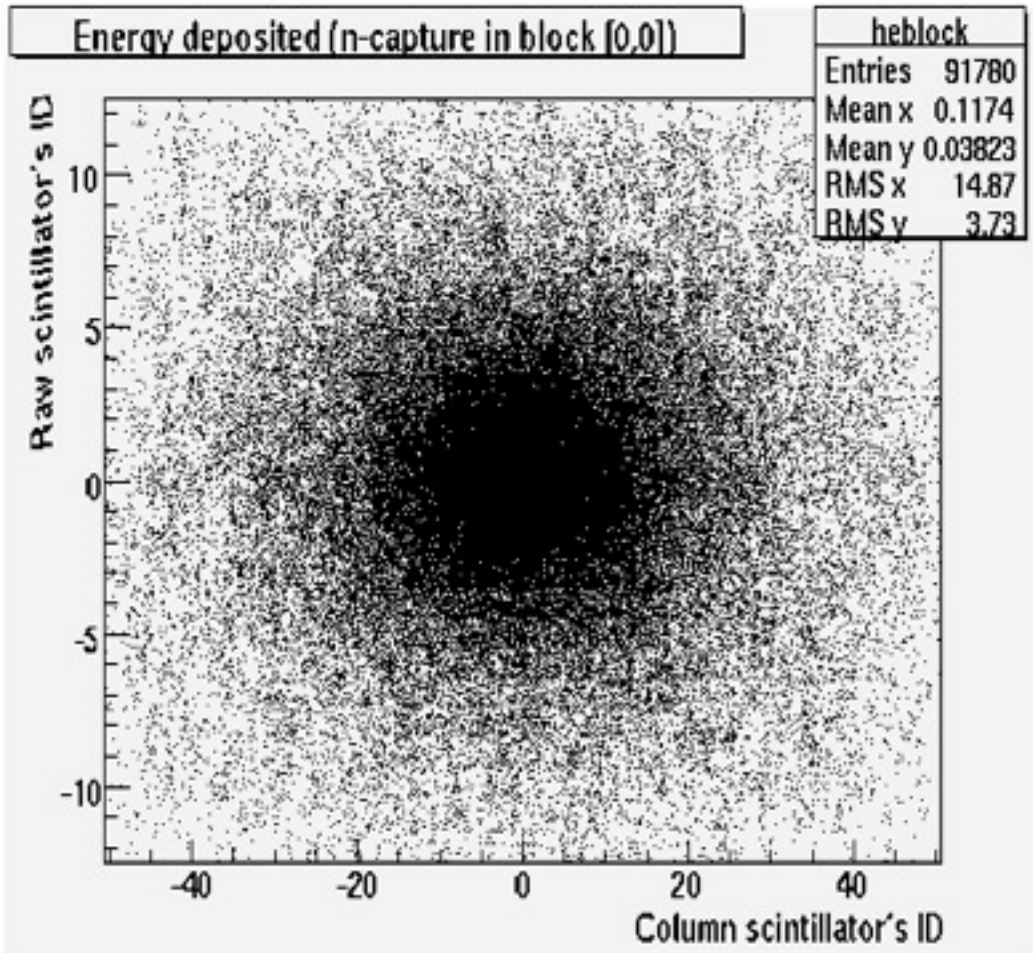




Monte Carlo simulations

Example:

Space distribution of the energy deposit after the $^{157}\text{Gd}(n, \gamma)$ reaction



End of 2009:

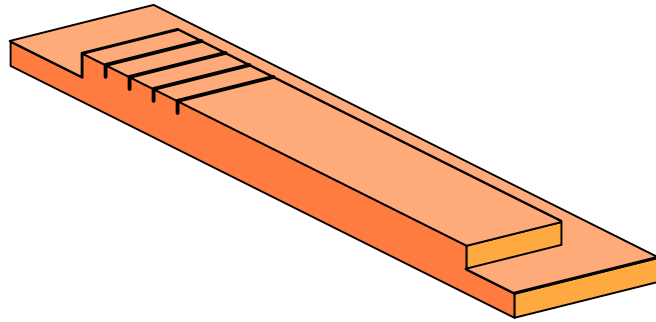
Start

the real

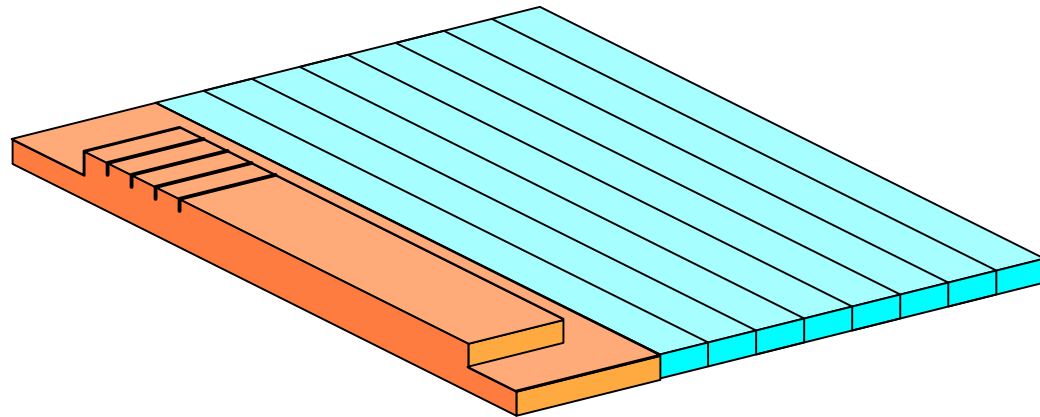
creation!

