

# KTA1

Slides to introductory lecture on  
14.10.2013

# Meilensteine der Kern- und Teilchenphysik

- 1897 Discovery of electron
- 1900  $\alpha$ ,  $\beta$  and  $\gamma$  radioactivity
- 1905 Photon identified as quantum of electromagnetic field
- 1911 Discovery of atomic nucleus
- 1912 Discovery of cosmic rays
  - Invention of cloud chamber
- 1913 Bohr model of atom
- 1919 Discovery of proton
- 1923 de Broglie wave-particle duality
- 1925 Introduction of electron spin
- 1926 Wave mechanics
- 1927 Uncertainty Principle
- 1928 Dirac wave equation
- 1930 Neutrino hypothesis
- 1931 Operation of first cyclotron and of Van der Graaff accelerator
- 1932 Discovery of positron
  - Discovery of neutron
- 1933 Discovery of electromagnetic showers
- 1934 Theory of beta decay
  - Discovery of Čerenkov effect
- 1935 Yukawa theory of nuclear forces
- 1936 Breit-Wigner resonance formula
- 1937 First evidence for mesotron (= muon)
- 1939 Observation of mesotron (= muon) decay
- 1940 Spin-statistics theorem
- 1945 Phase stability in accelerators (synchrotron principle)
- 1946 First proposal of Big Bang model
  - Two-meson hypothesis

1947	Discovery of pion and $\pi \rightarrow \mu$ decay in cosmic rays Prediction of muon-induced nuclear fusion Two-meson hypothesis (again) Discovery of $V$ particles	1965	Observation of cosmic microwave background radiation Introduction of colour quantum number and vector gluons
1948	Quantum electrodynamics Observation of $K \rightarrow 3\pi$ decay Pion production at accelerators	1967	Baryon asymmetry of universe (Sakharov criteria)
1950	Spark chamber invented Semiconductor detector invented Discovery of neutral pion and $\pi^0 \rightarrow 2\gamma$ decay	1968	Weinberg–Salam–Glashow electroweak model Deep inelastic $ep$ scattering. Bjorken scaling and partons
1951	Observation of $\Lambda$ hyperon and neutral kaon, $K_S^0$	1970	Invention of multiwire proportional chamber Proposal of fourth quark (charm)
1952	Evidence for $\Delta(1232)\pi p$ resonance Strong focussing principle for synchrotron Invention of bubble chamber	1972	Solar neutrino deficit ( $^{37}\text{Cl}$ experiment) Fermilab Tevatron operates CKM matrix for weak quark decays
1953	Evidence for $\Sigma$ and $\Xi$ hyperons First $V$ events at accelerator: associated production First hypernucleus event $\tau-\theta$ ( $= K\pi 3/K\pi 2$ ) paradox	1973	QCD as field theory of interquark interactions Neutrino scattering experiments confirm that partons are quarks Discovery of neutral weak currents
1954	Prediction of long-lived $K_L^0$ Invention of strangeness quantum number and classification	1974	Discovery of $J/\psi$ and $\psi' c\bar{c}$ resonances
1956	Observation of antiproton Detection of (anti)neutrinos from reactor Experimental evidence for $K_L^0$ Proposal for colliding-beam accelerators	1975	Charmed baryons and mesons Discovery of $\tau$ lepton $e^+e^- \rightarrow$ quark jets
1957	Observation of muon-induced nuclear fusion Two-component neutrino, $V - A$ theory Parity non-conservation in weak decays Resolution of $\tau-\theta$ paradox	1976	CERN SPS operates
1958	$(\pi \rightarrow e)/(\pi \rightarrow \mu)$ branching ratio Neutrino helicity measurement	1977	Discovery of $\Upsilon (= b\bar{b})$ states Emergence of Standard Model
1959	Operation of CERN PS, Brookhaven AGS	1978	Parity violation in polarised electron–deuterium scattering
1961	$K_L-K_S$ regeneration Discovery of $\rho$ , $\omega$ , $\eta$ pion resonances	1979	$e^+e^- \rightarrow$ three jets (PETRA)
1962	Pion $\beta$ -decay $\pi^+ \rightarrow \pi^0 e^+ \nu$ First accelerator neutrino beams and interactions $\nu_\mu$ and $\nu_e$ as separate neutrino flavours	1980	Evidence for $\Upsilon(3S)$ and $\Upsilon(4S)$ (CESR)
1963	Cabibbo theory of hadronic weak decays	1981	Observation of mesons and baryons containing $b$ quarks
1964	Streamer chamber invented Introduction of quarks and quark model First evidence for $\Omega^-$ hyperon Discovery of $CP$ violation in $K^0$ decay Higgs mechanism of spontaneous symmetry breaking	1983	Discovery of $Z^0$ and $W^\pm$ bosons
		1987	Observation of $B^0-\bar{B}^0$ mixing SN 1987A Supernova neutrino burst
		1990	$Z^0$ produced at $e^+e^-$ colliders LEP and SLC Number of neutrino flavours $N_\nu = 3$ from $Z^0$ width
		1993	Solar neutrino deficit confirmed in gallium experiments Atmospheric neutrino flavour anomaly Precise measurements of $Z^0$ decay parameters confirm Standard Model
		1995	Discovery of $t$ quark at Fermilab collider
		1997	$e^+e^- \rightarrow W^+W^-$ pair production at LEP 200 collider