Carriage frames

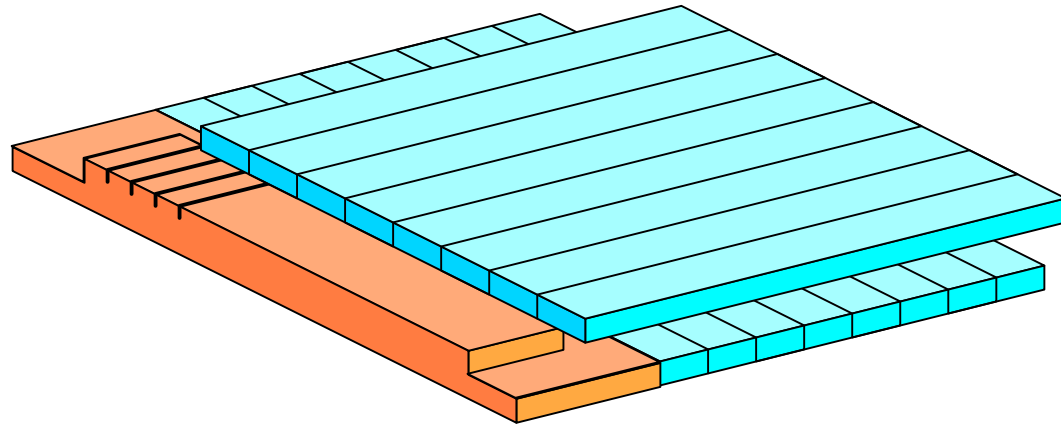
Carriage frames



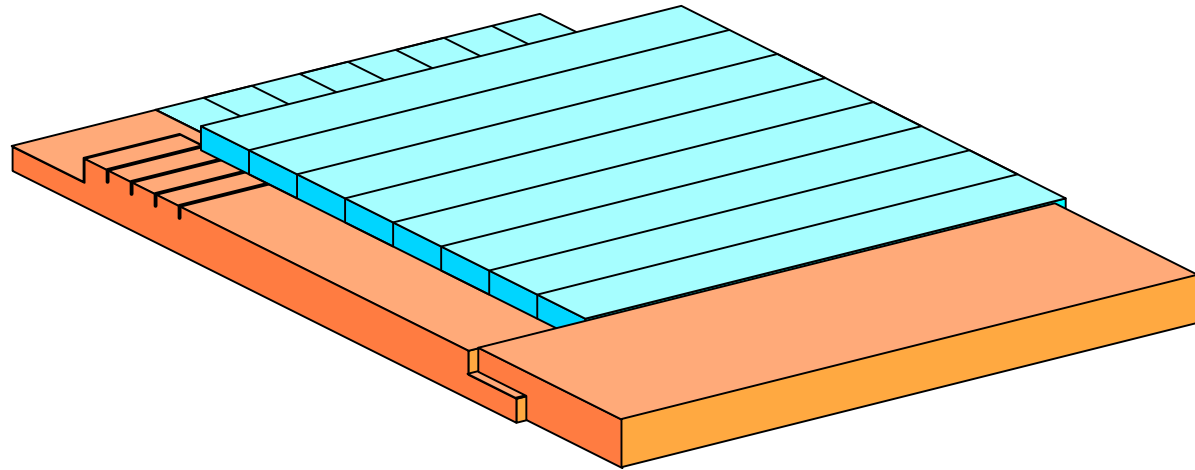
Carriage frames

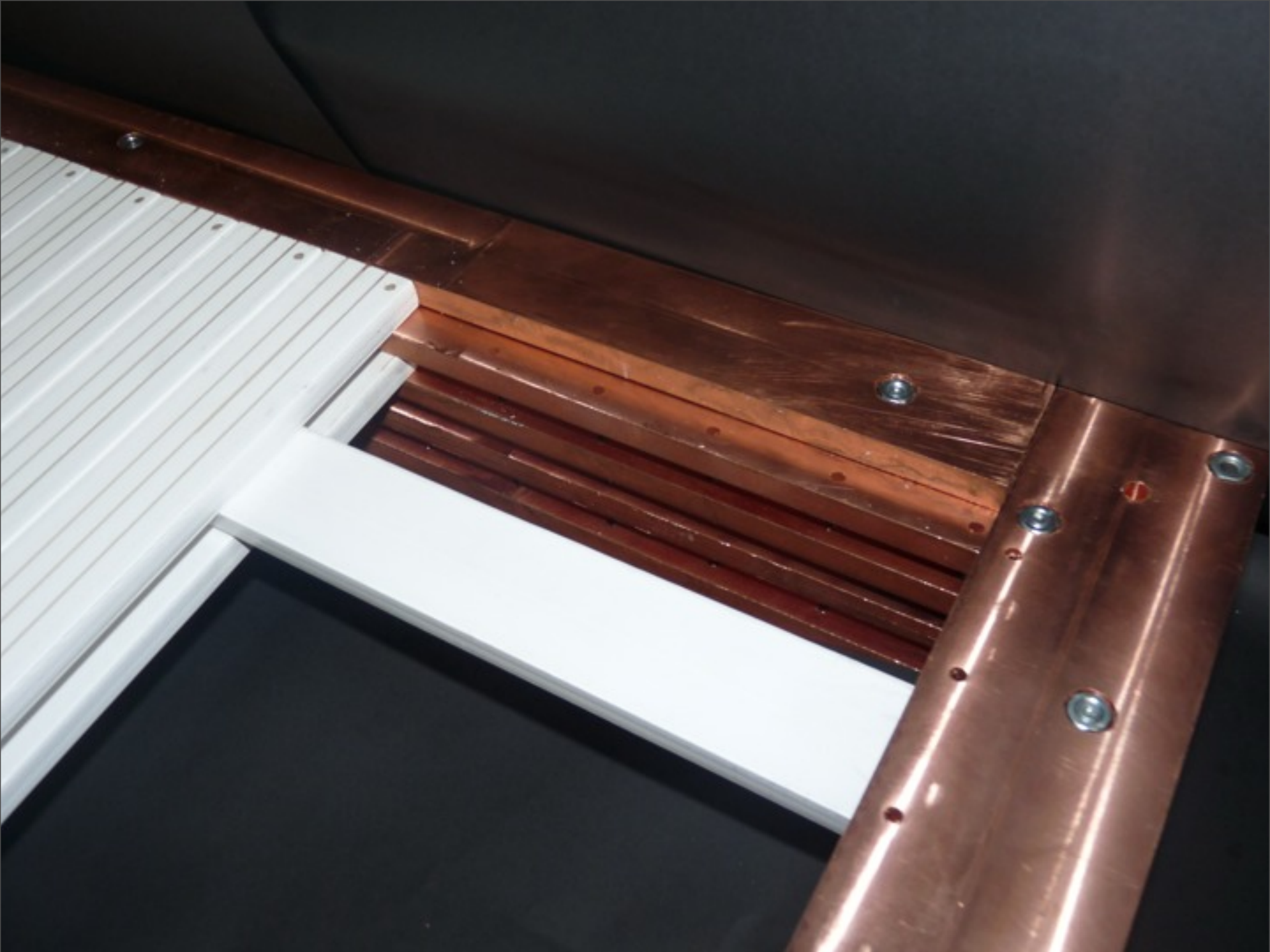


Carriage frames

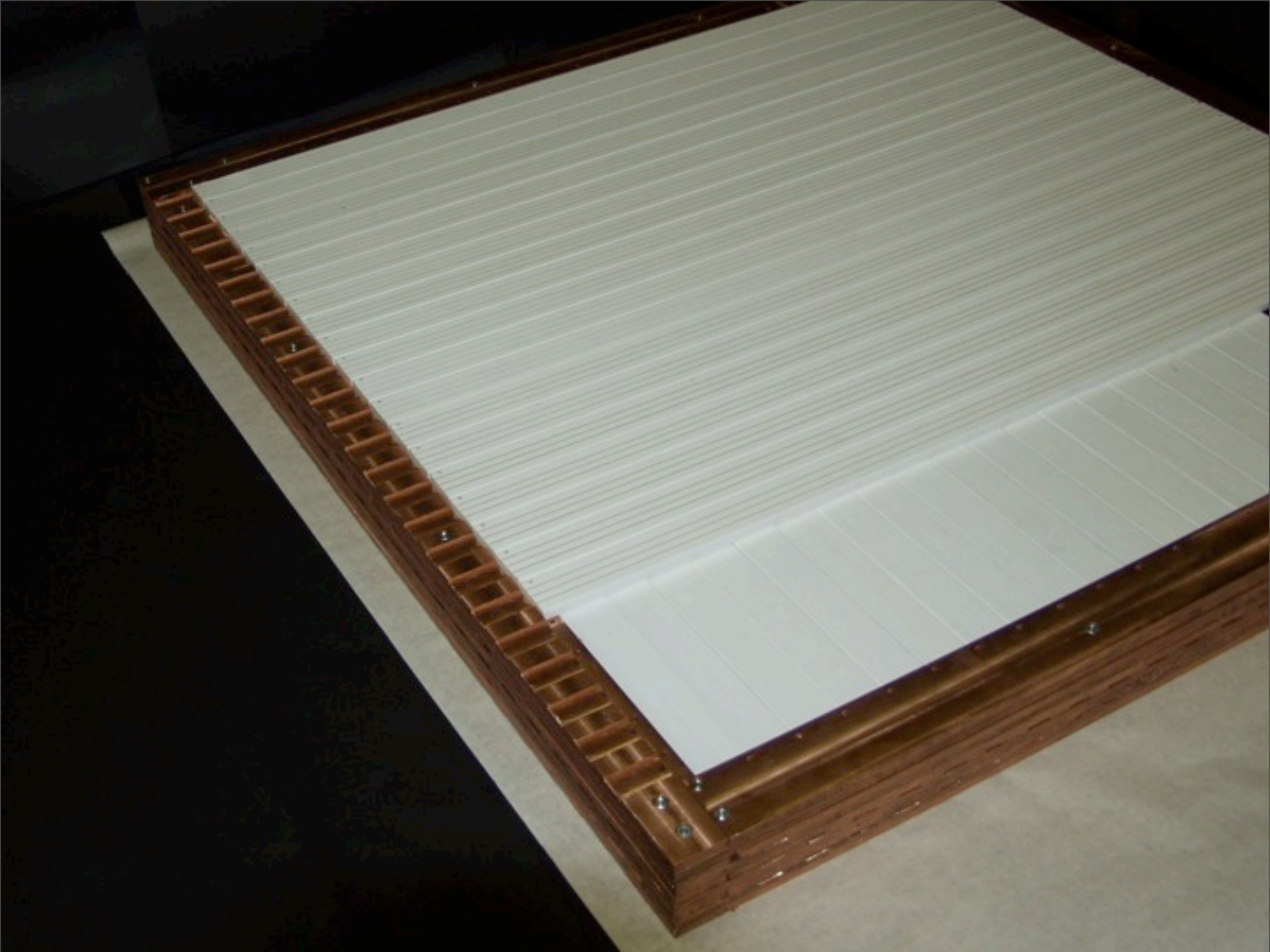


Carriage frames

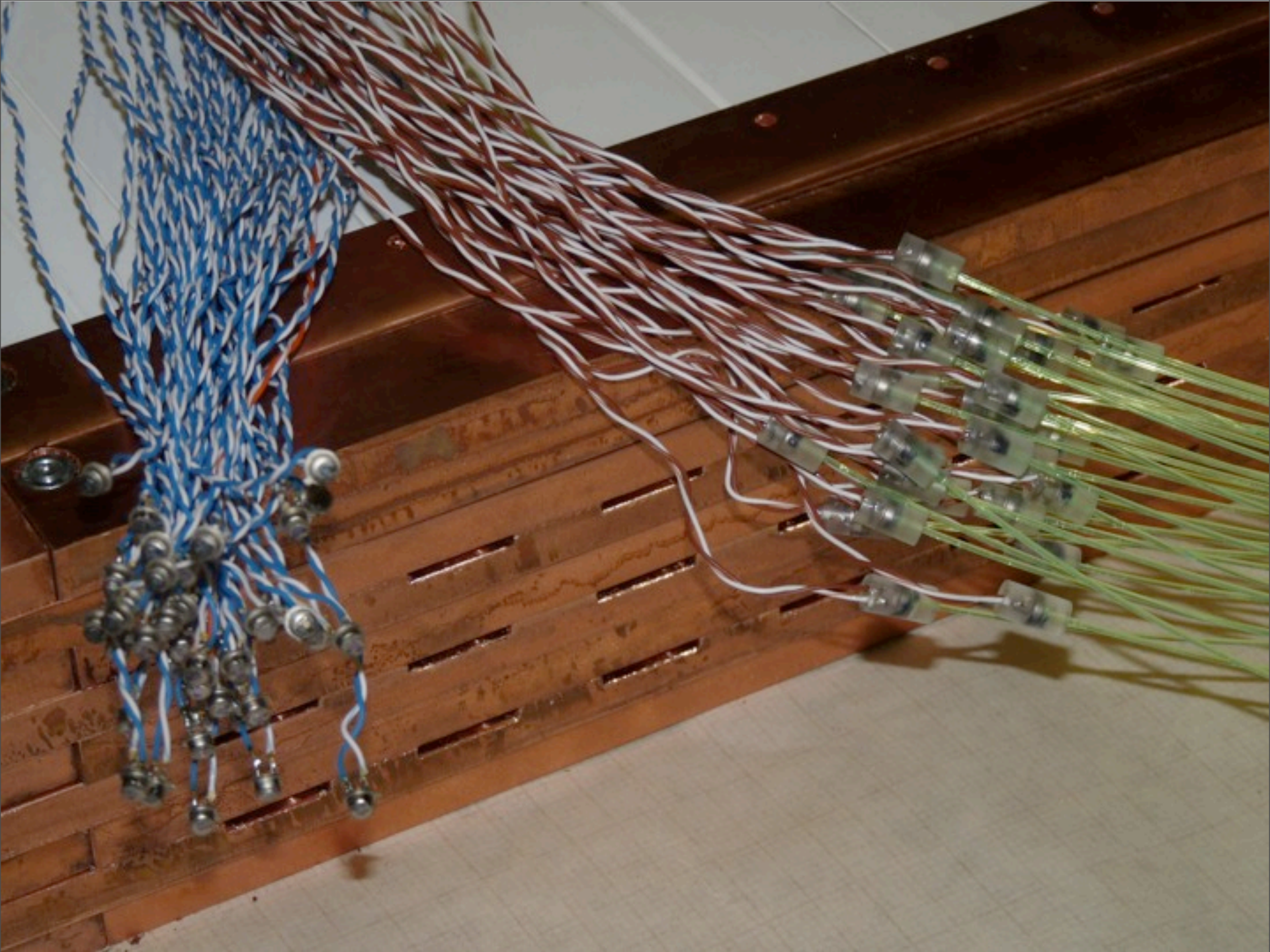




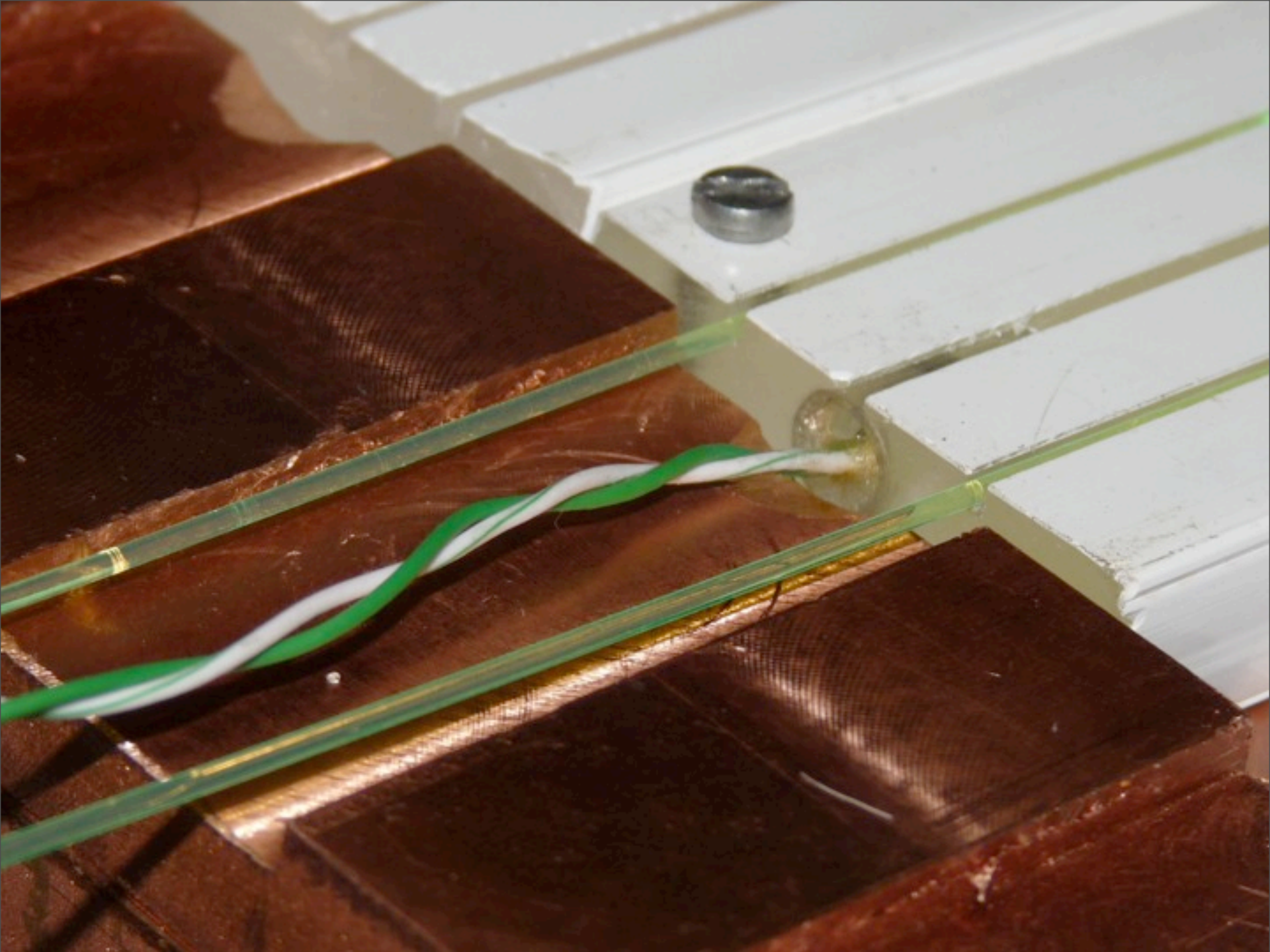
Wednesday, February 9, 2011



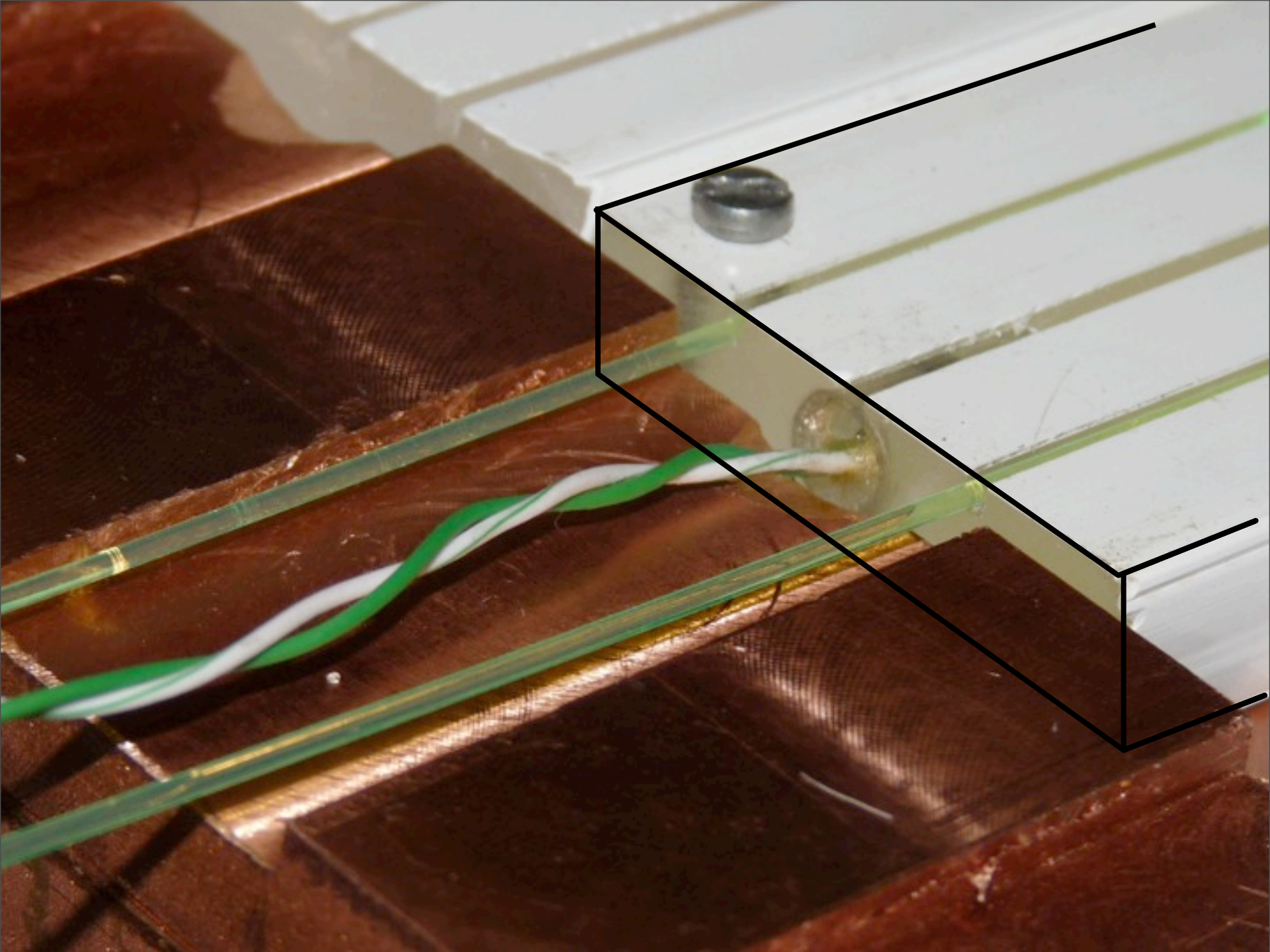
Wednesday, February 9, 2011



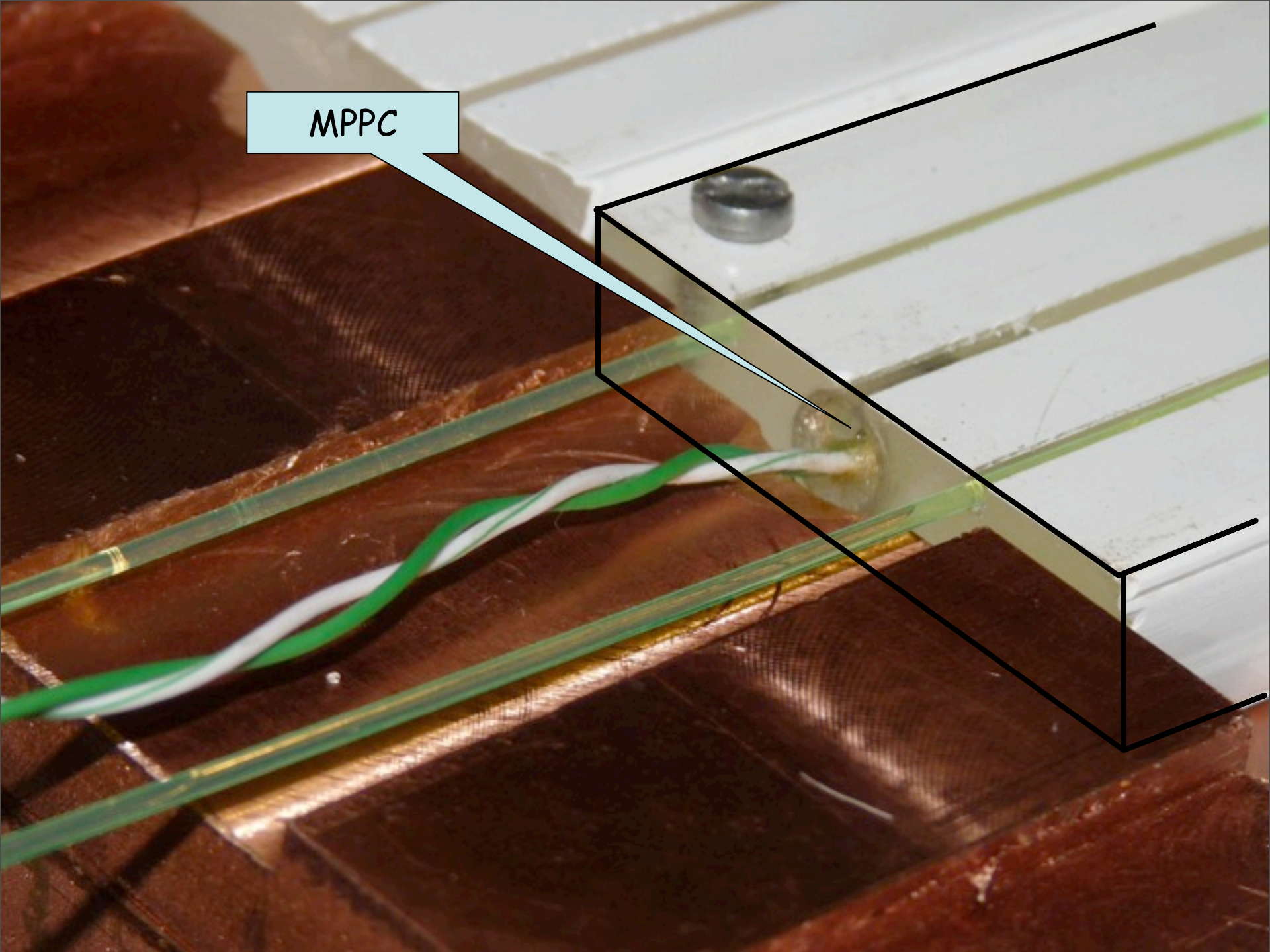
Wednesday, February 9, 2011



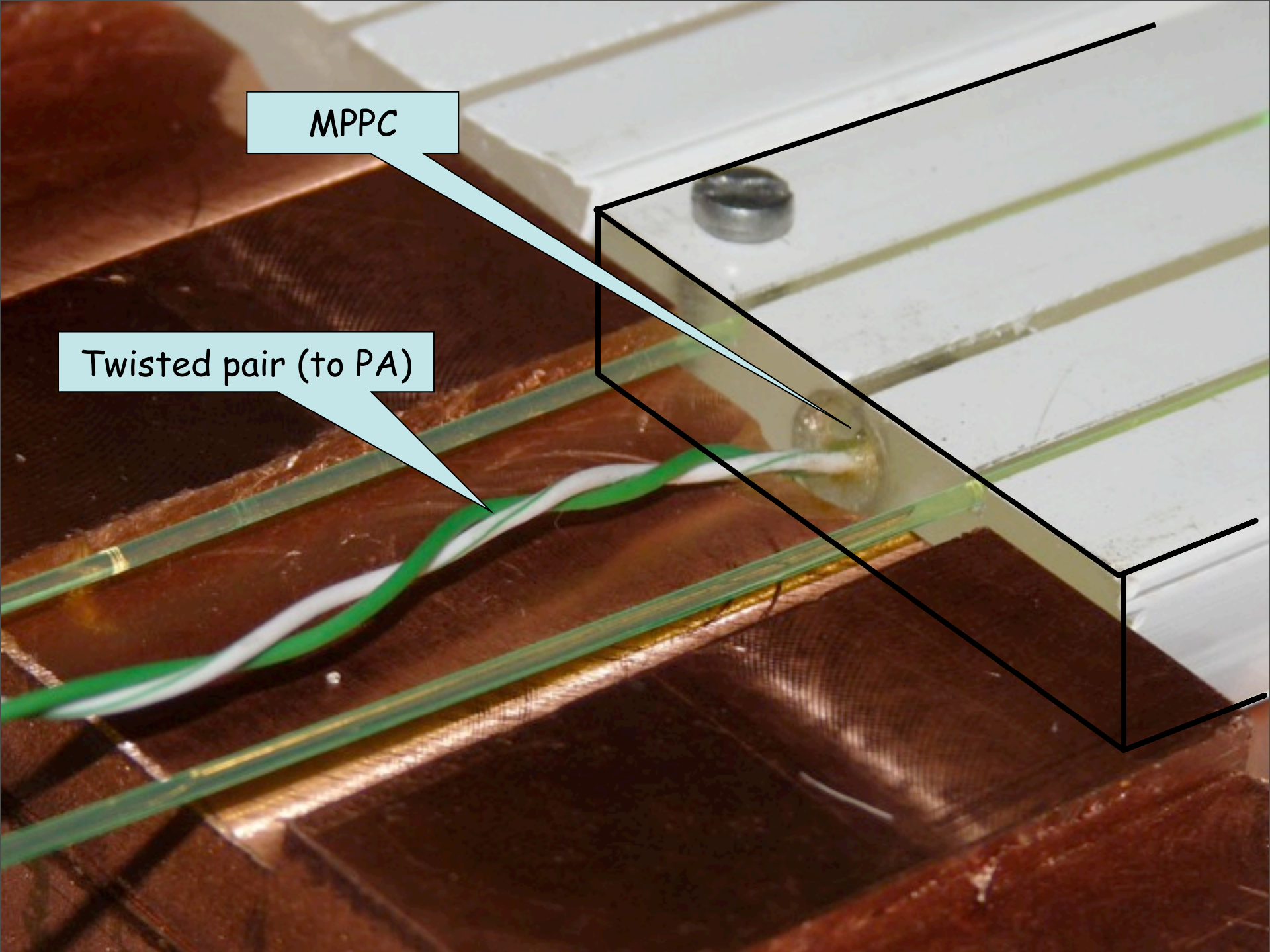
Wednesday, February 9, 2011



Wednesday, February 9, 2011

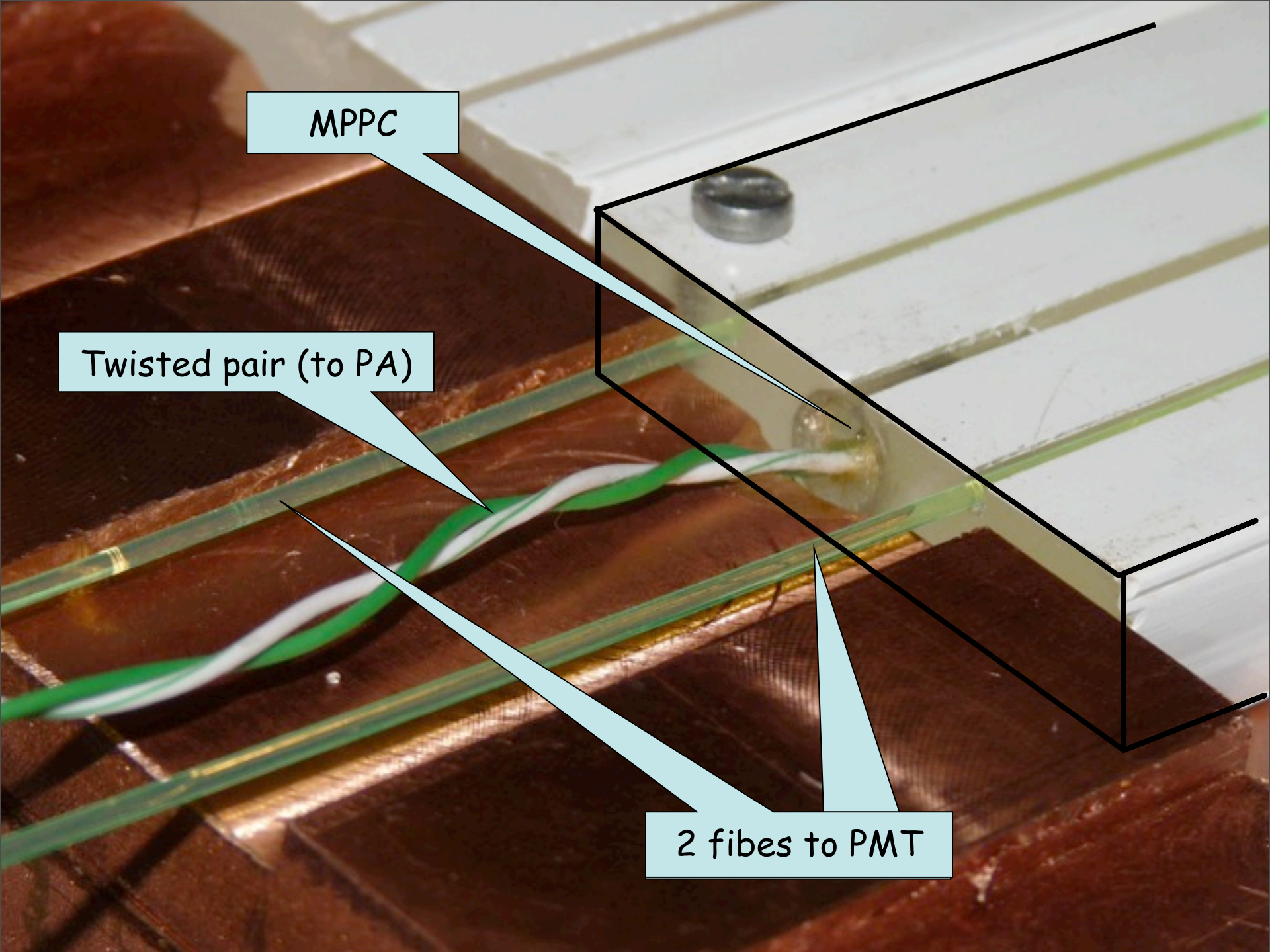


MPPC



MPPC

Twisted pair (to PA)



MPPC

Twisted pair (to PA)

2 fibers to PMT

Mounting table (with Cu bottom shielding)