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1998: Super-Kamiokande: Atmospheric neutrino oscillations

2000: tau-neutrino detection

2002: SNO: neutrino flavor conversion of solar neutrinos

2002: CP violation in B-mesons

2007: direct detection of solar Be-7 neutrinos

2011: non-zero mixing angle theta\_13

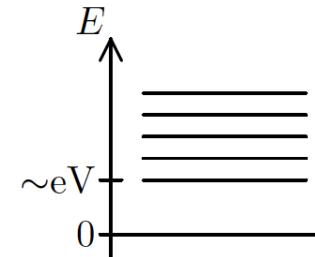
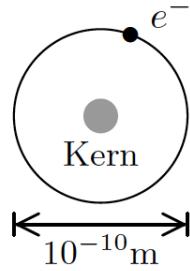
2012: Discovery of Higgs particle



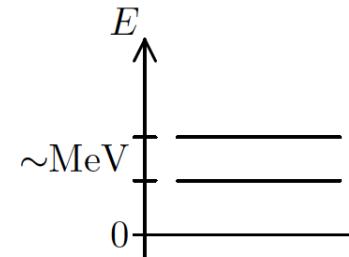
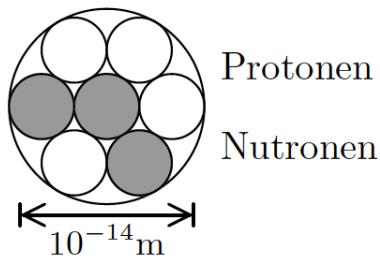
The Nobel Prize in Physics 2013  
 François Englert, Peter Higgs

# Längenskala und Hierarchie der Struktur der Materie

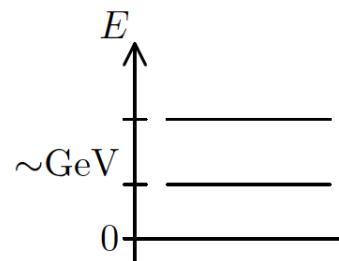
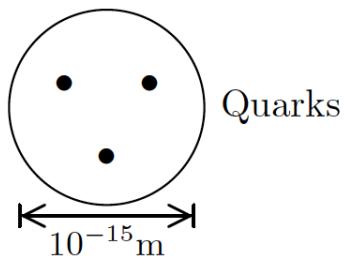
- Atom:



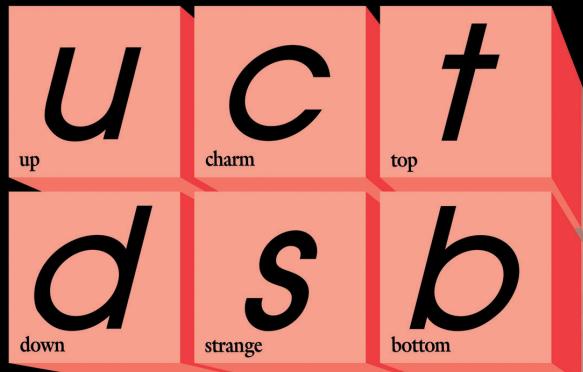
- Kern:



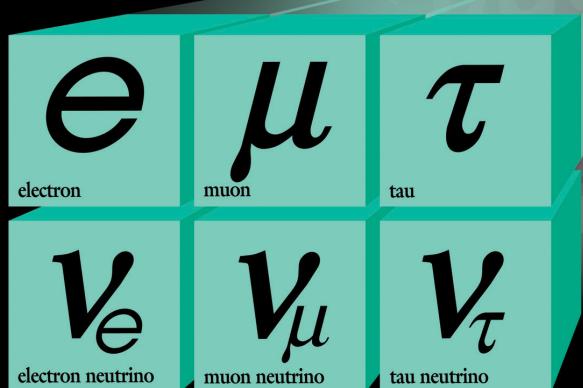
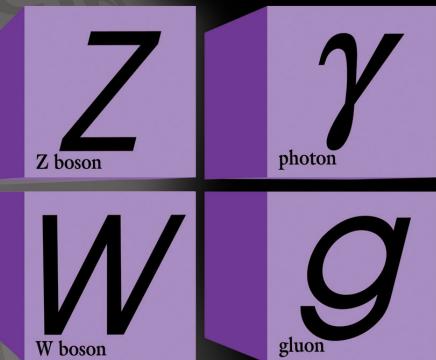
- Nukleon:



# Quarks



# Forces

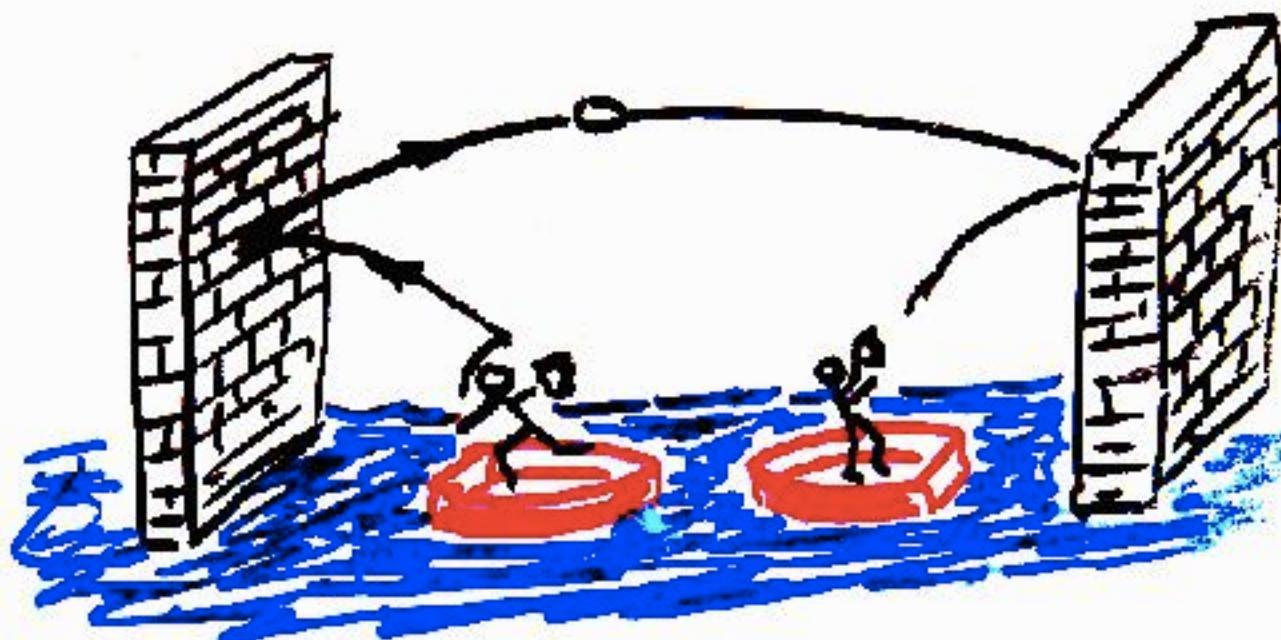


# Leptons

# Wechselwirkung vermittelt durch Austauschteilchen

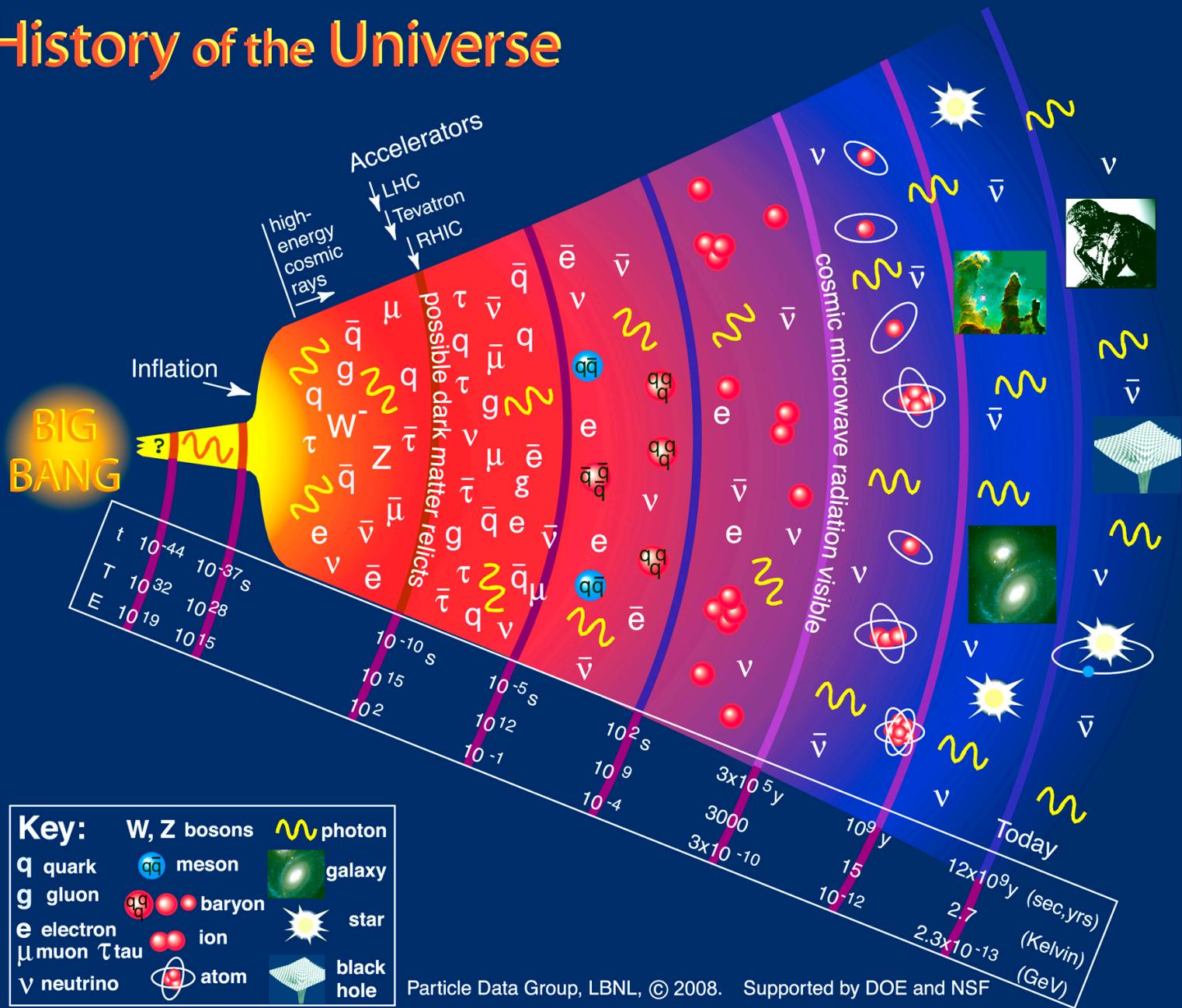


Abstoßung



Anziehung

# History of the Universe



Particle Data Group, LBNL, © 2008. Supported by DOE and NSF

Urknalltheorie der Entstehung des Universums.

Stützen: kosmische Hintergrundstrahlung, primordiale Nukleosynthese (Bild der leichten Kerne ( $A \leq 7$ )).

**Zeitskalen:**

$t \simeq 10^{-44}$  s: Planck-Ära,  $10^{19}$  GeV;  $U_{pot} \approx -mc^2$

$t \simeq 10^{-36}$  s: GUT-Brechung,  $10^{15}$  GeV

$t \simeq 10^{-10}$  s: Brechung elektromagnetischer und schwacher Wechselwirkung;  $10^2$  GeV ( $\sim M_{W,Z_0}c^2$ )

$t \simeq 10^{-6}$  s: Quark-Confinement;  $1$  GeV  $\rightarrow$  Bildung der Nukleonen  
 $p + \bar{p} \longrightarrow 2\gamma$   
 $n + \bar{n} \longrightarrow 2\gamma$

$t \simeq 1$  s:  $\nu$ -Entkopplung;  $1$  MeV  
 $e^- + e^+ \longrightarrow 2\gamma$

$t \simeq 5$  min: Bildung leichter ( $A \leq 7$ ) Kerne;  $10^2$  keV  
z.B.  $p + n \longrightarrow {}^2\text{H} + \gamma$  ( $E_\gamma = 2,2$  MeV)

$t \simeq 10^5$  a: Photonen entkoppeln;  $\simeq$  eV  
 $p + e^- \longrightarrow \text{H-Atom}$

$t \gtrsim 10^6$  a: Strukturbildung