Mounting of the first section



Wednesday, February 9, 2011



Wednesday, February 9, 2011



Wednesday, February 9, 2011



Wednesday, February 9, 2011



Wednesday, February 9, 2011

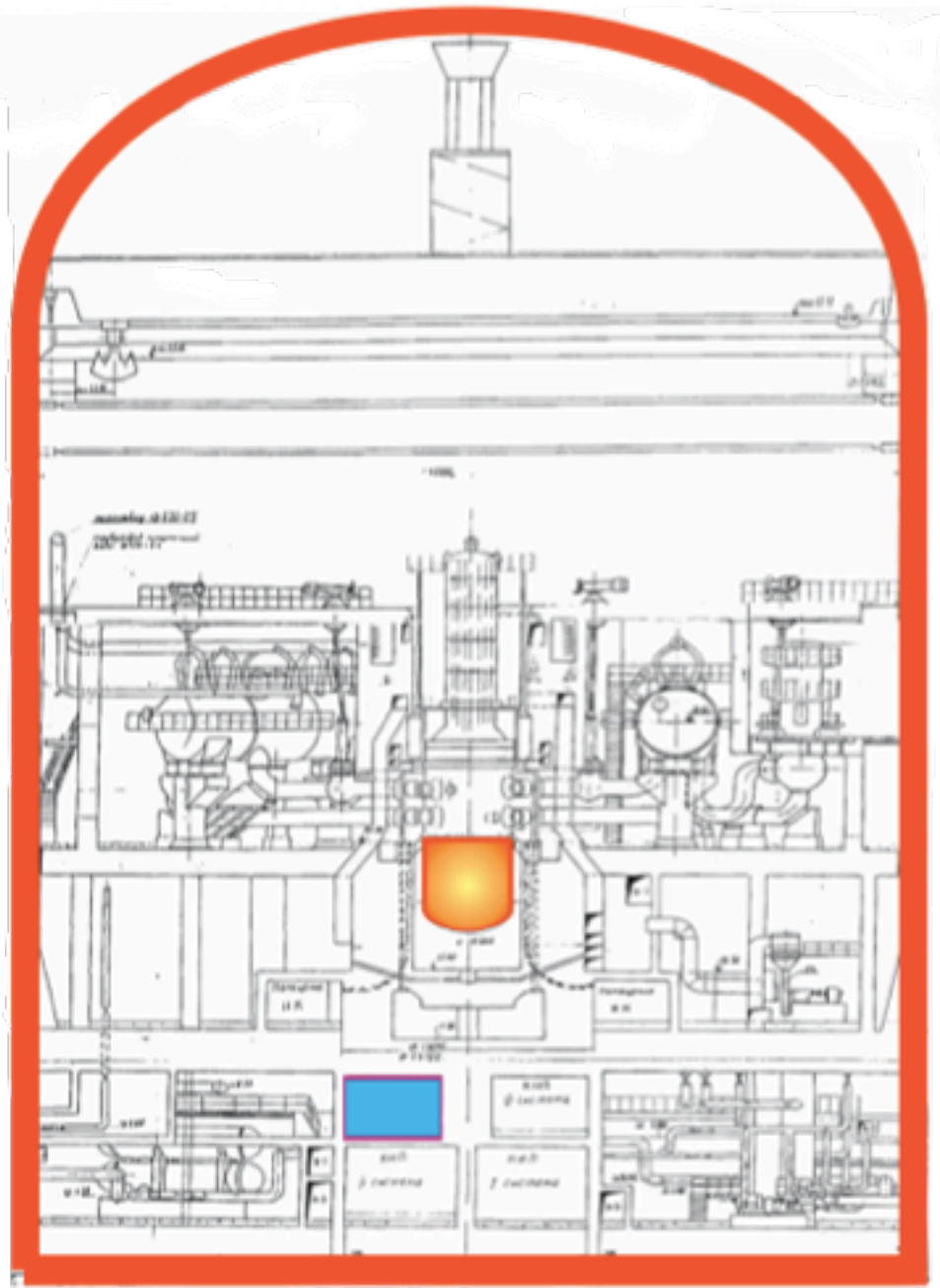


Wednesday, February 9, 2011

The section #0 was dismantled and sent to Moscow (our ITEP-members of the team are playing with it now).

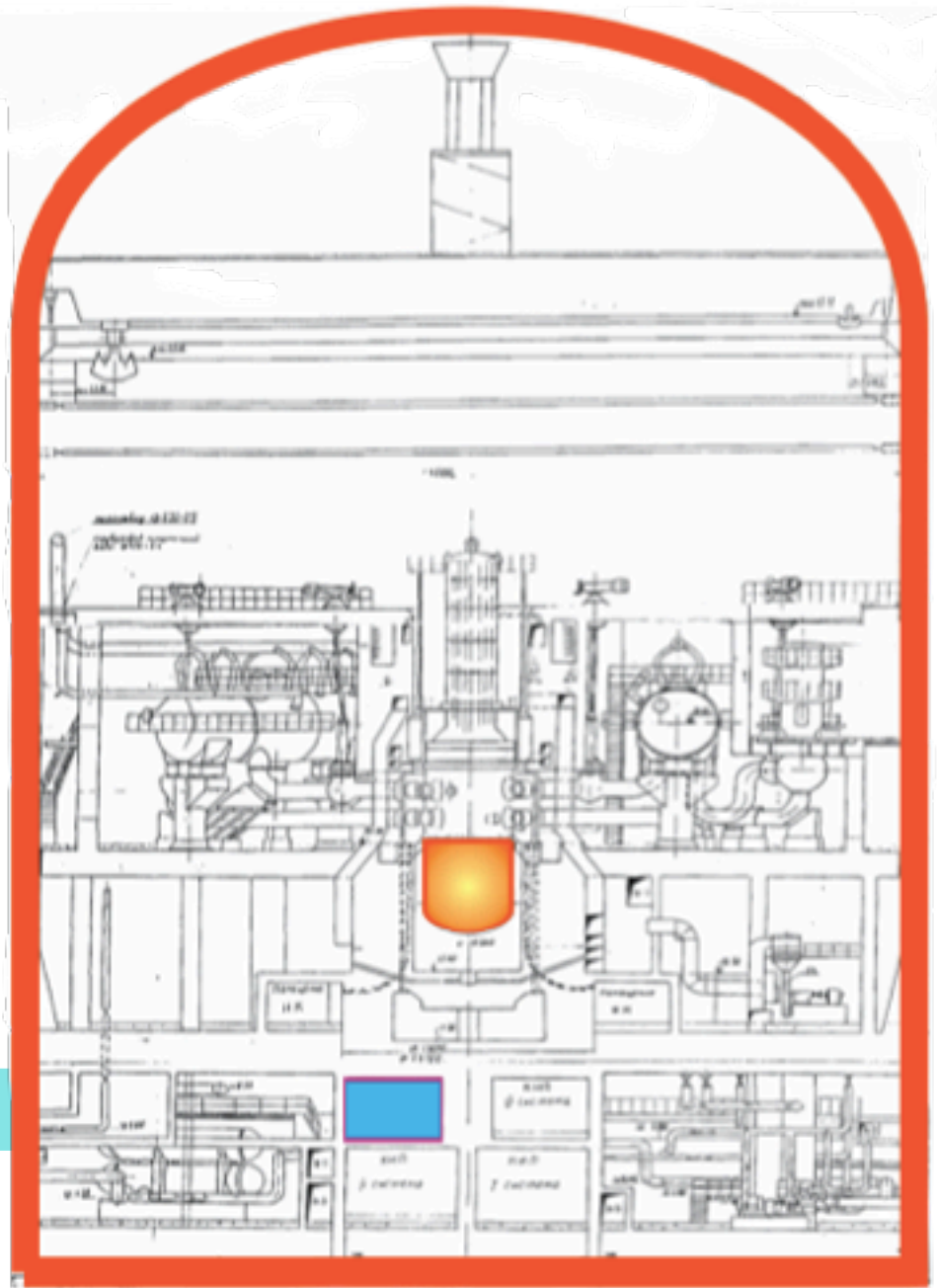
After some corrections we have started production of the sections #1, #2, etc.

**The GEMMA-1 site:
Reactor unit #2 of the
“Kalinin” Nuclear Power Plant
(400 km North from Moscow)**



The GEMMA-1 site:
Reactor unit #2 of the
“Kalinin” Nuclear Power Plant
(400 km North from Moscow)

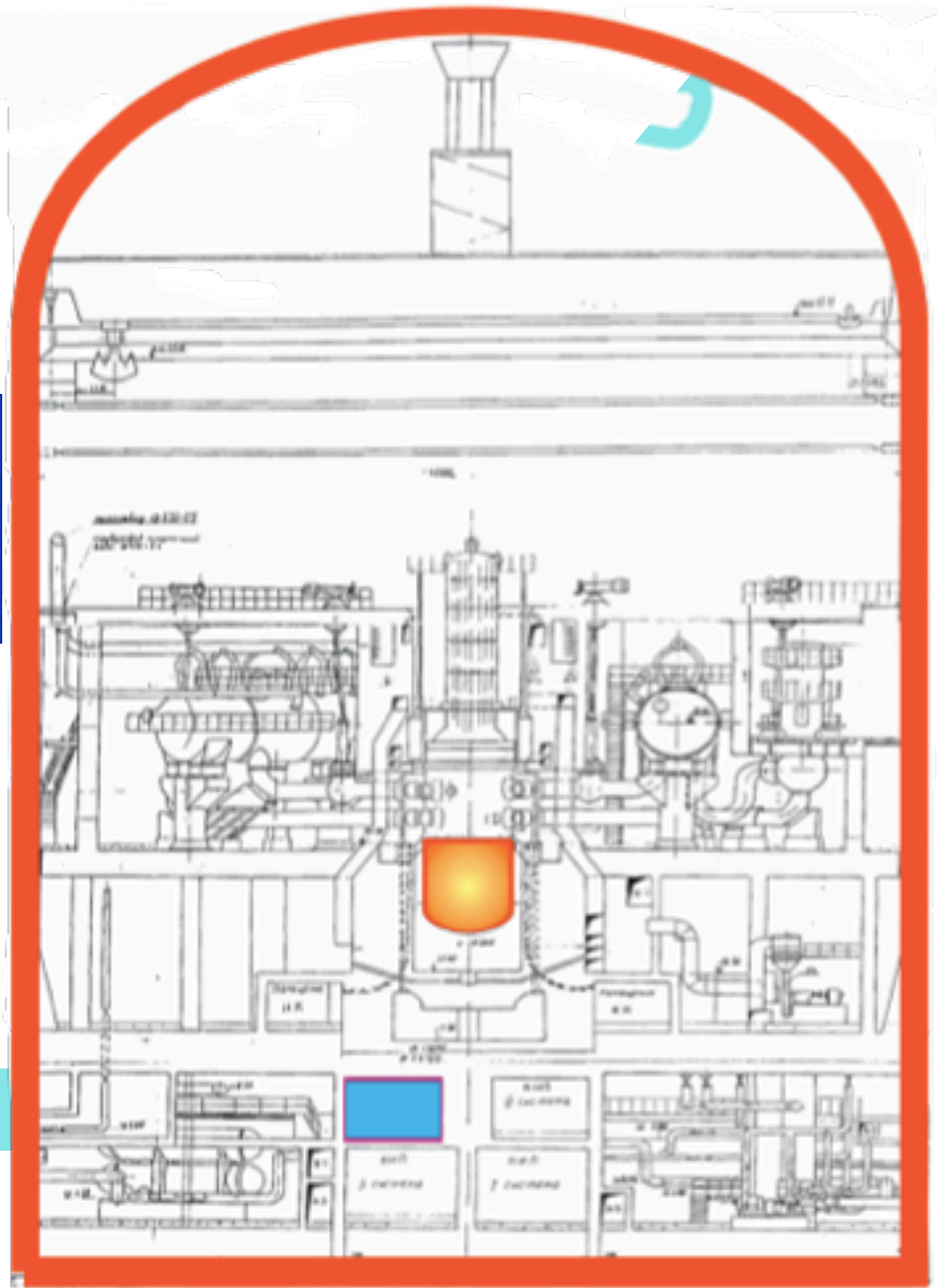
Technological room
just under reactor
14 m only!
 2.7×10^{13} n/cm²/s



The GEMMA-1 site:
Reactor unit #2 of the
“Kalinin” Nuclear Power Plant
(400 km North from Moscow)

Overburden
(reactor, building, shielding, etc.) :
~70 m of W.E.

Technological room
just under reactor
14 m only!
 2.7×10^{13} n/cm²/s



GEMMA background conditions

- **γ -rays** were measured with Ge detector. The main sources are: ^{137}Cs , ^{60}Co , ^{134}Cs .
- **Neutron** background was measured with ^3He counters, i.e., thermal neutrons were counted. Their flux at the facility site turned out to be **30 times lower** than in the outside laboratory room.
- **Charged** component of the cosmic radiation (**muons**) was measured to be **5 times lower** than outside.



КАЭС



Wednesday, February 9, 2011

КАЭС



S
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КАЭС

***To be started
on Sept. 2011***

**S
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b
l
e**



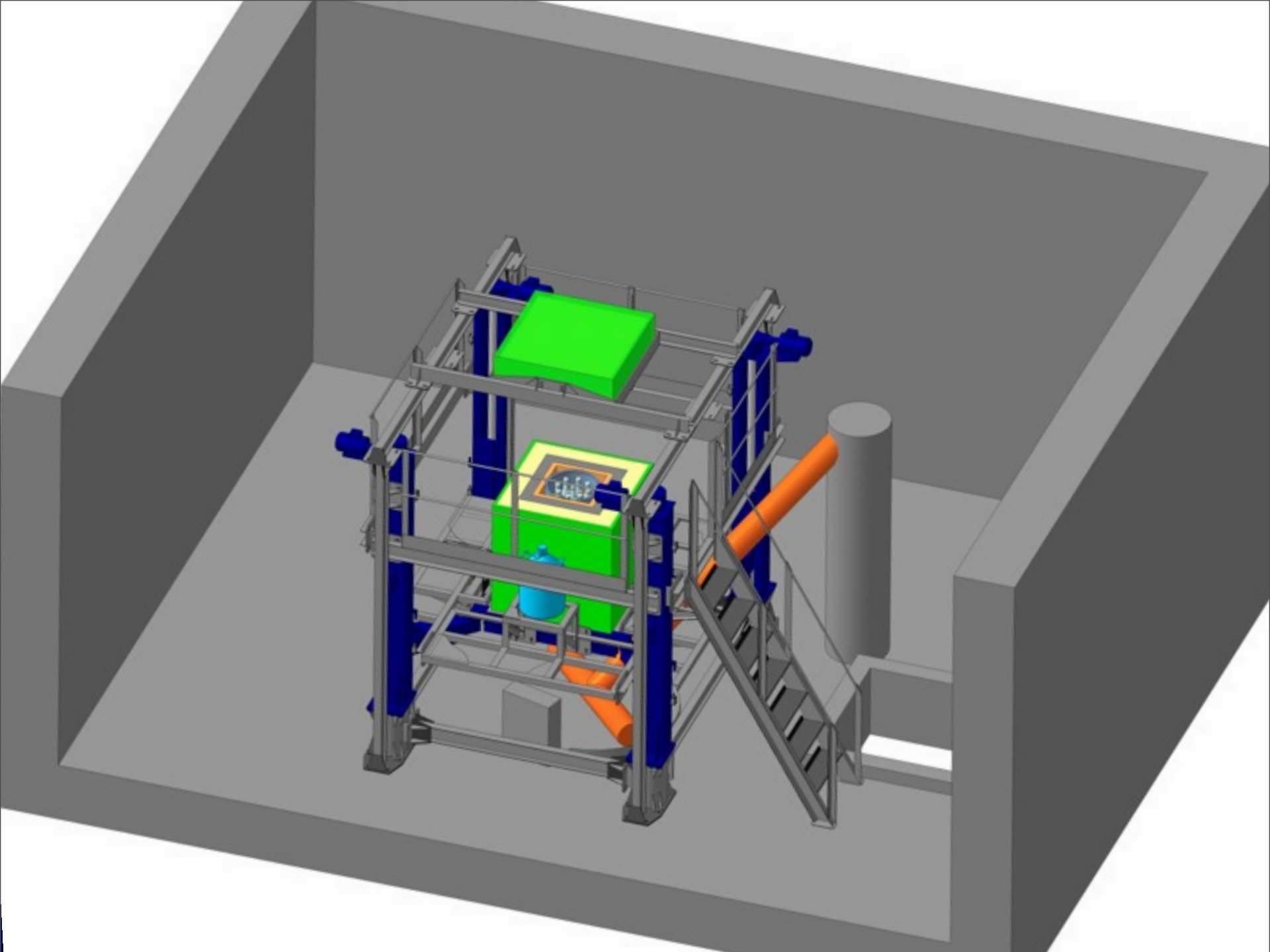
КАЭС

**To be started
on Sept. 2011**

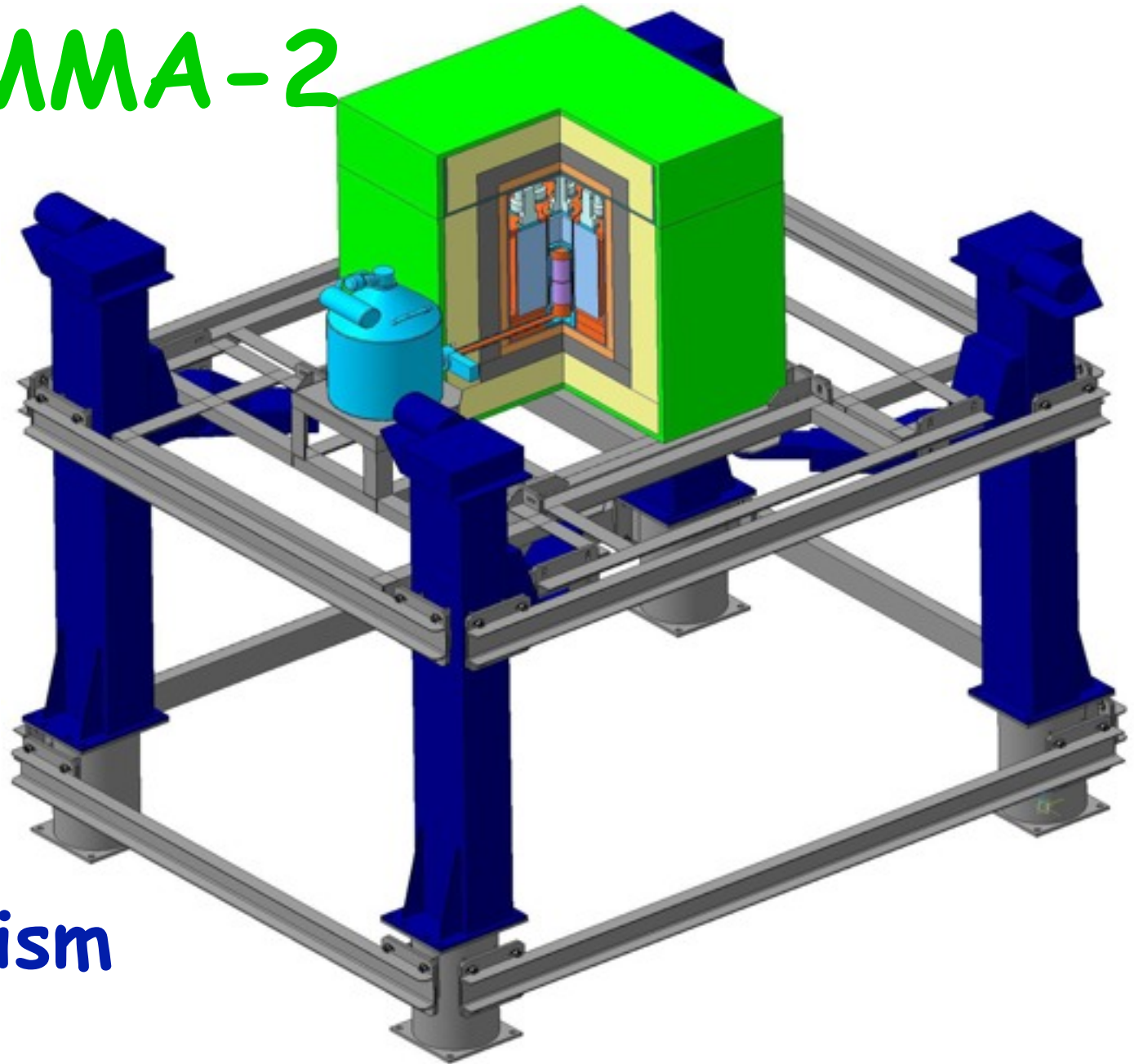
**14 → 10 m
 $5 \cdot 10^{13}$ v/cm²/s**

**S
i
t
e

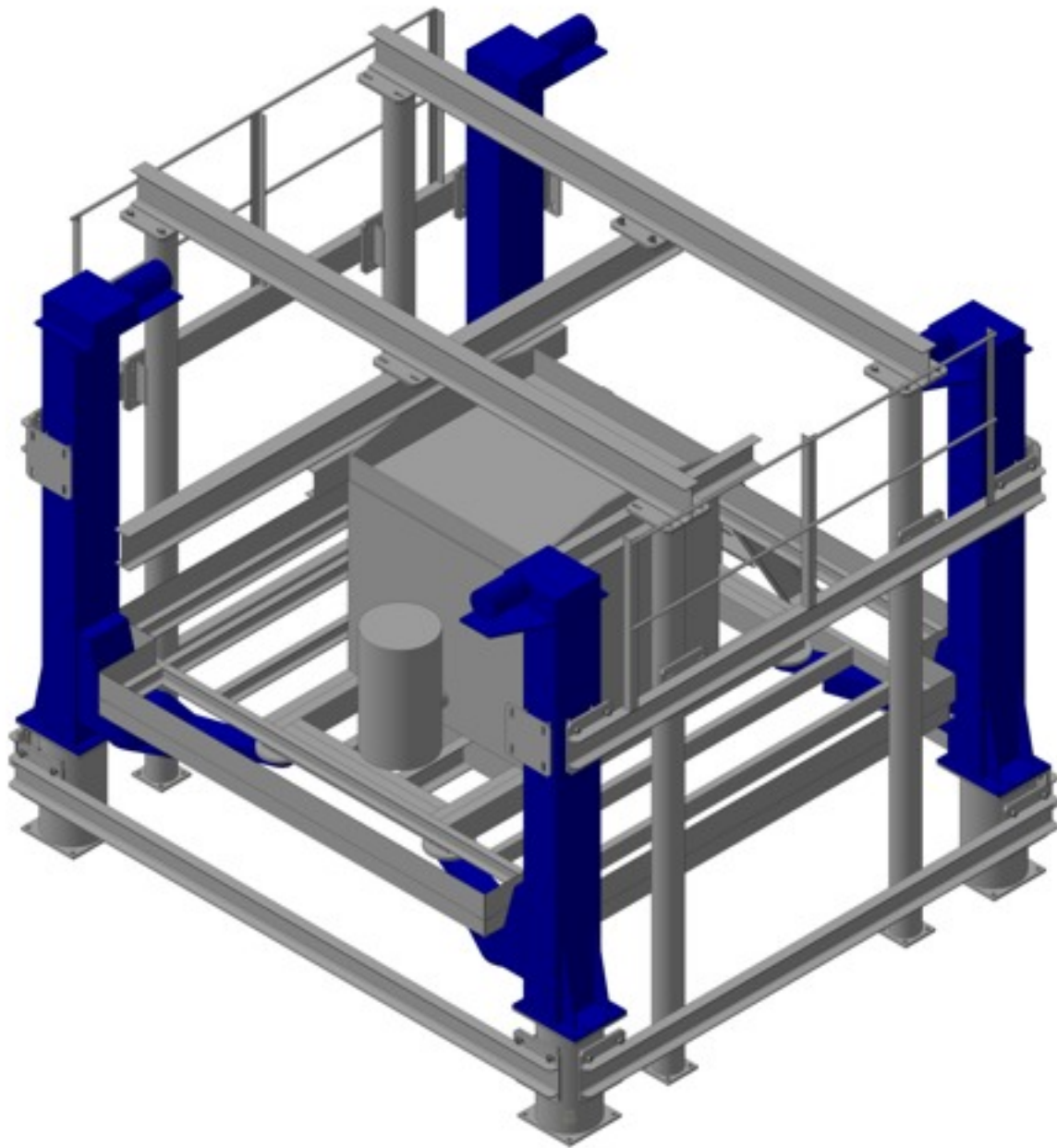
a
v
a
i
l
a
b
l
e**

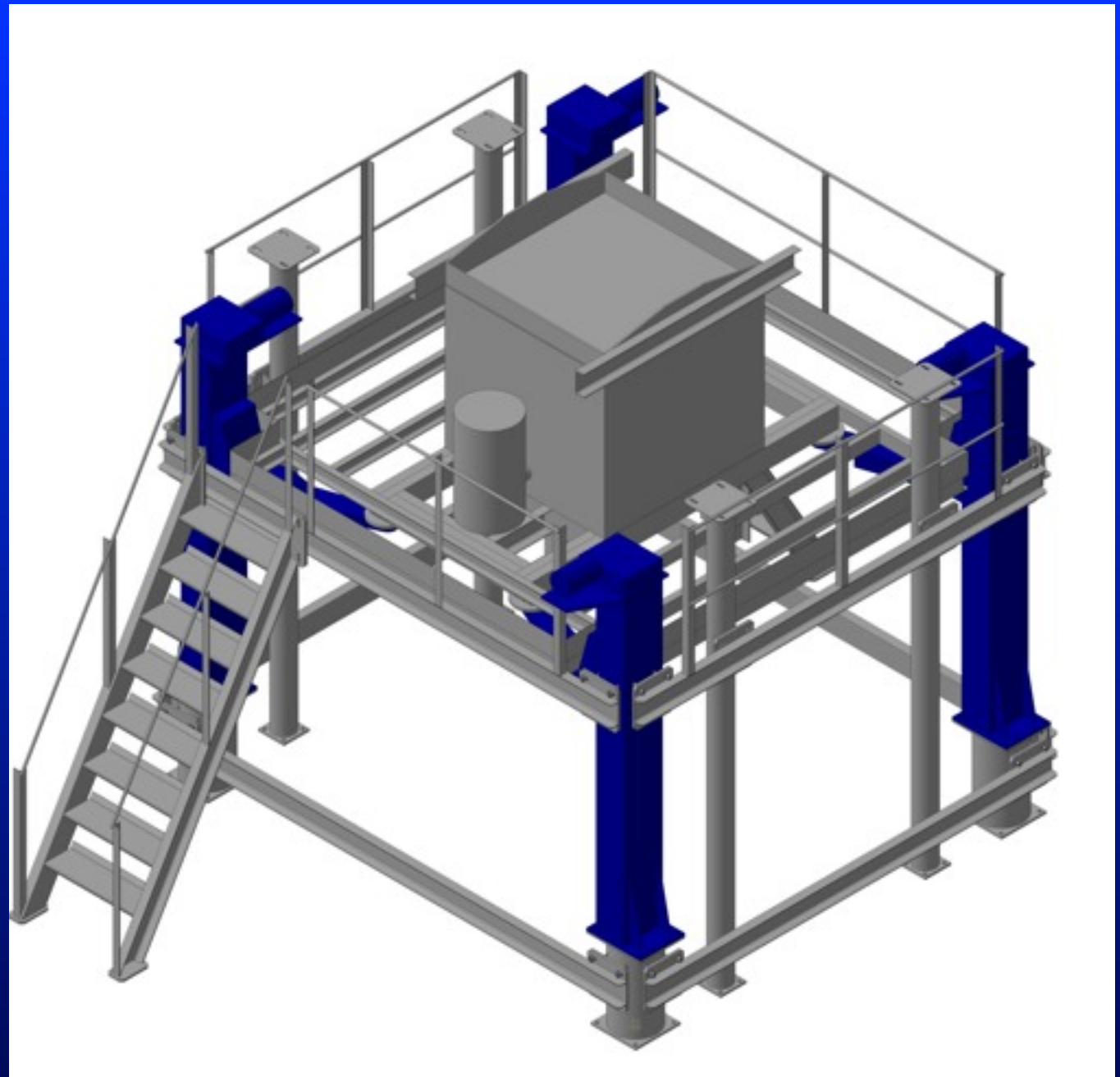


GEMMA-2



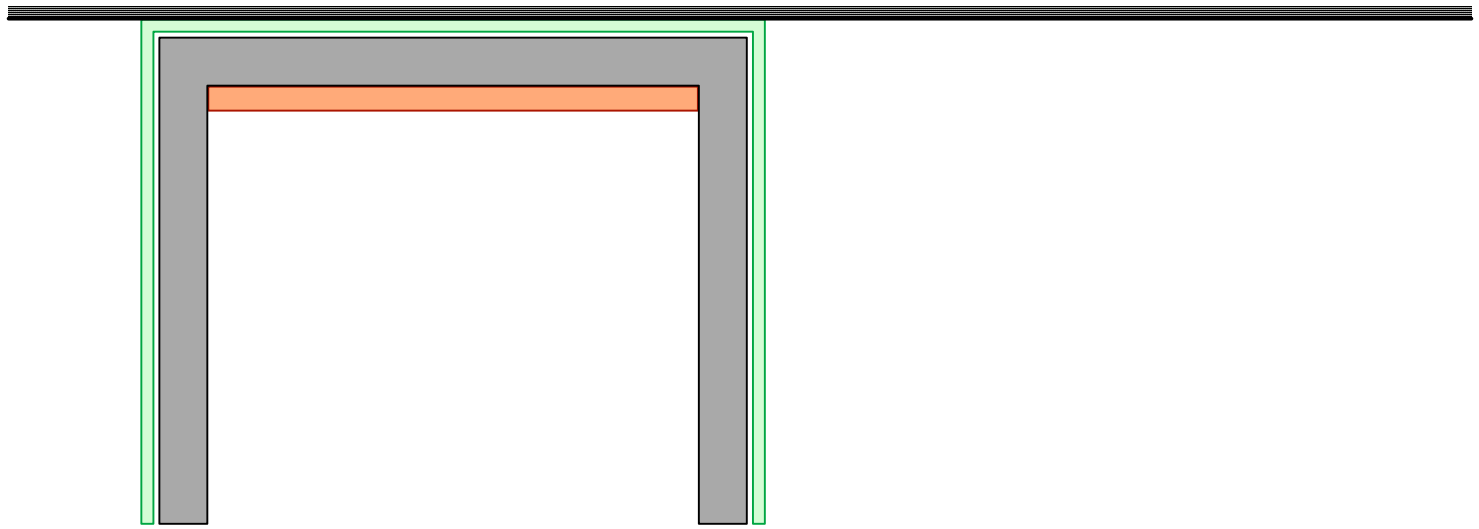
Lifting
mechanism

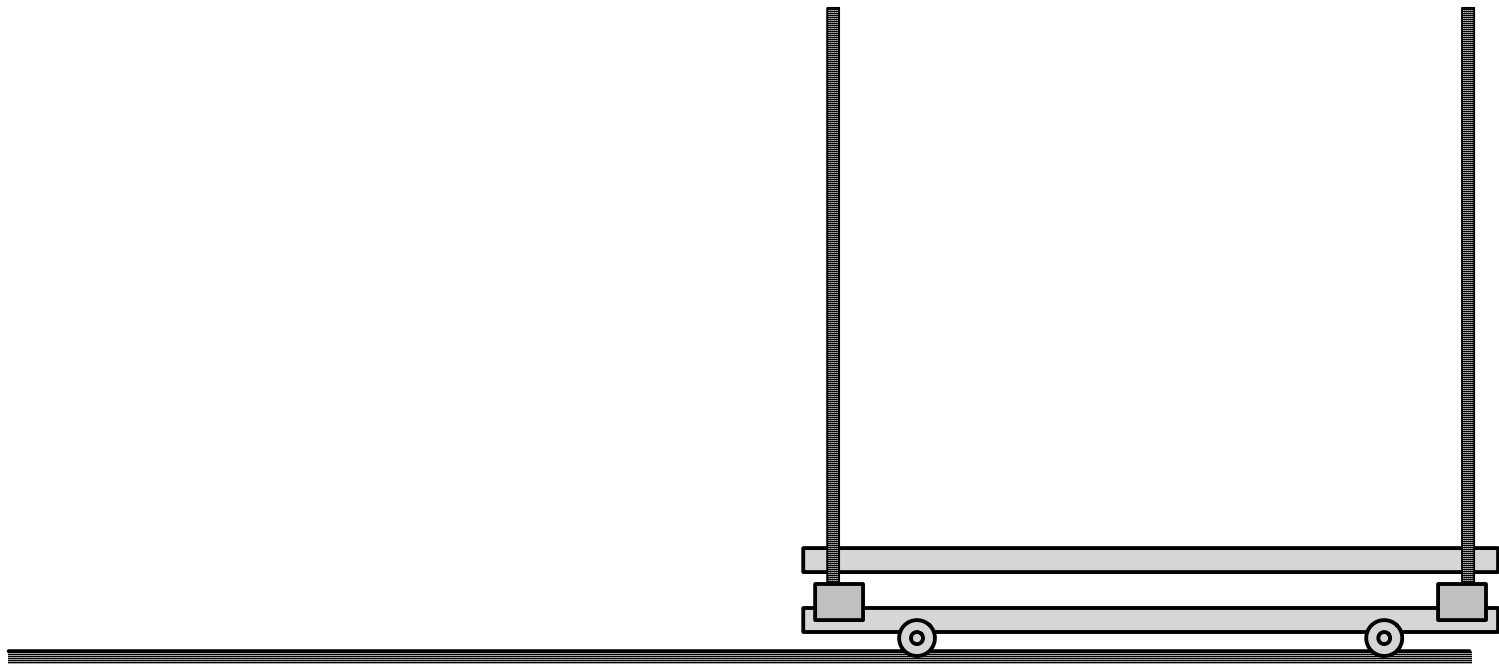
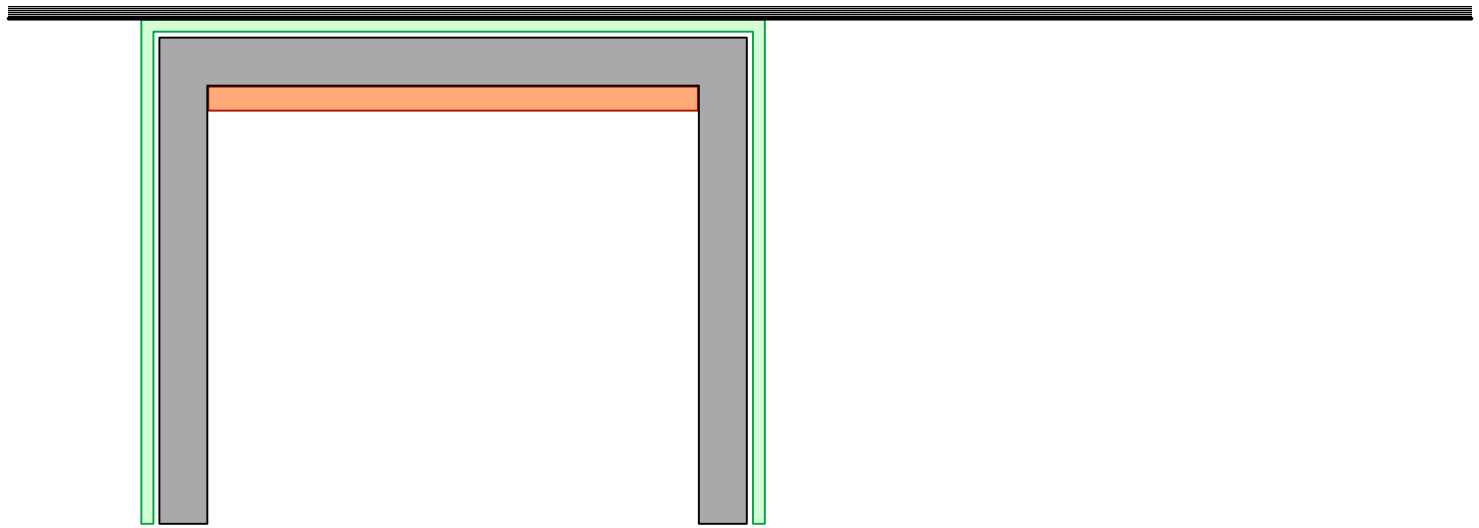


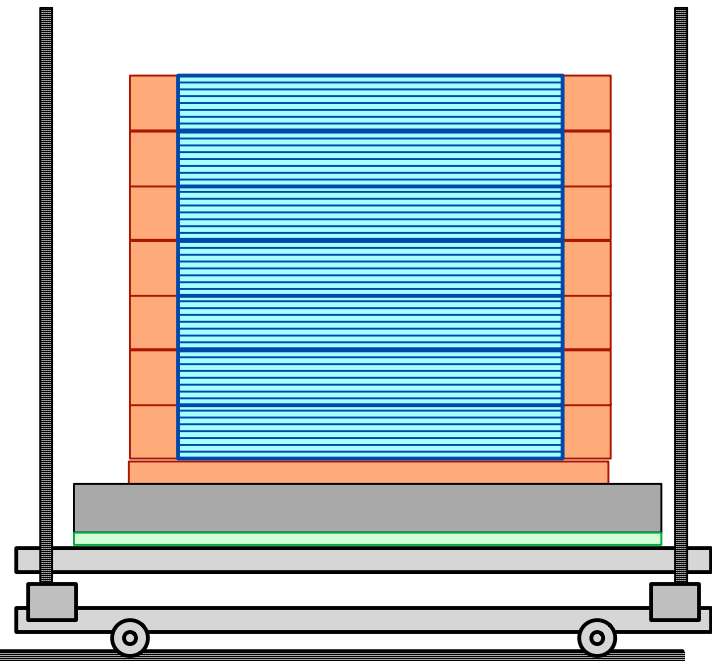
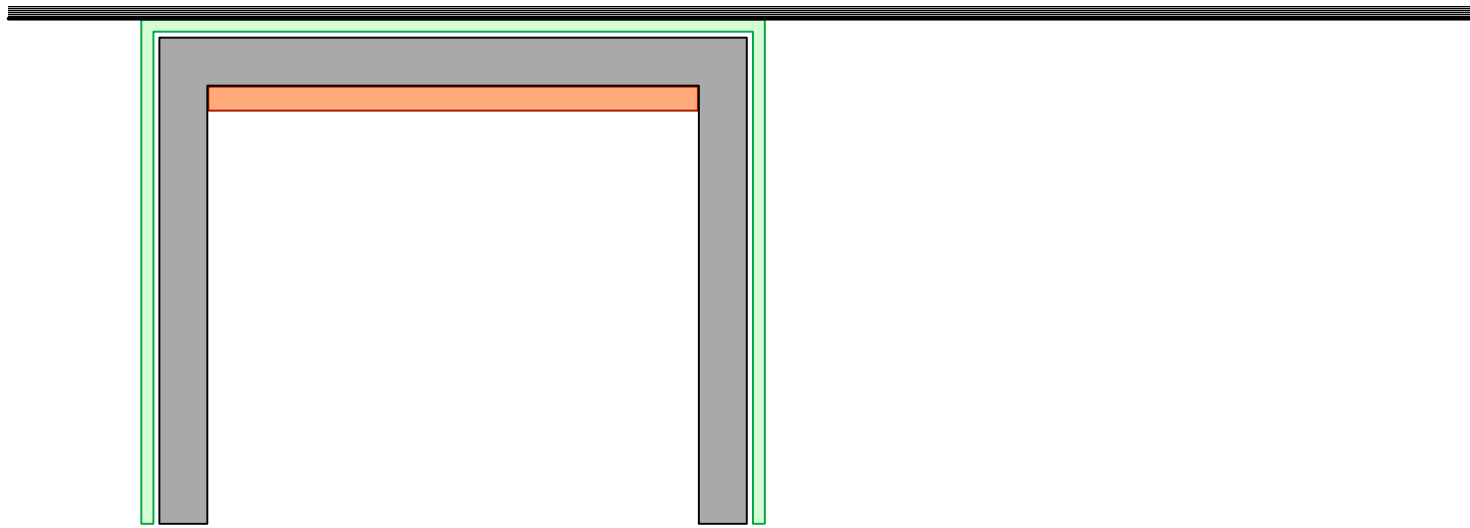


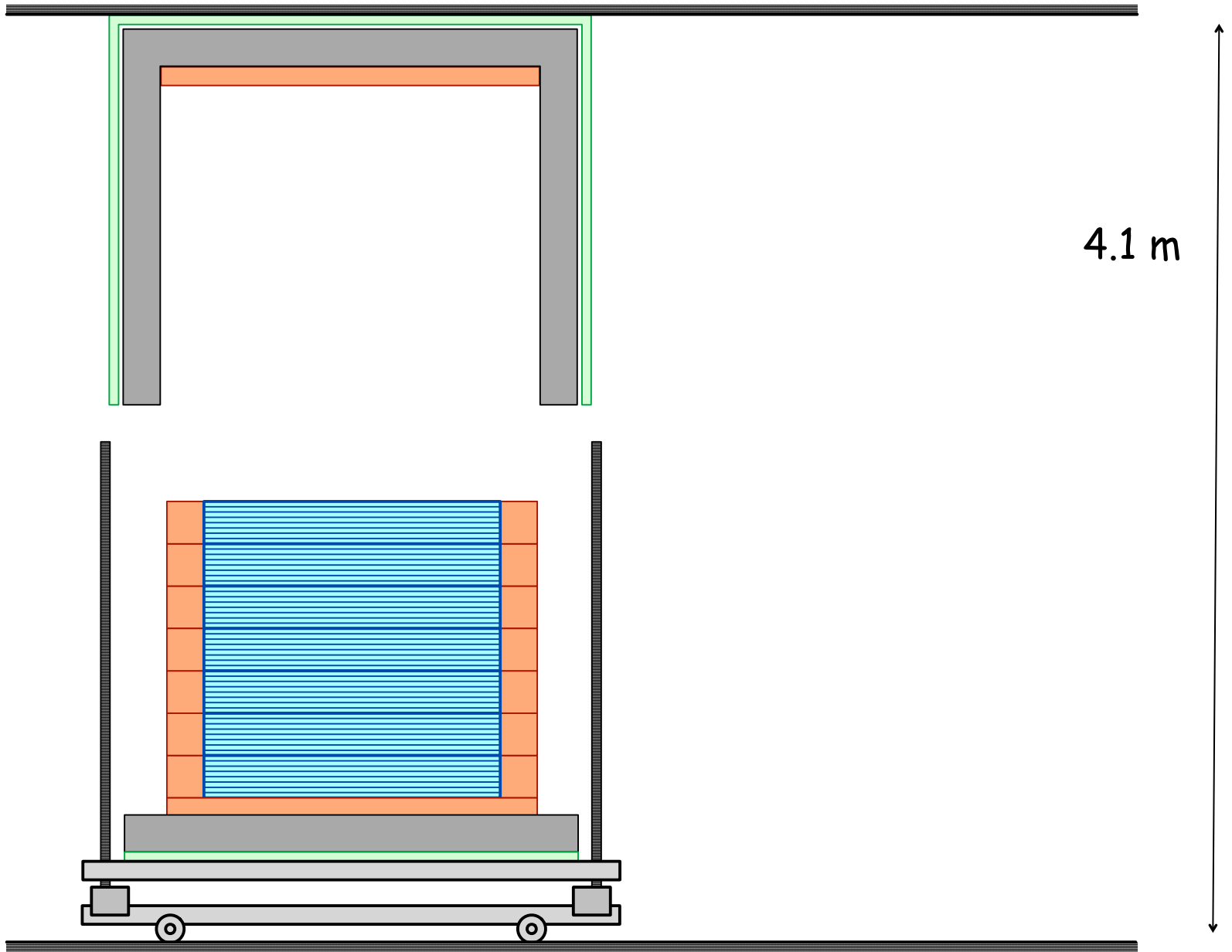


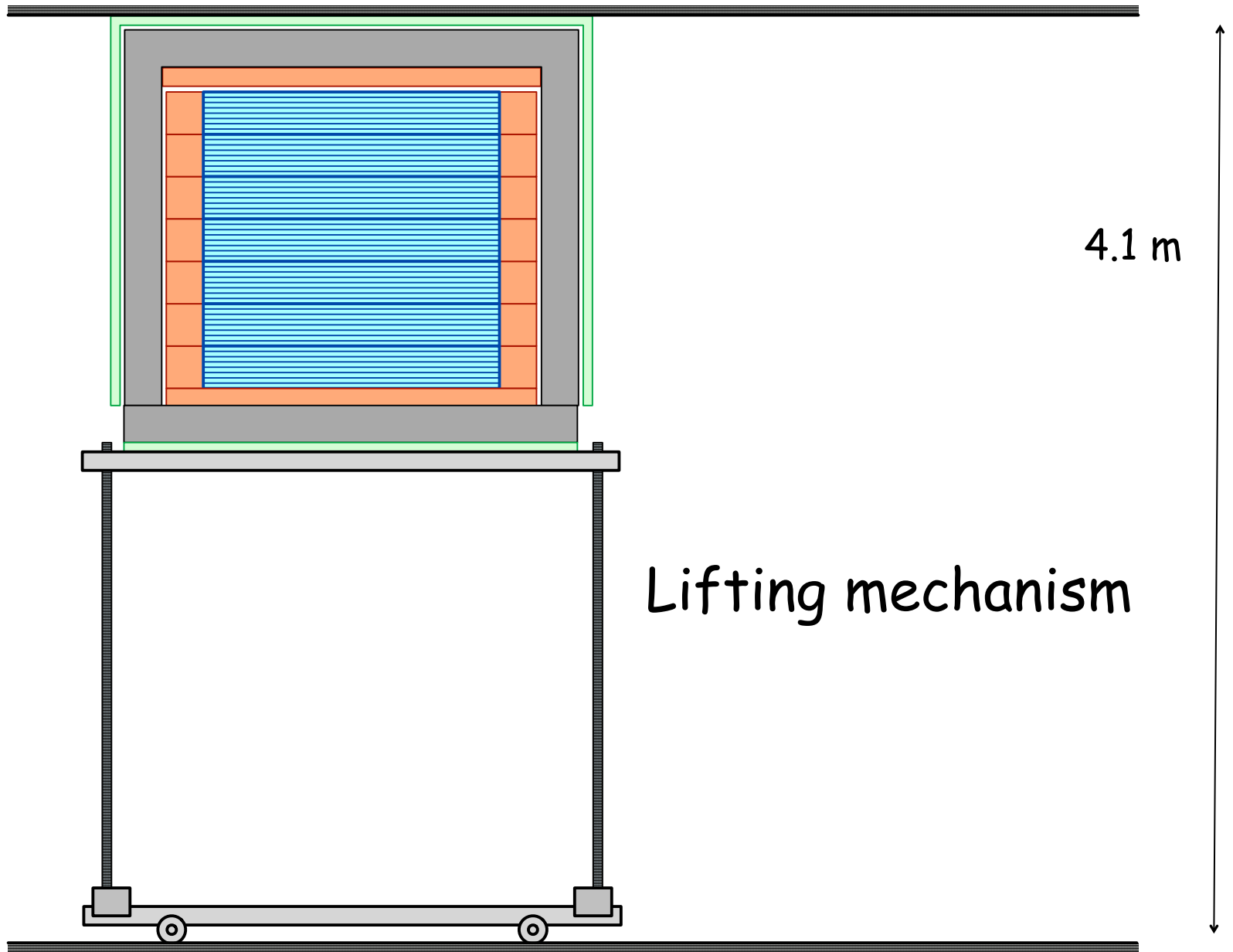
Could we make
the same
for the
DANSS











Important features

(resp. to conventional liquid scint.)

- [Handling is much safer](#) (*not caustic, spontaneously igniting, volatile or solvent*)
→ no restrictions to move the detector very close to the reactor core → higher neutrino flux → **better sensitivity**.
- [High segmentation](#) (2500 cells) → space information → better IBD-signature → stronger **BCKG suppression**.
- [PS](#) is not doped with Gd, but [interlayered](#) with it → better **quality** and **stability** of the scintillator.
- [WLS-fibers](#) improve homogeneity of light-collection → better **energy resolution**.
- Each cell in a module is looked through with both [individual MPPC](#) (*high QE, but bad noise and range*) and [common PMT](#) (*lower QE, but better range and stability*) → coincidence mode and combination of **PMT and MPPC advantages**.
- [Sectional structure](#) → possibility of “**partial**” **operation, renewal** and **upgrade**.
- Use of [lifting mechanism](#) → measurements at **different distance** → more **reliable data interpretation**.

Expected parameters:

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- Sensitive volume: 1 m³

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- Composition: 5 sections (1m × 1m × 0.2m)

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- Composition: 5 sections ($1\text{m} \times 1\text{m} \times 0.2\text{m}$)
or 10 sections ($1\text{m} \times 1\text{m} \times 0.1\text{m}$)

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of $(5\mathbf{X} + 5\mathbf{Y})$ modules = 2500 scint. strips
{ $1 \text{ module} = 5 \times 10 = 50 \text{ strips}$ }

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- Count rate: **9200 IBD-events per day**

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- Background: **40-50 events per day**

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- Due date: section №0 – **Feb 2010**

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- Background: **40-50 events per day**
- Due date: section №0 – **Feb 2010**
sections №1-3 – **2011** №4-5 – **2012**